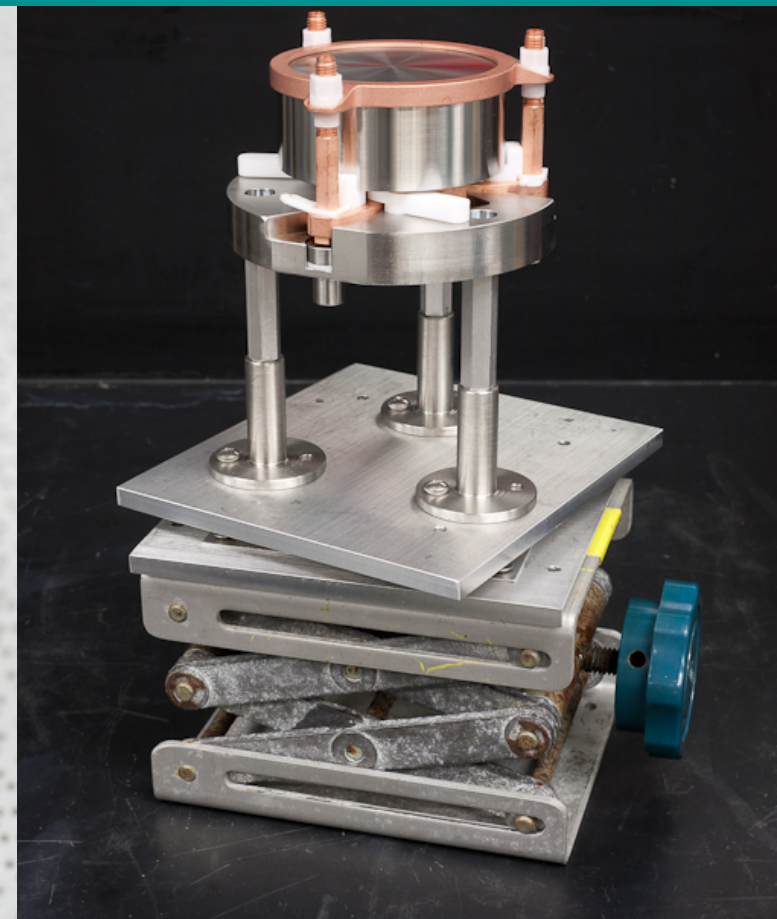
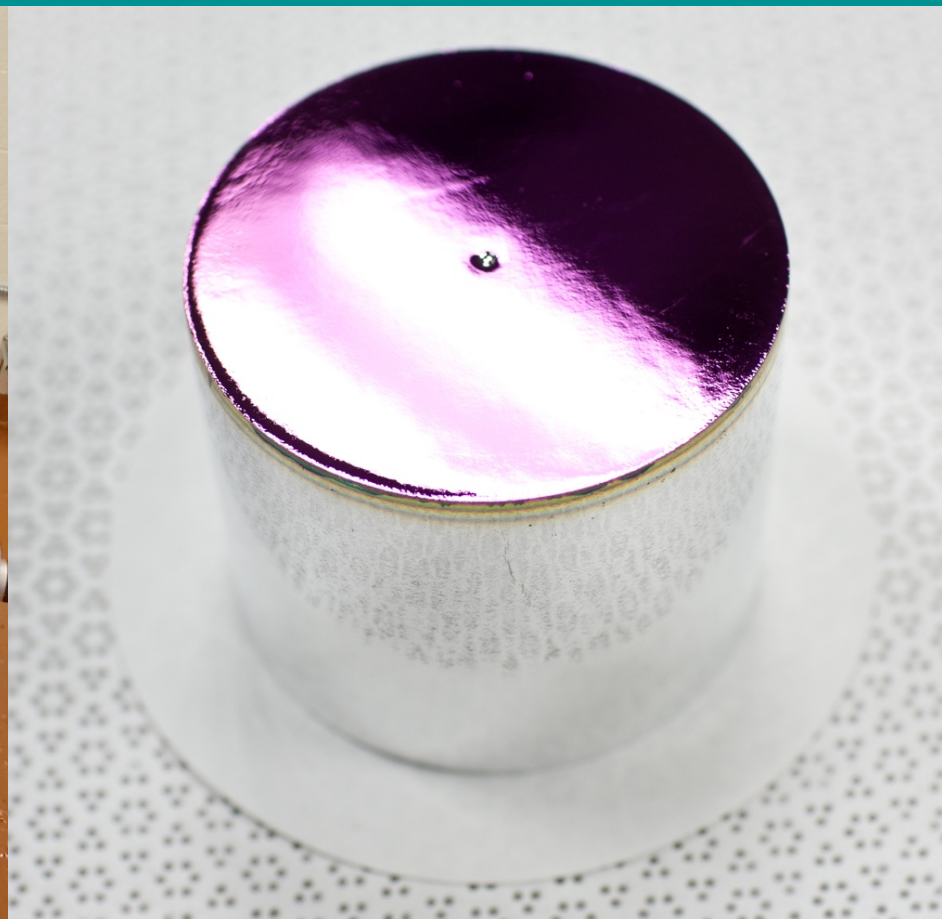


# Searching for neutrinoless double-beta decay of germanium-76 in the presence of backgrounds



**Alexis Schubert**  
**MAJORANA Collaboration**

# The MAJORANA Collaboration



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Joint Institute for Nuclear Research, Dubna, Russia

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Rodriguez, Michael Ronquest, Harry Salazar, Wenqin Xu

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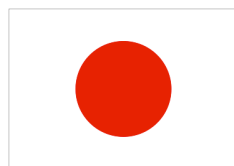
**Dustin Combs**, **Lance Leviner**, David G. Phillips II, Albert Young

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Jim Beene, Fred Bertrand, Greg Capps, Alfredo Galindo-Uribarri,  
Kim Jeskie, David Radford, Robert Varner, Brandon White,  
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Mary Kidd

University of Alberta, Edmonton, Alberta

Aksel Hallin

University of North Carolina, Chapel Hill, North Carolina and TUNL

**Padraic Finnerty**, Florian Fraenkle, **Graham K. Giovanetti**, Matthew P. Green,  
Reyco Henning, Mark Howe, **Sean MacMullin**, **Kyle Snavelly**,  
**Jacqueline Strain**, **Kris Vorren**, John F. Wilkerson

University of South Carolina, Columbia, South Carolina

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Vince Guiseppe, Kirill Pushkin, **Nathan Snyder**

University of Tennessee, Knoxville, Tennessee

Yuri Efremenko, Sergey Vasiliev

University of Washington, Seattle, Washington

Tom Burritt, Jason Detwiler, Peter J. Doe, Greg Harper, **Jonathan Leon**,  
David Peterson, R. G. Hamish Robertson, **Alexis Schubert**, Tim Van Wechel





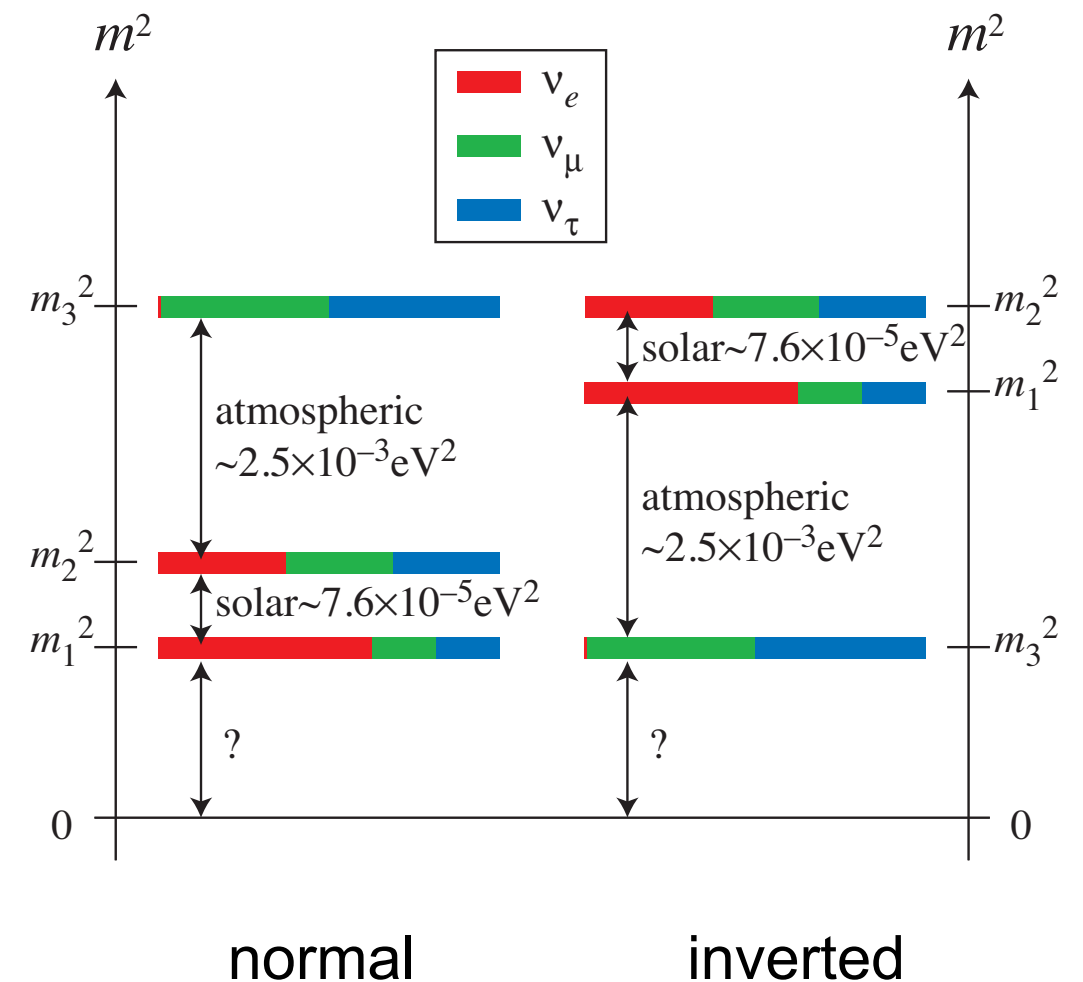
# Outline

- Neutrinoless double-beta decay
- The MAJORANA DEMONSTRATOR
- Studies with an R&D detector



# Neutrino questions

- What is the absolute mass scale of neutrinos?
- What is the neutrino mass hierarchy?
- Is the neutrino its own antiparticle (a Majorana particle)?
- Is lepton number a conserved quantity?

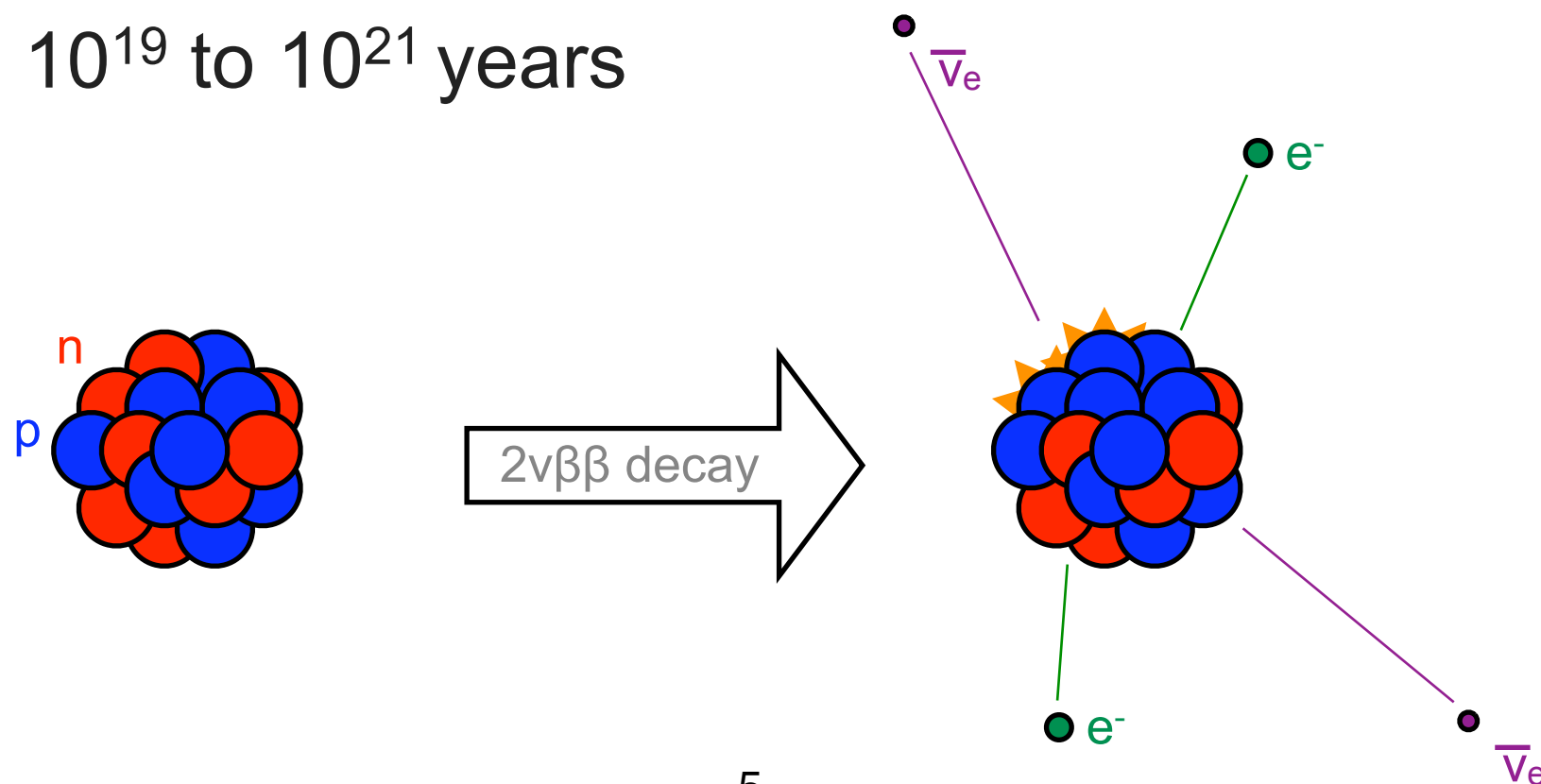


[hitoshi.berkeley.edu/neutrino](http://hitoshi.berkeley.edu/neutrino)



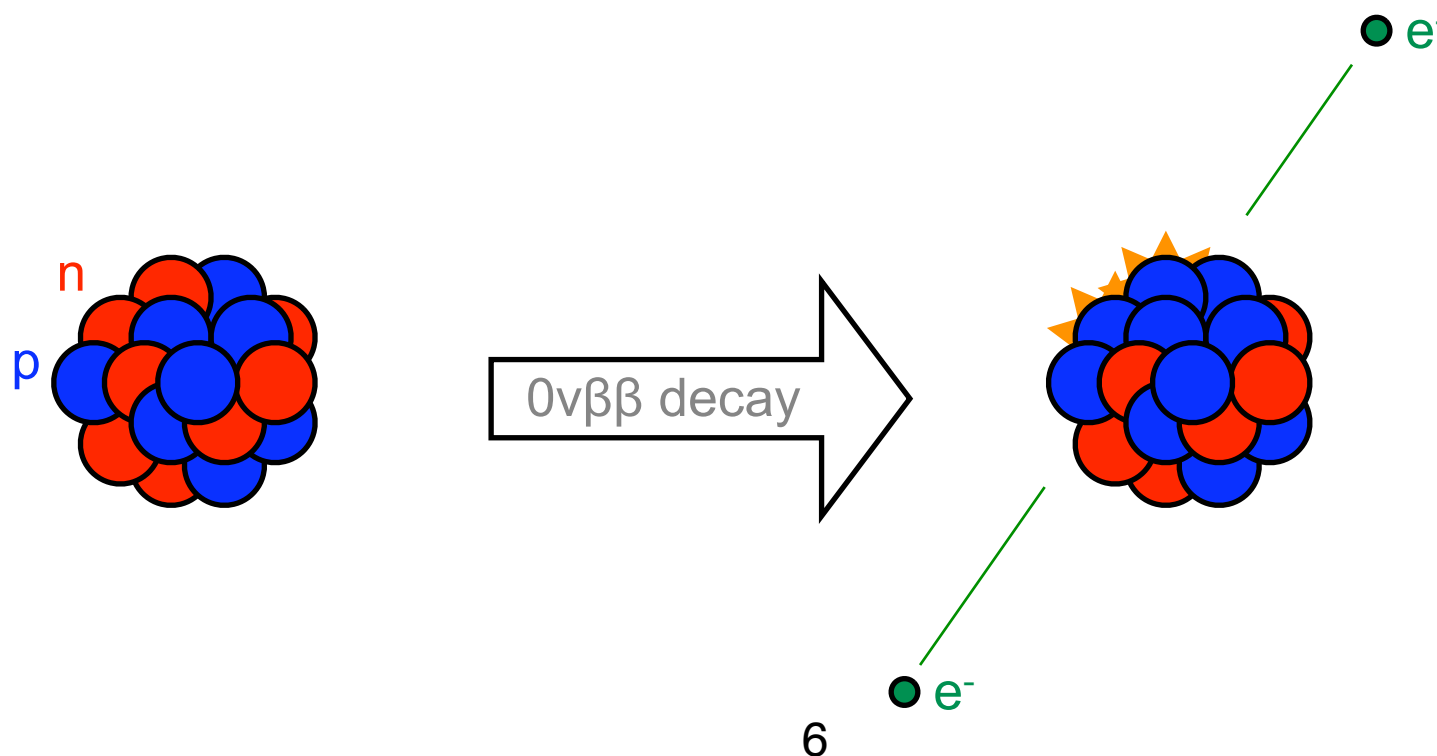
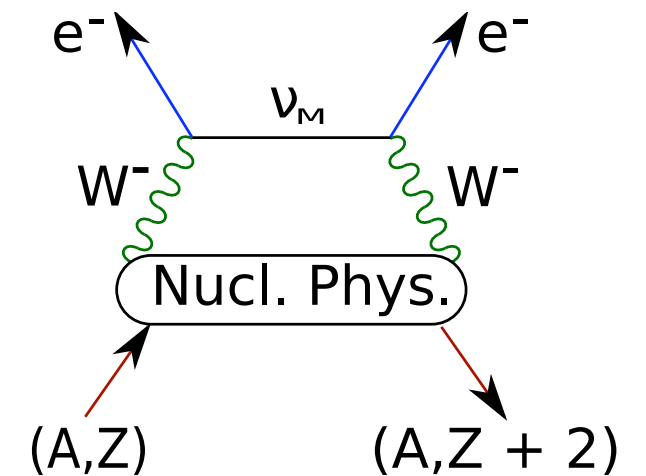
# Double-beta decay

- Process that occurs for some nuclei with even number of protons and neutrons
- Occurs with the emission of two neutrinos ( $2\nu\beta\beta$ )
- Observed in many nuclei
- $T_{1/2} \sim 10^{19}$  to  $10^{21}$  years



# Neutrinoless double-beta decay ( $0\nu\beta\beta$ )

- Observation would indicate:
  - Neutrino is a Majorana particle
  - Lepton number is violated
- Information about mass may be available

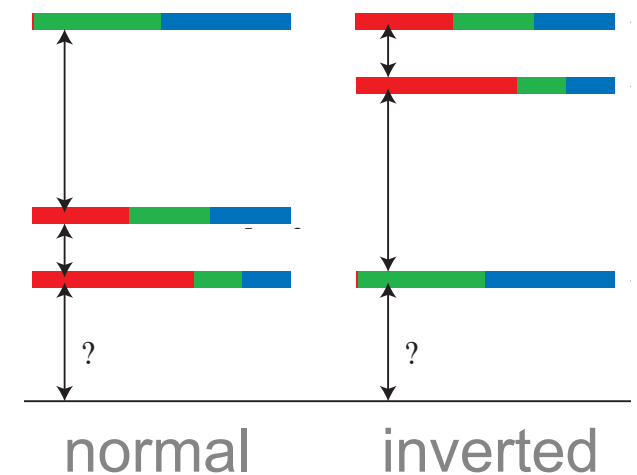
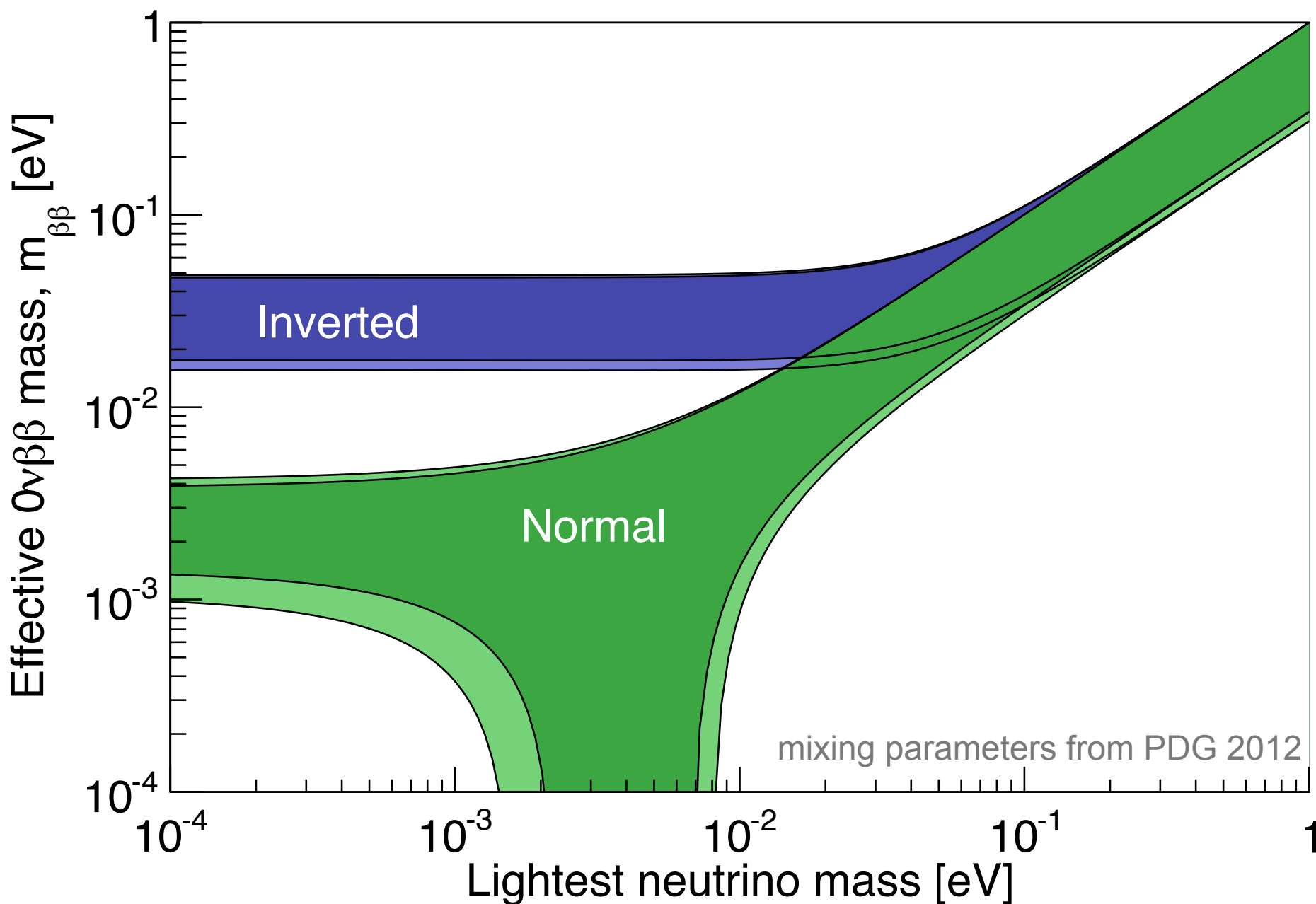




# $0\nu\beta\beta$ and neutrino mass

**decay rate:**  $[T_{1/2}^{0\nu\beta\beta}]^{-1} = G^{0\nu\beta\beta}(E_0, Z) (M^{0\nu\beta\beta})^2 \langle m_{0\nu\beta\beta} \rangle^2$

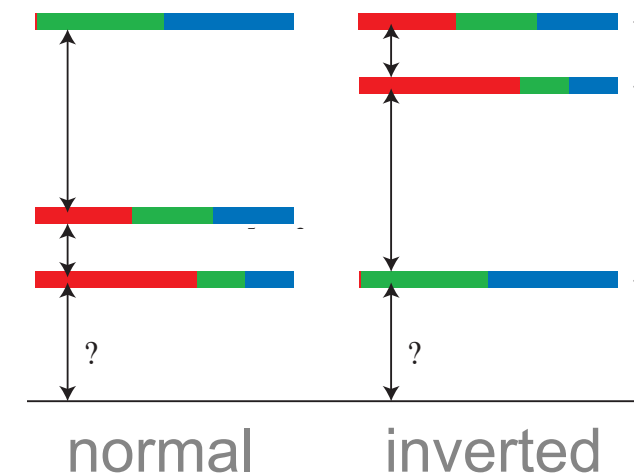
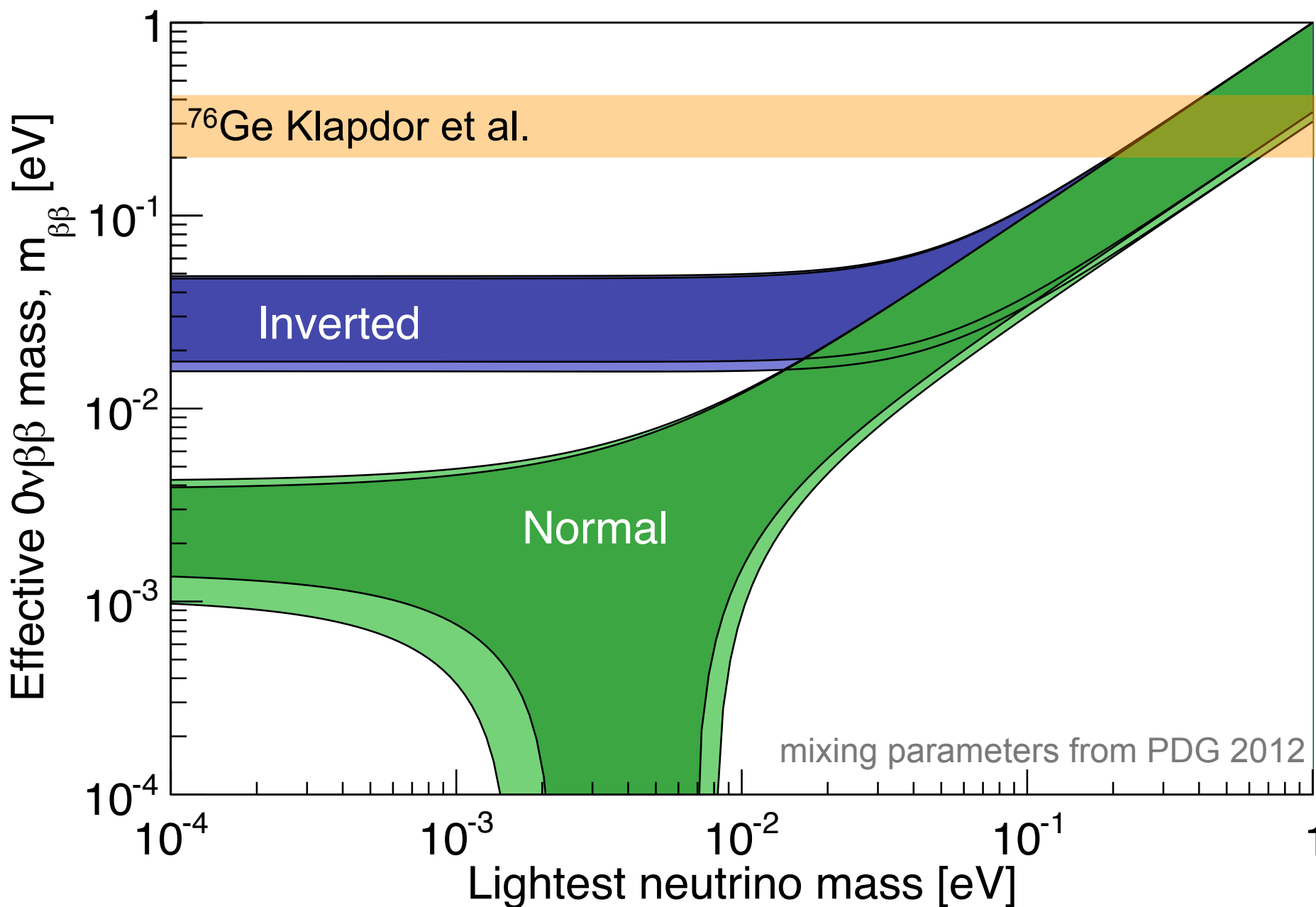
**eff. mass:**  $m_{0\nu\beta\beta} = \left| |U_{e1}|^2 m_1 + |U_{e2}|^2 m_2 e^{i\phi_2} + |U_{e3}|^2 m_3 e^{i\phi_3} \right|$



# $0\nu\beta\beta$ and neutrino mass

**decay rate:**  $[T_{1/2}^{0\nu\beta\beta}]^{-1} = G^{0\nu\beta\beta}(E_0, Z) (M^{0\nu\beta\beta})^2 \langle m_{0\nu\beta\beta} \rangle^2$

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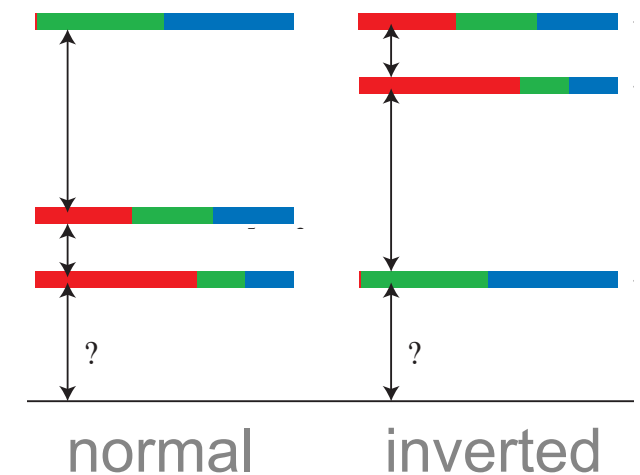
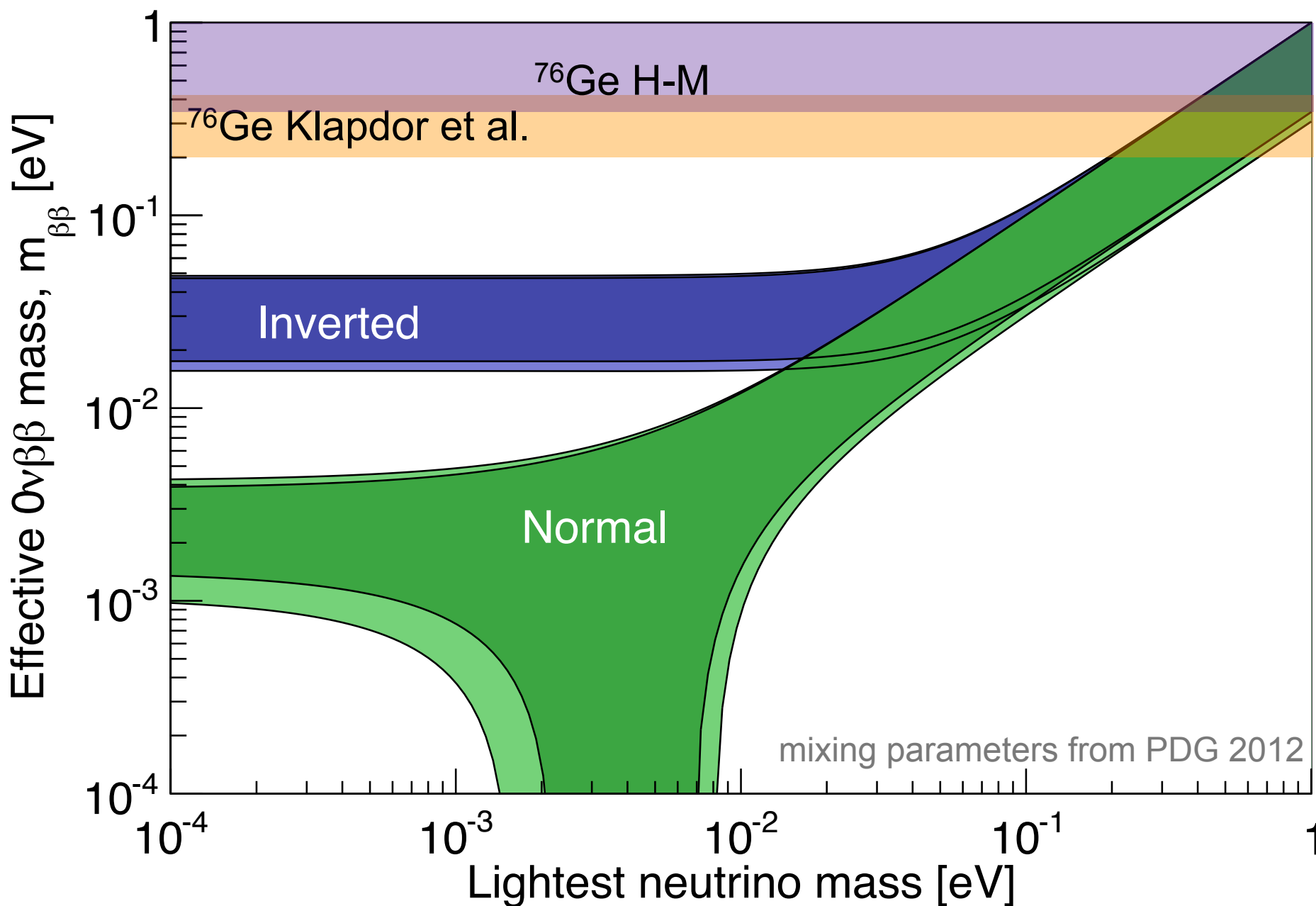
Klapdor: Mod. Phys. Lett. A21 (2006) 1547  
 HM: Eur Phys. Journal A12 (2001) 147  
 EXO-200: Phys. Rev. Lett. 109 (2012) 032505  
 KamLAND-Zen: Phys. Rev. C 86 (2012) 021601



# $0\nu\beta\beta$ and neutrino mass

**decay rate:**  $[T_{1/2}^{0\nu\beta\beta}]^{-1} = G^{0\nu\beta\beta}(E_0, Z) (M^{0\nu\beta\beta})^2 \langle m_{0\nu\beta\beta} \rangle^2$

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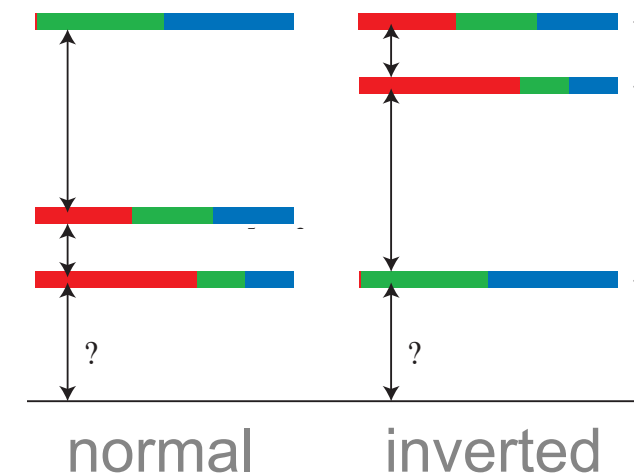
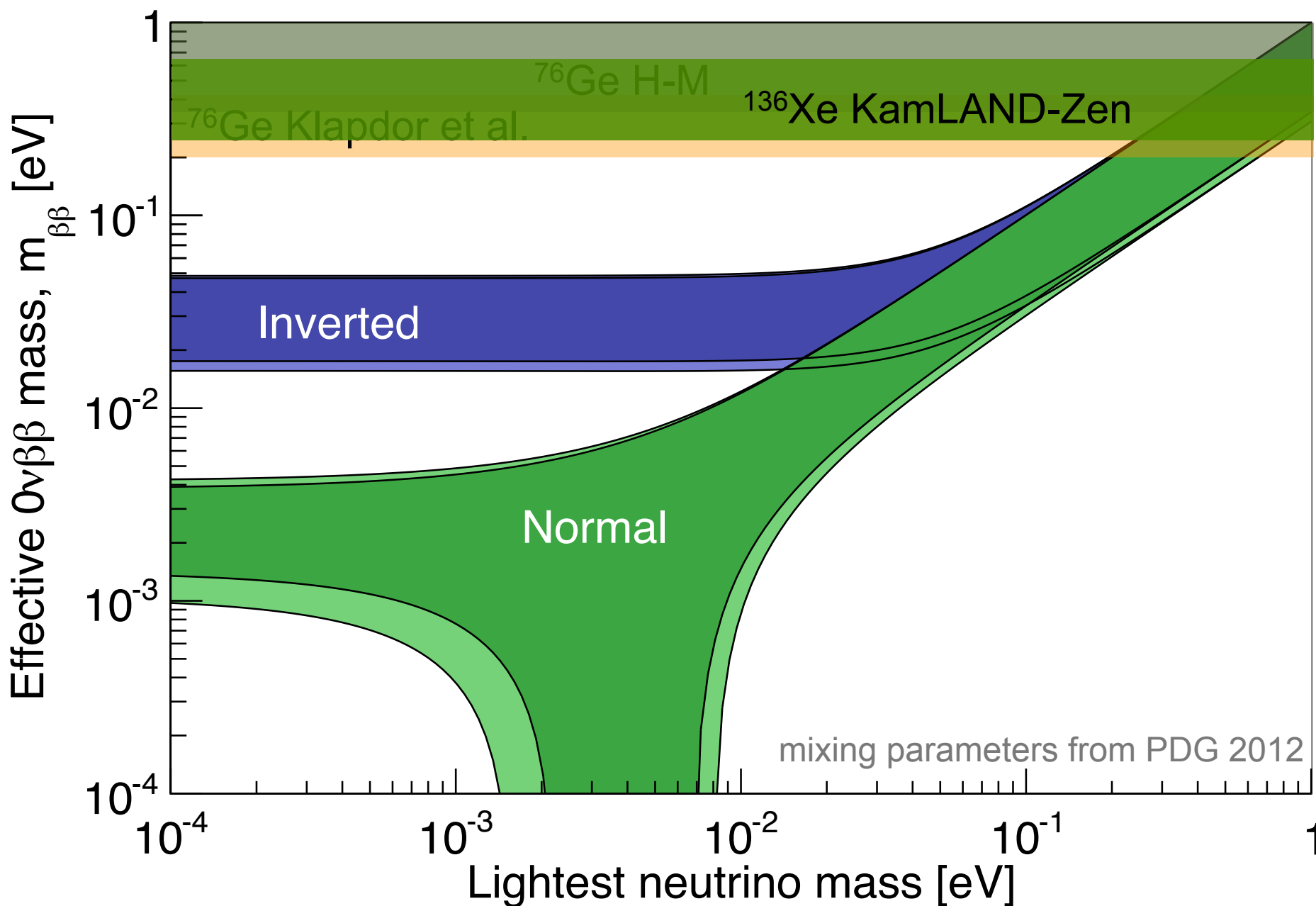


Klapdor: Mod. Phys. Lett. A21 (2006) 1547  
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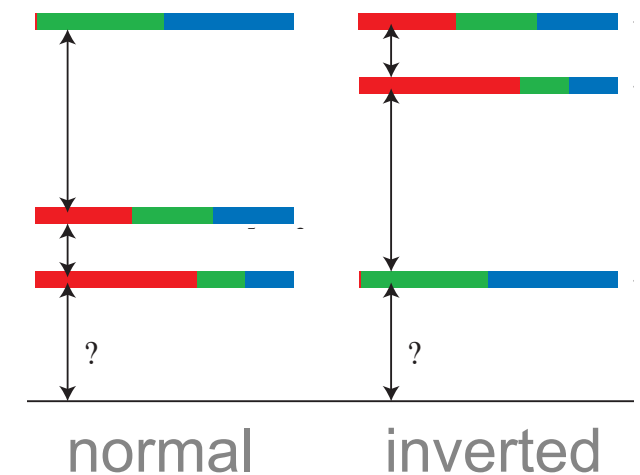
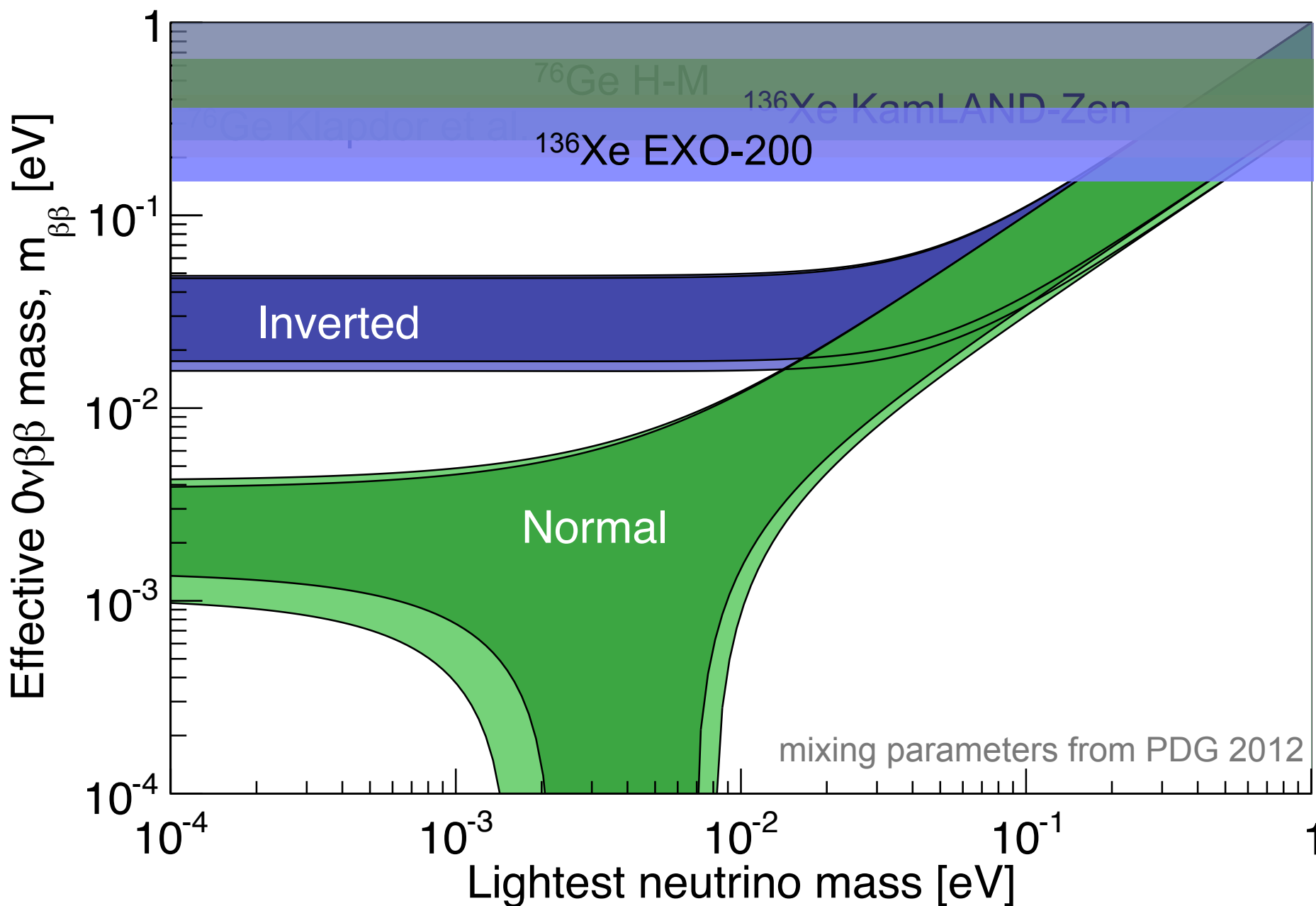
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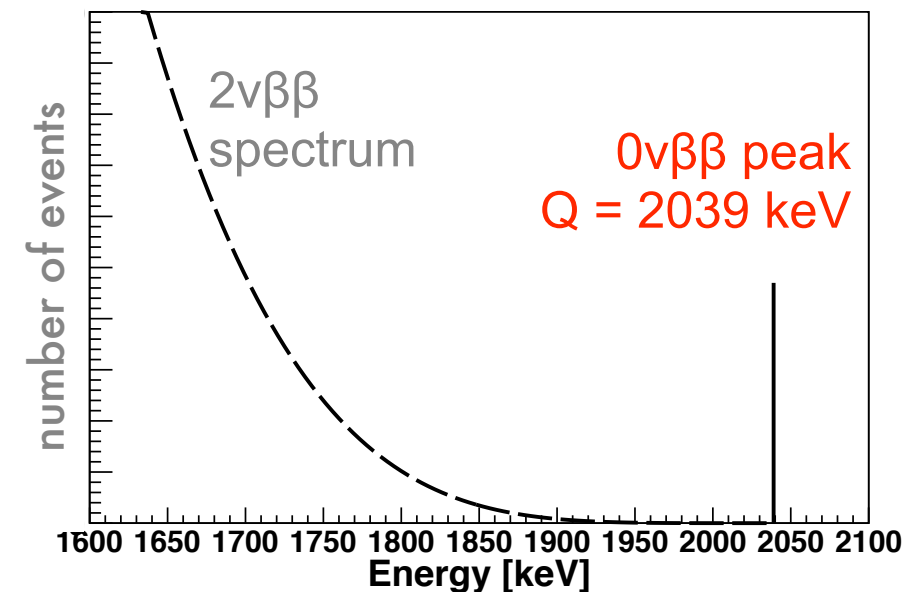


Klapdor: Mod. Phys. Lett. A21 (2006) 1547  
 HM: Eur Phys. Journal A12 (2001) 147  
 EXO-200: Phys. Rev. Lett. 109 (2012) 032505  
 KamLAND-Zen: Phys. Rev. C 86 (2012) 021601

# Criteria for $0\nu\beta\beta$ experiment

- Large mass of source
- Extremely low background rate
- Best possible background identification techniques

Sum of electron energies ( $^{76}\text{Ge}$ )



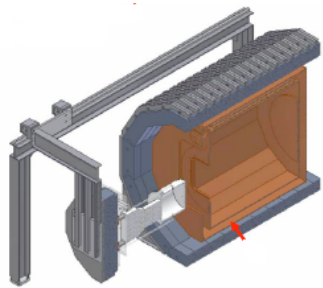
# Criteria for $0\nu\beta\beta$ observation

- Peak at the correct energy
- Full energy spectrum, including backgrounds, understood
- Observe in several different isotopes in independent experiments

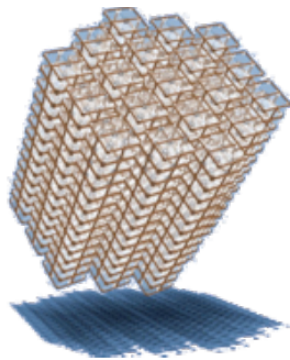


# $0\nu\beta\beta$ experiments

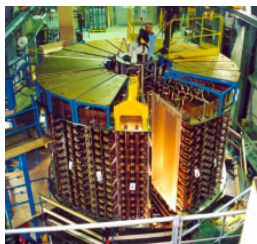
EXO-200



CUORE



NEMO



Collaboration	Isotope	Technique	mass ( $0\nu\beta\beta$ isotope)	Status
CANDLES	Ca-48	305 kg $\text{CaF}_2$ crystals - liq. scint	0.3 kg	Construction
CARVEL	Ca-48	$^{48}\text{CaWO}_4$ crystal scint.		
GERDA I	Ge-76	Ge diodes in LAr	15 kg	Operating
II		Point contact Ge in LAr or LN	30-35 kg	Construction
MAJORANA DEMONSTRATOR	Ge-76	Point contact Ge	30 kg	Construction
1TGe (GERDA & MAJORANA)	Ge-76	Best technology from GERDA and MAJORANA	~ tonne	R&D
NEMO3	Mo-100 Se-82	Foils with tracking	6.9 kg 0.9 kg	Complete
SuperNEMO Demonstrator	Se-82	Foils with tracking	7 kg	R&D
MOON	Mo-100	Mo sheets	200 kg	R&D
CAMEO	Cd-116	$\text{CdWO}_4$ crystals		
COBRA	Cd-116, Te-130	$\text{CdZnTe}$ detectors	10 kg	R&D
CUORICINO	Te-130	$\text{TeO}_2$ Bolometer	10 kg	Complete
CUORE	Te-130	$\text{TeO}_2$ Bolometer	206 kg	Construction
KamLAND-ZEN	Xe-136	2.7% in liquid scint.	380 kg	Operating
NEXT-100	Xe-136	High pressure Xe TPC	80 kg	R&D
EXO200	Xe-136	Xe liquid TPC	160 kg	Operating
EXO	Xe-136	Xe liquid TPC	~ tonne	R&D
DCBA	Nd-150	Nd foils & tracking chambers	< kg	R&D
SNO+	Nd-150	0.1% $^{150}\text{Nd}$ suspended in Scint	44 kg	Construction

Operating

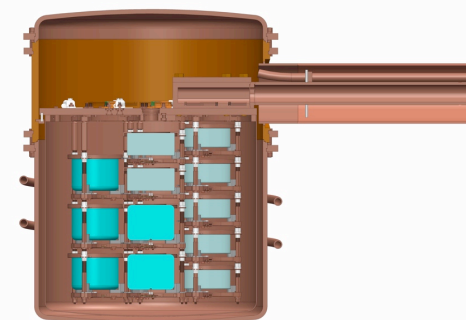
Commissioning

Construction

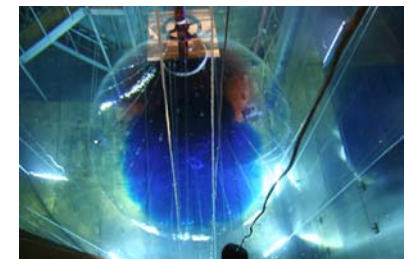
GERDA



MAJORANA

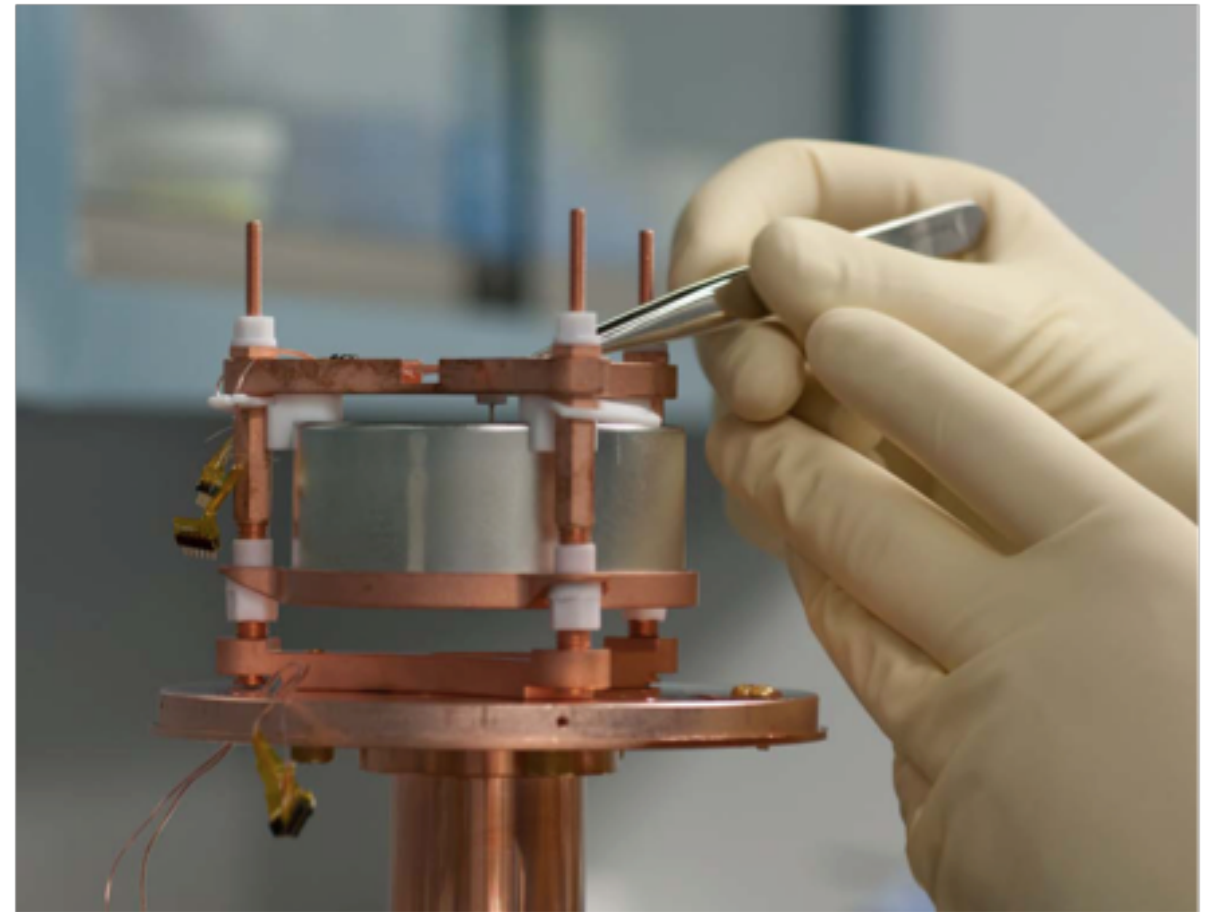


KamLAND-Zen



# germanium detectors

- Detector is source: demonstrated ability to enrich from 7.4% to 86%  $^{76}\text{Ge}$
- Ge diodes are intrinsically high purity
- Excellent energy resolution: 0.13% FWHM at Q-value of 2039 keV
- Commercially available
- P-type point contact detectors
  - extremely low noise
  - low energy threshold



A stainless-steel detector blank in a prototype MAJORANA detector mount





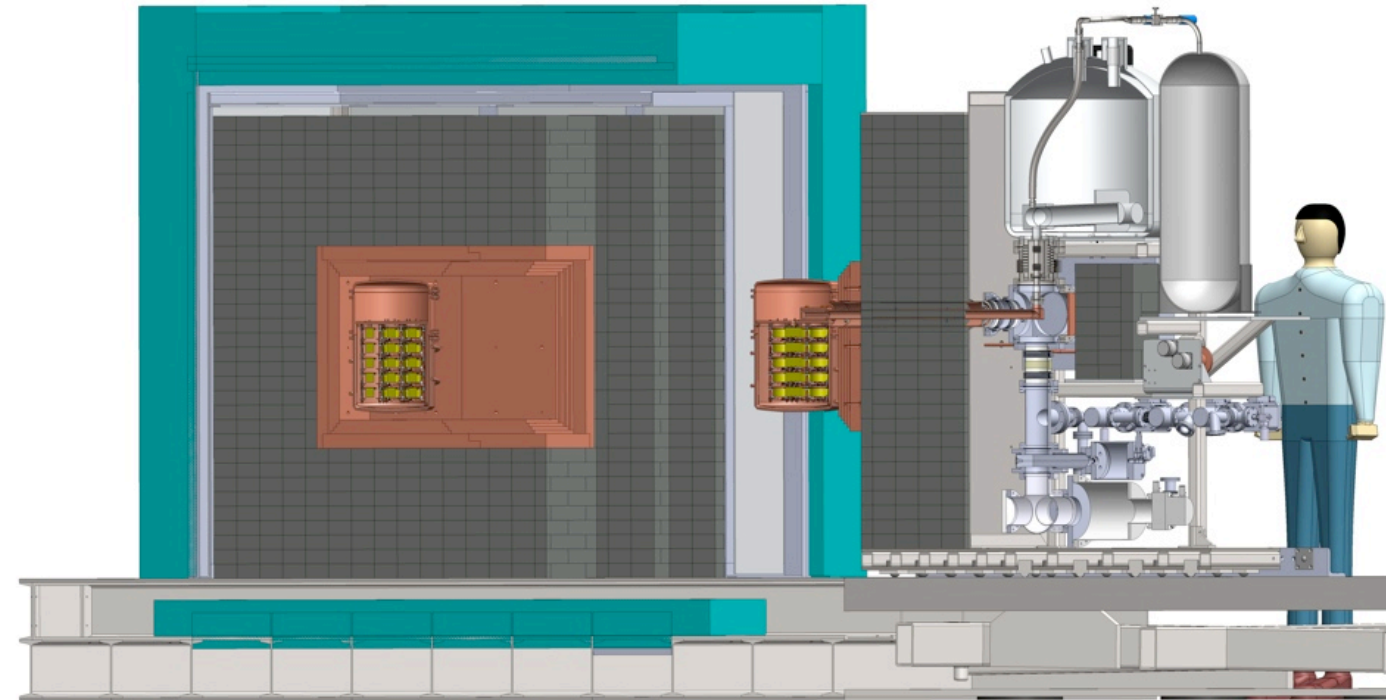
# GERDA



# MAJORANA



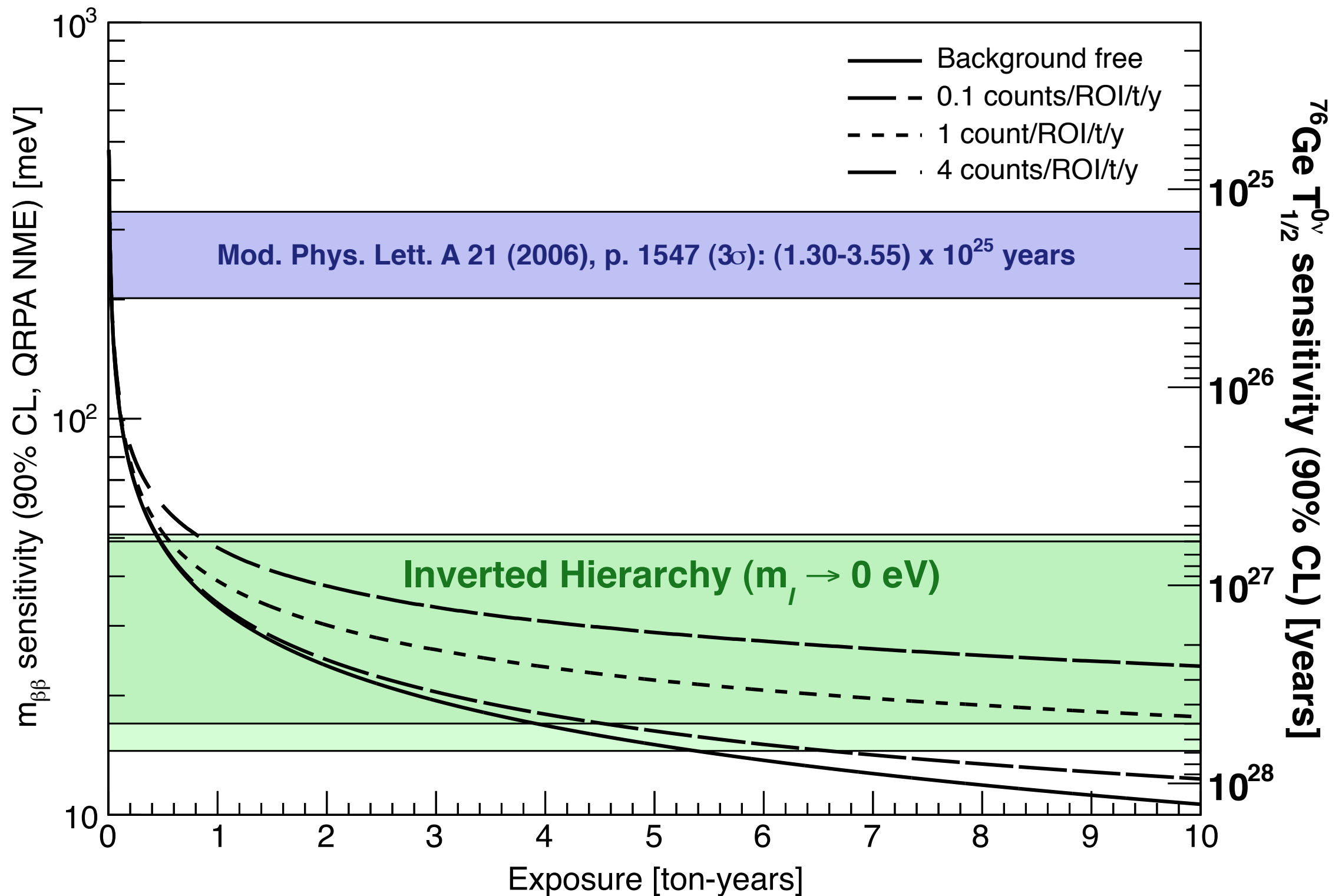
- **Design:** detectors submerged in liquid Argon at LNGS, Italy
- **Shield:** LAr, H<sub>2</sub>O
- **Phase I:** 18 kg enr-Ge (2011)
- **Phase II:** 20 kg enr-Ge (2013)



- **Design:** detectors in high-purity electroformed copper cryostats at Sanford Lab, US
- **Shield:** copper, lead
- **DEMONSTRATOR:** 30 kg of enr-Ge

Open exchange of knowledge and technologies  
**Future goal:** merge for tonne-scale experiment

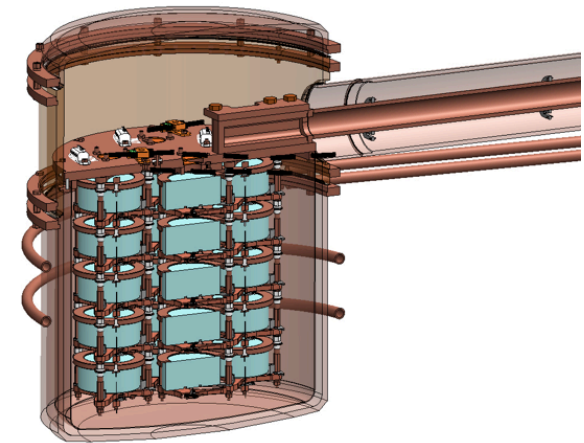
# Sensitivity of a tonne-scale $^{76}\text{Ge}$ experiment



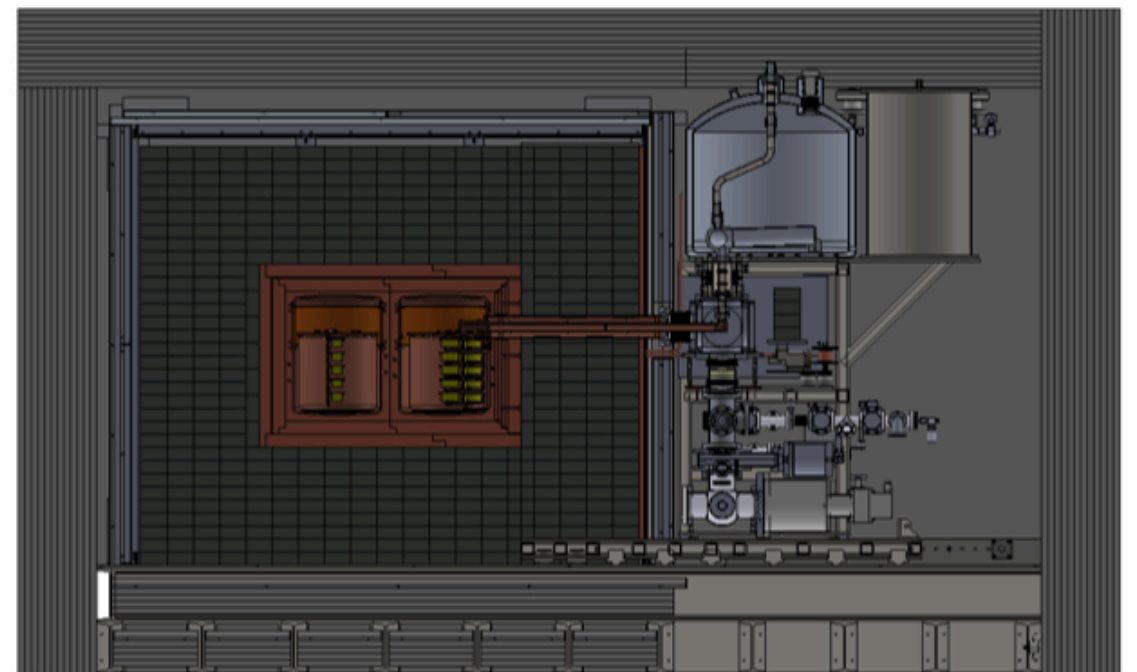
# The MAJORANA DEMONSTRATOR

Funded by DOE Office of Nuclear Physics and NSF Particle and Nuclear Astrophysics,  
with additional contributions from international collaborators.

- Goals:**
- Demonstrate backgrounds low enough to justify building a tonne scale experiment
  - Establish feasibility to construct & field modular arrays of Ge detectors
  - Test Klapdor-Kleingrothaus claim
  - Low-energy dark matter (light WIMPs) search

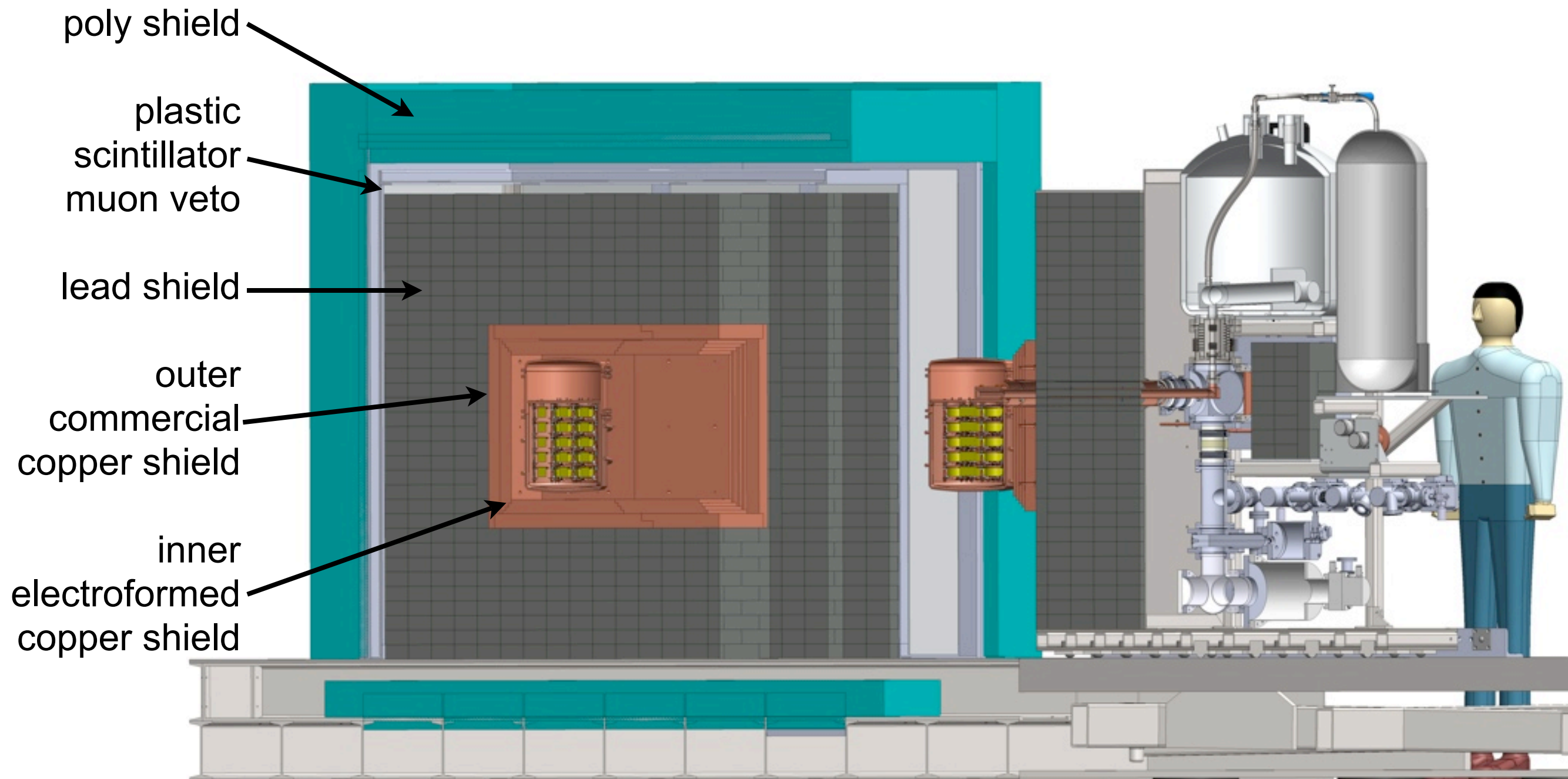


- **Located underground at 4850' Sanford Lab**
- **Background Goal in the  $0\nu\beta\beta$  peak region of interest (4 keV at 2039 keV)**  
**4 counts/ROI/t/y (after analysis cuts)**  
*scales to 1 count/ROI/t/y for a tonne experiment*
- **40-kg of Ge detectors**
  - 30-kg of 86% enriched  $^{76}\text{Ge}$  crystals & 10-kg of  $^{\text{nat}}\text{Ge}$
  - Detector Technology: P-type, point-contact.
- **2 independent cryostats**
  - ultra-clean, electroformed Cu
  - 20 kg of detectors per cryostat
  - naturally scalable
- **Compact Shield**
  - low-background passive Cu and Pb shield with active muon veto





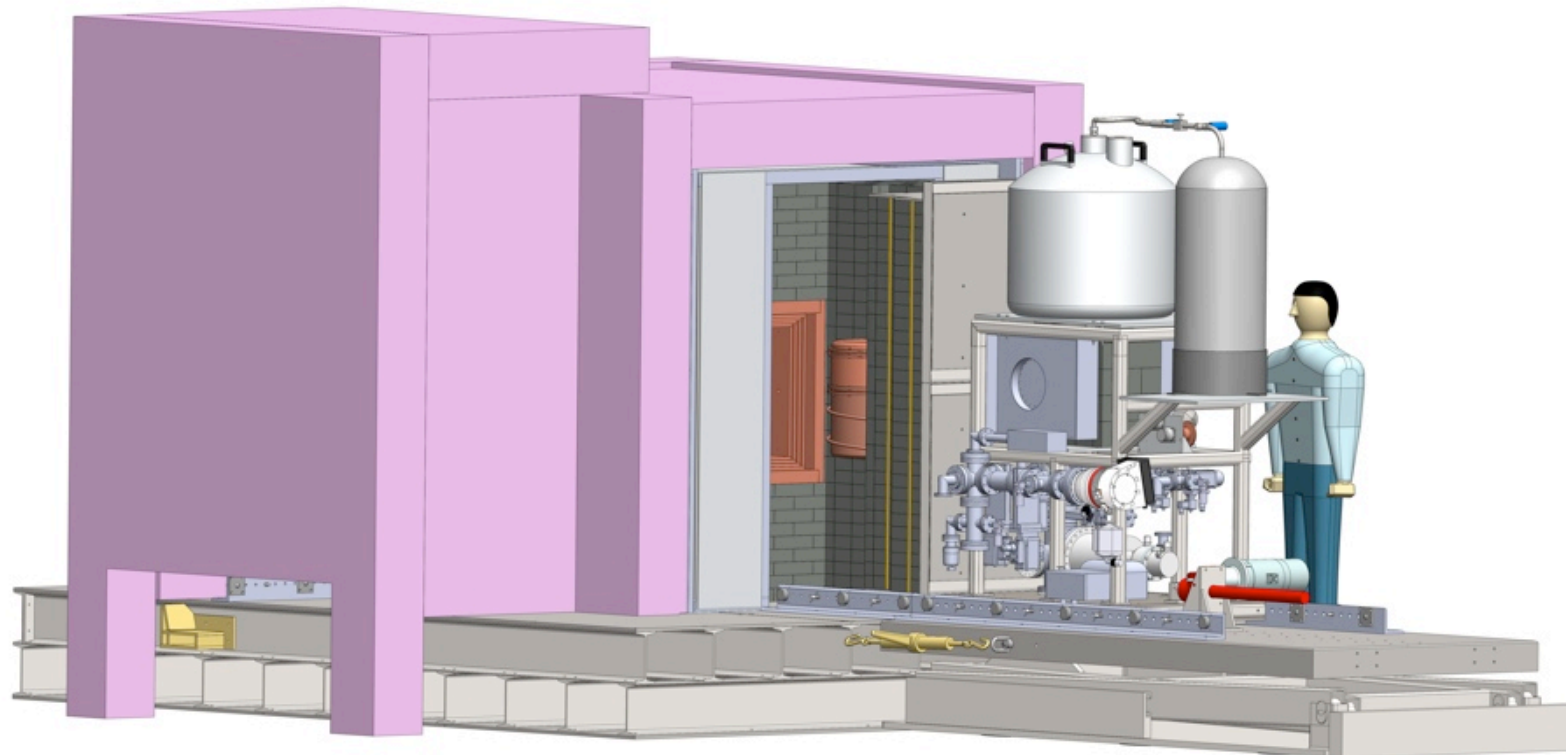
# The MAJORANA DEMONSTRATOR





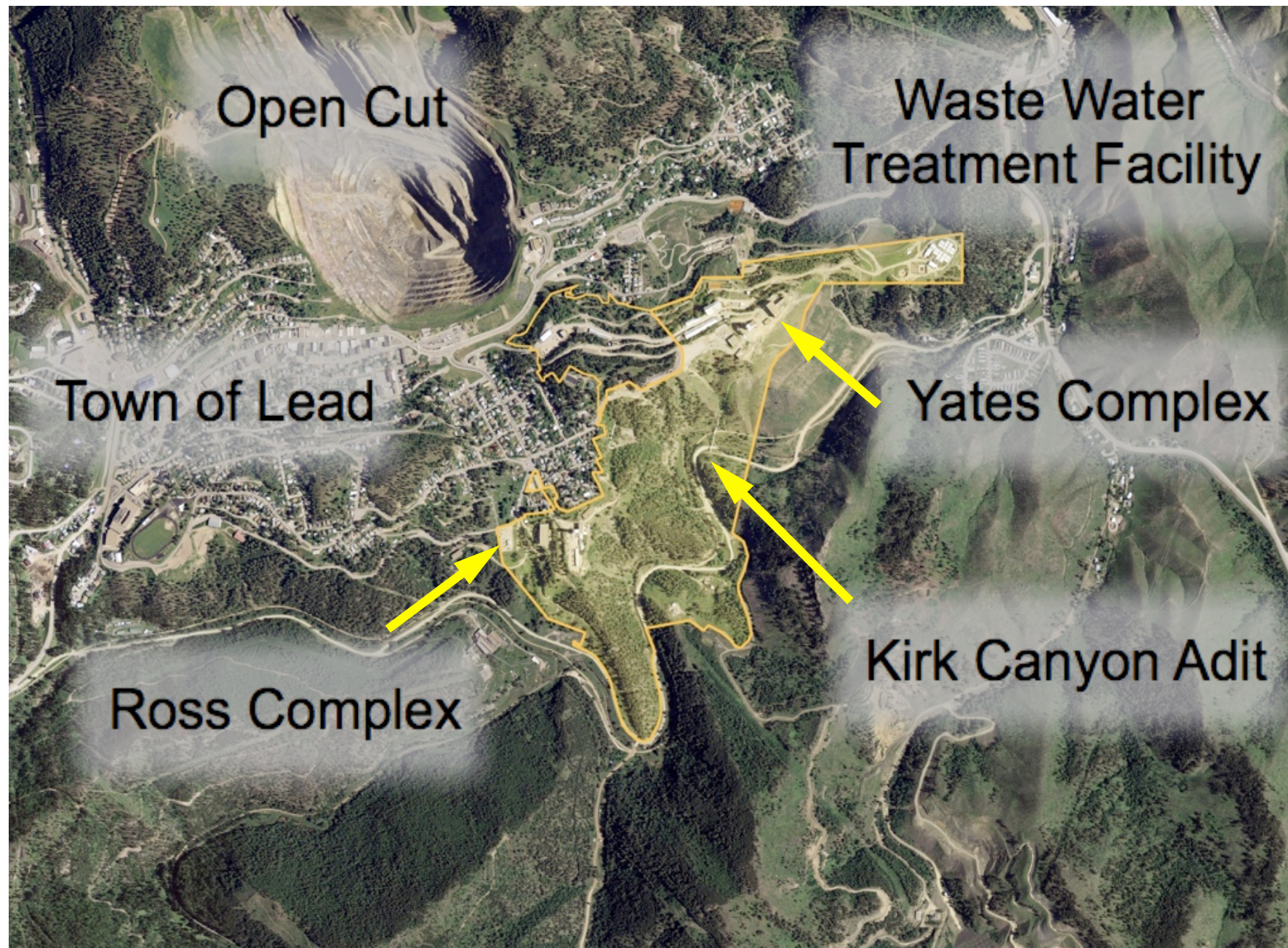
# DEMONSTRATOR Schedule

- **Three steps:**
  - **Prototype cryostat:** 2 strings natural Ge **Early 2013**
  - **Cryostat 1:** 3 strings enr. Ge, 4 strings nat. Ge **Fall 2013**
  - **Cryostat 2:** 7 strings enr. Ge **Fall 2014**





# Sanford Underground Research Facility

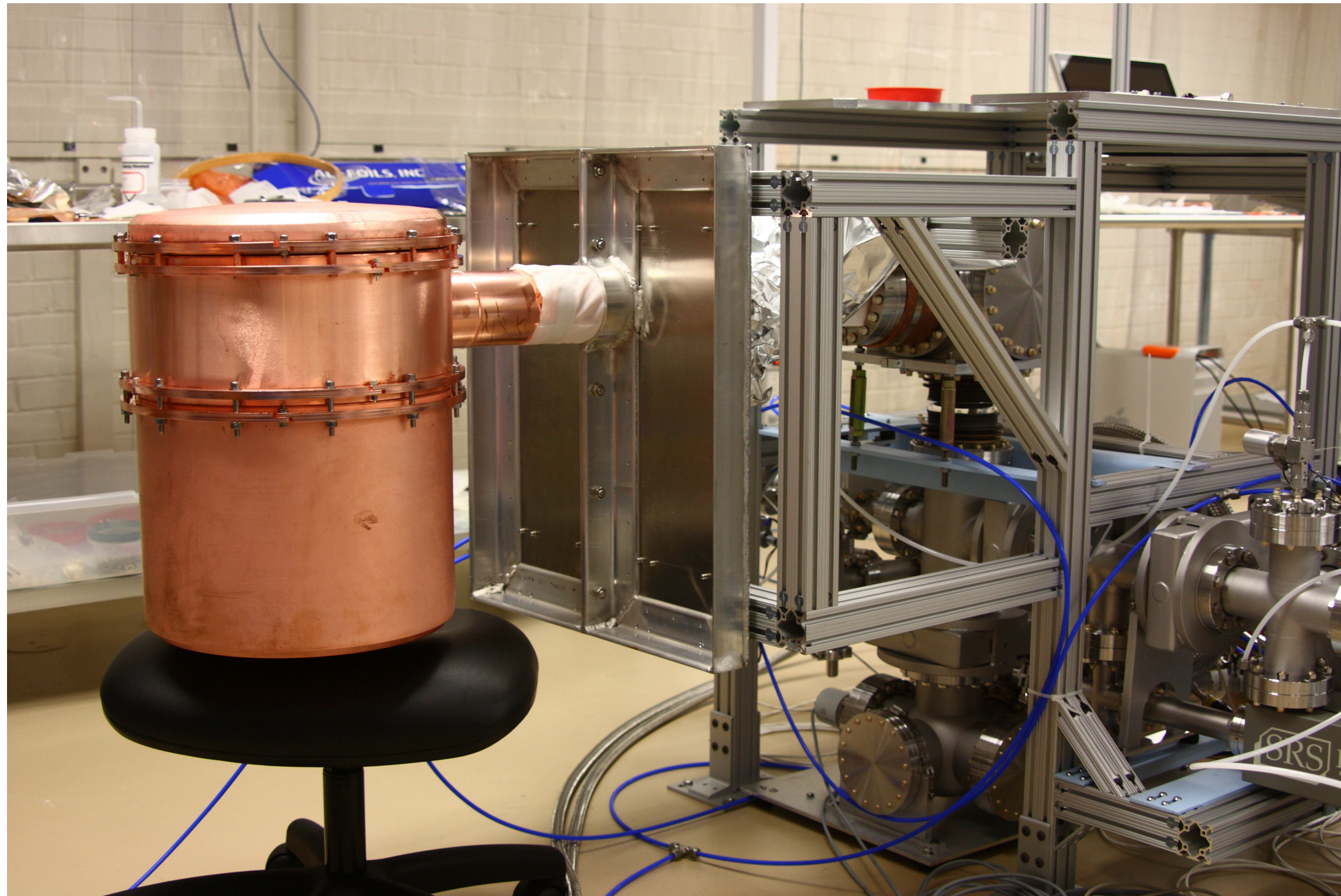
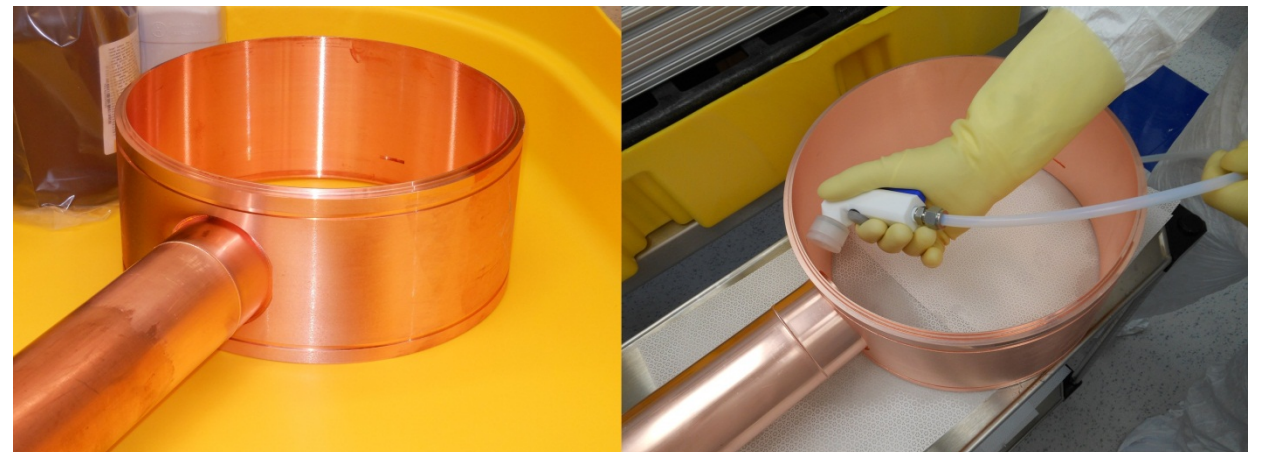


Home of the MAJORANA DEMONSTRATOR, Lead, SD

# Recent DEMONSTRATOR progress

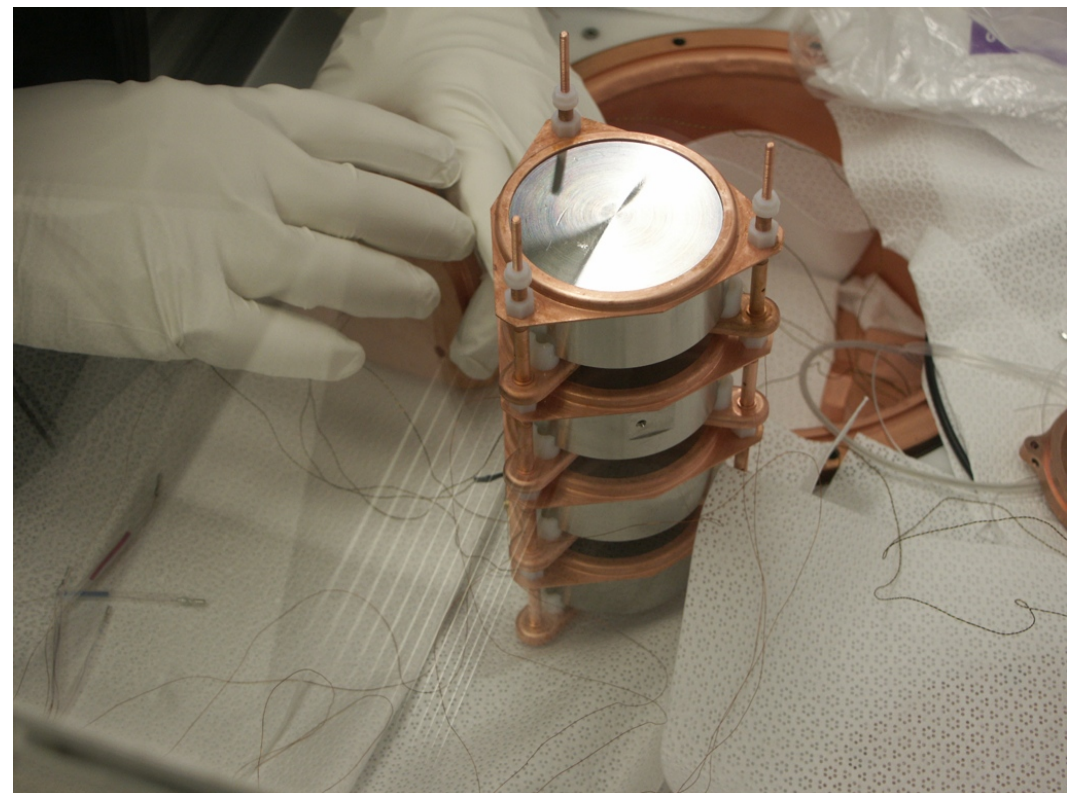
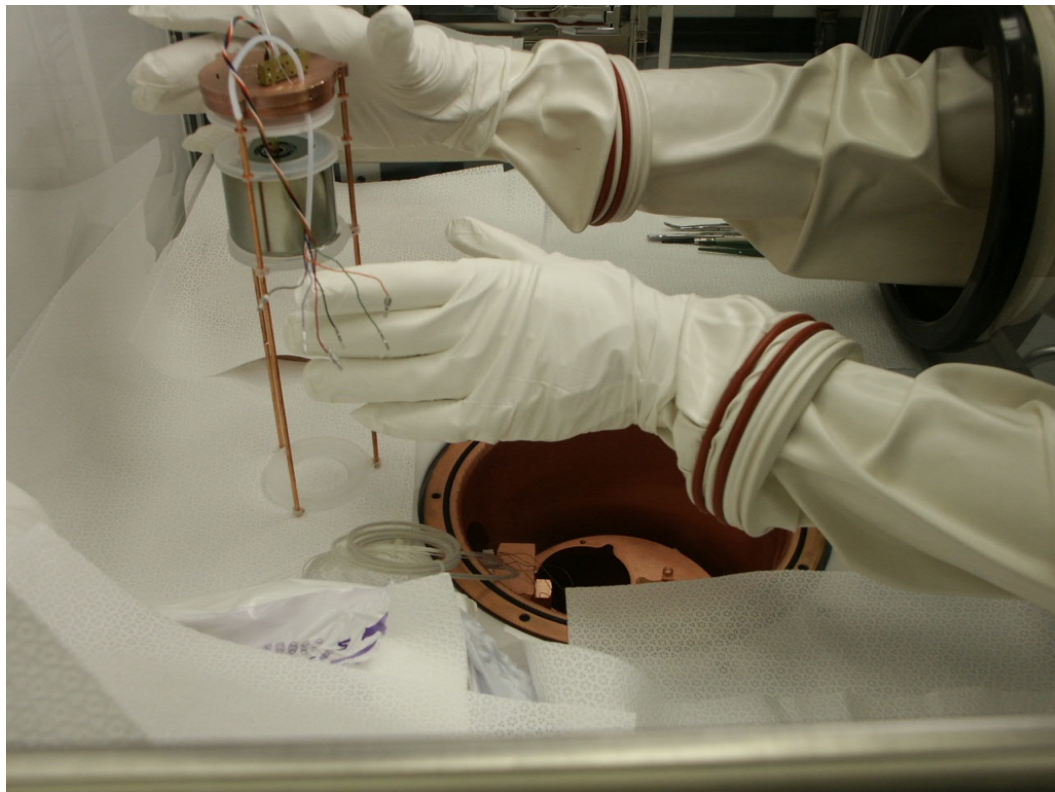


# prototype cryostat



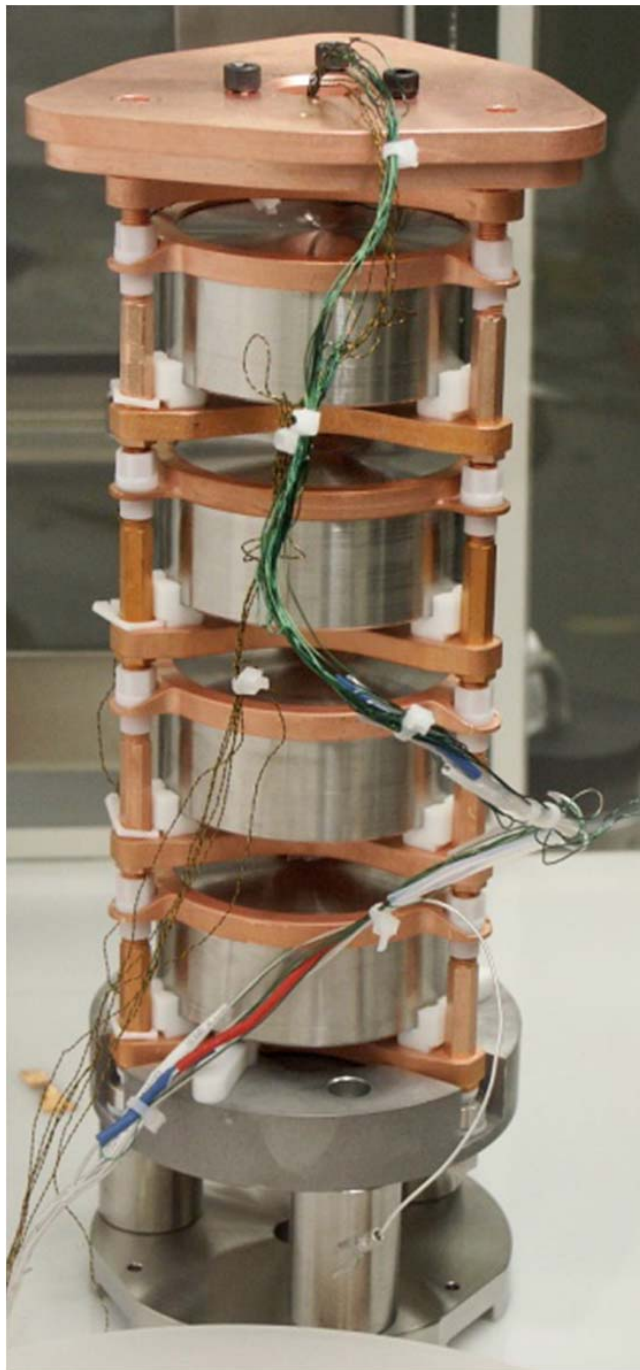


# detector string evolution





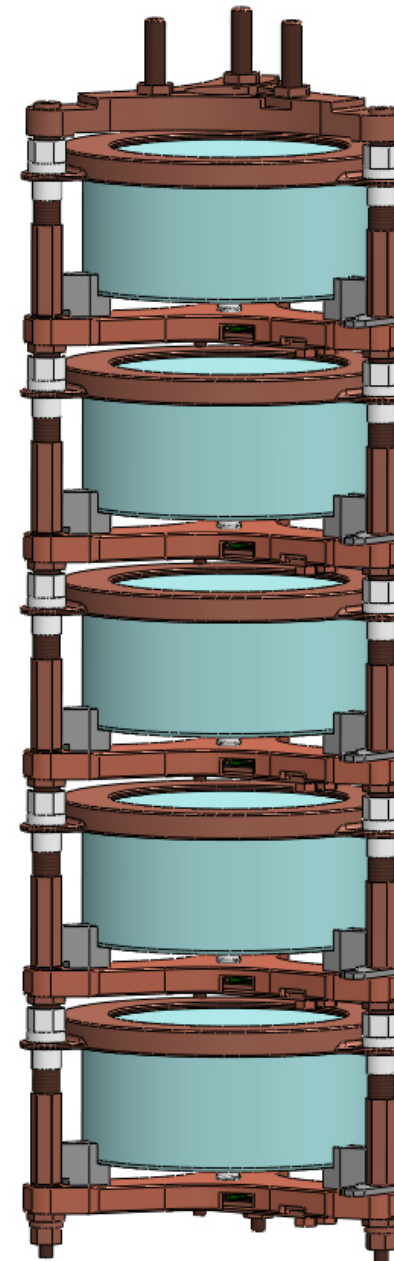
# detector string evolution



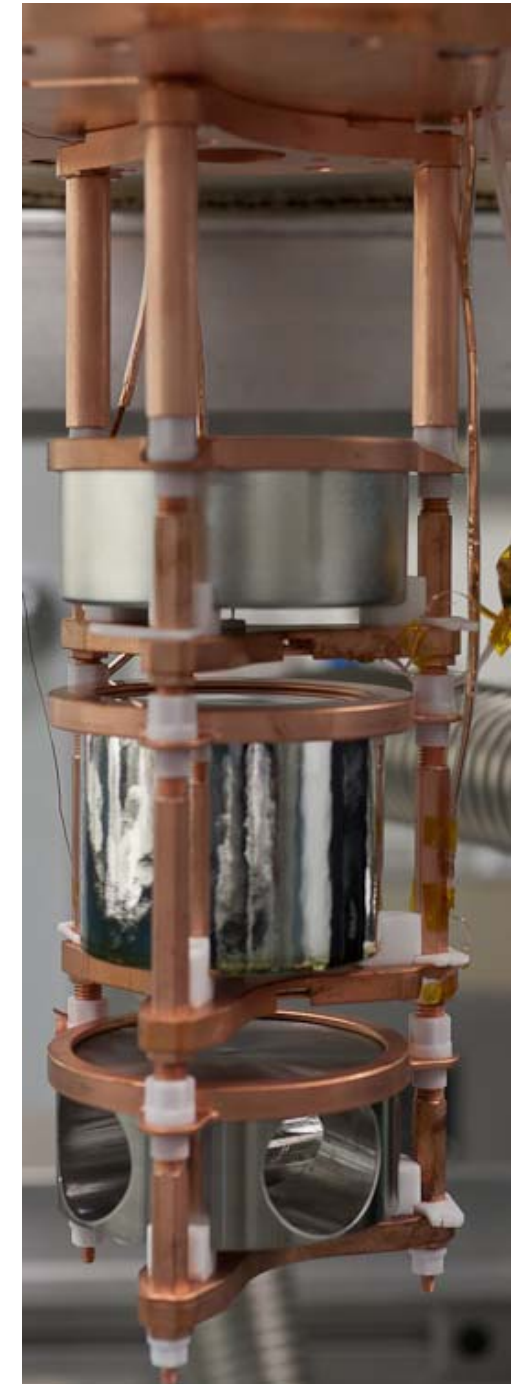
LANL thermal test string  
Jan 2011



LBNL March 2011 thermal  
tests



Design as released for  
Prototype production Feb  
2012: MJ80-02-195



LBNL Jan 2012 tests

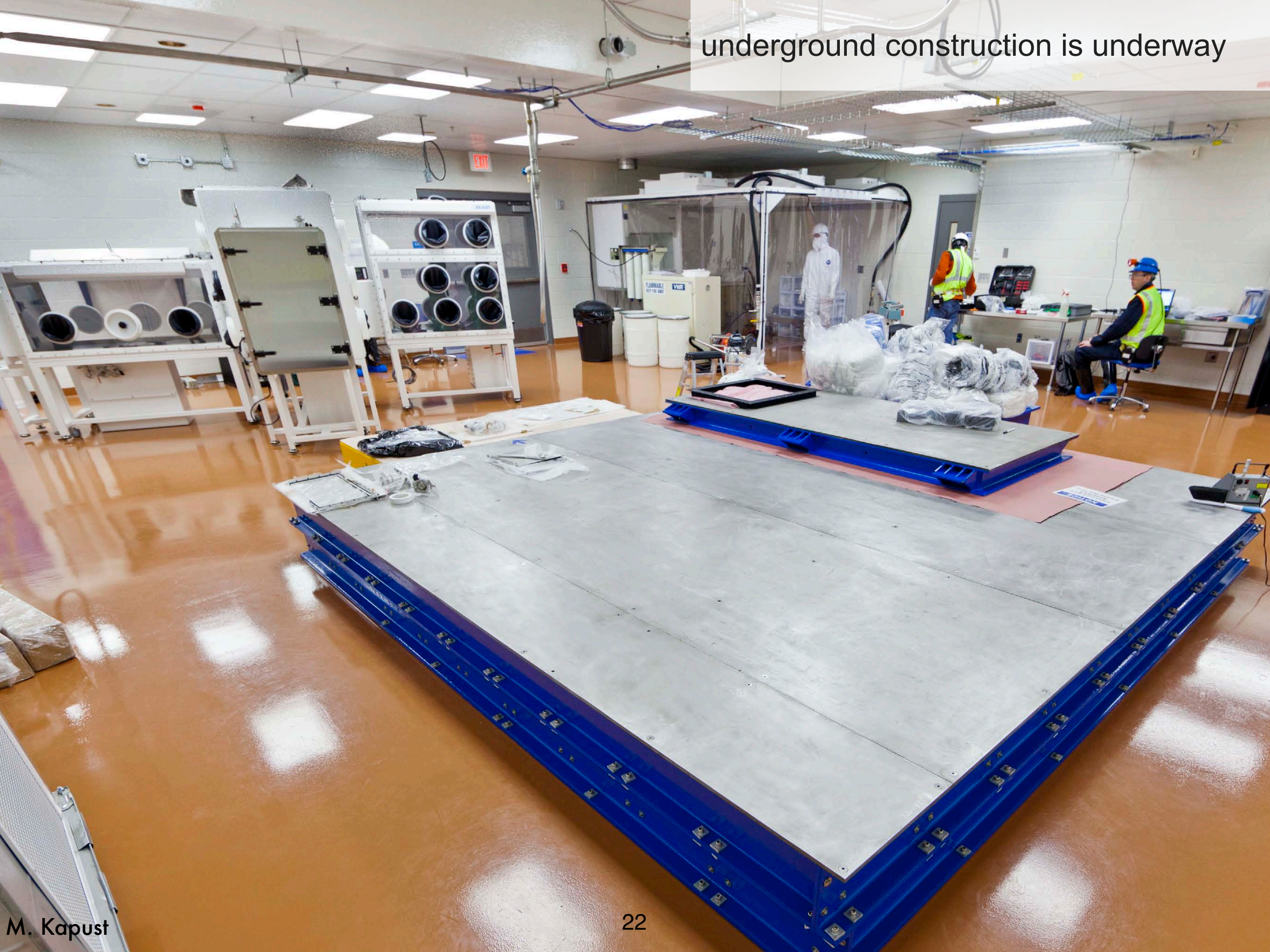


10 baths at Sanford producing copper since July 2011  
about 50% of EFCu complete, including major parts for cryostat 1





underground construction is underway







EXIT

NOTICE

NOTICE

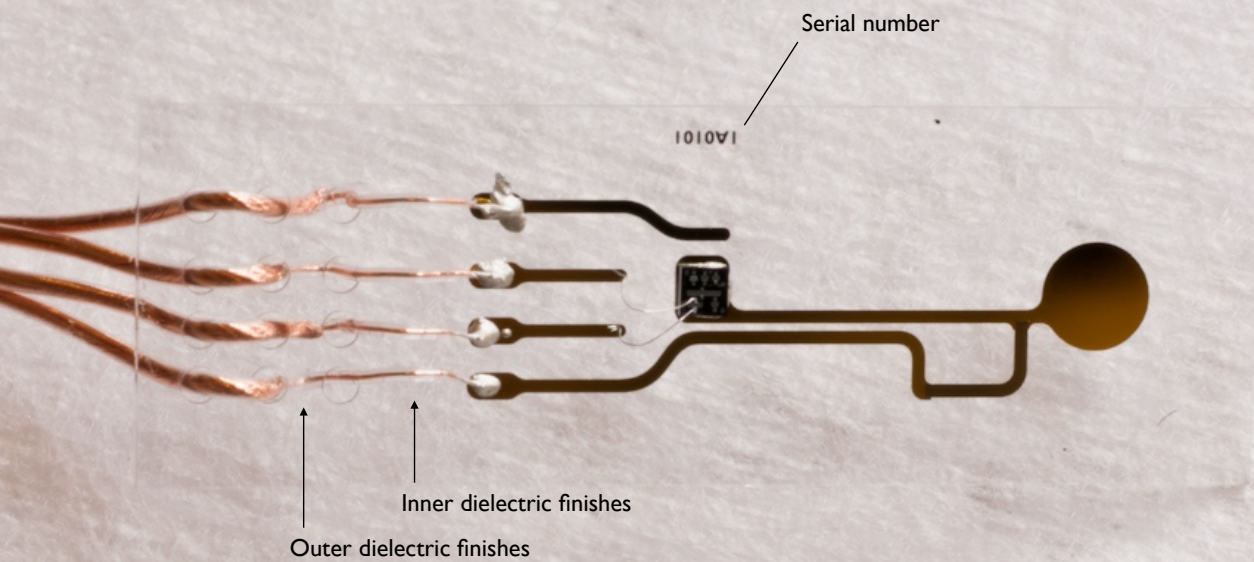






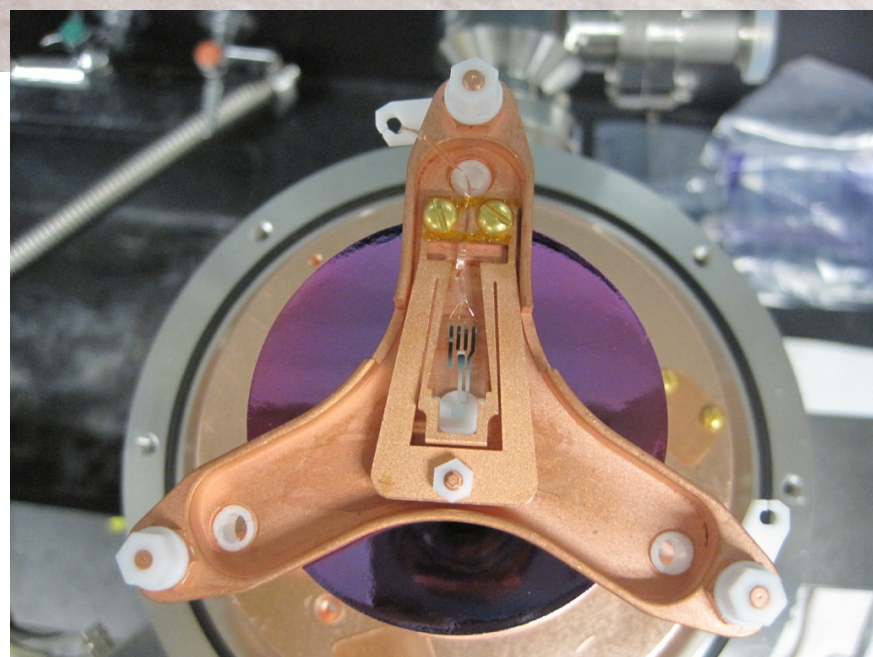
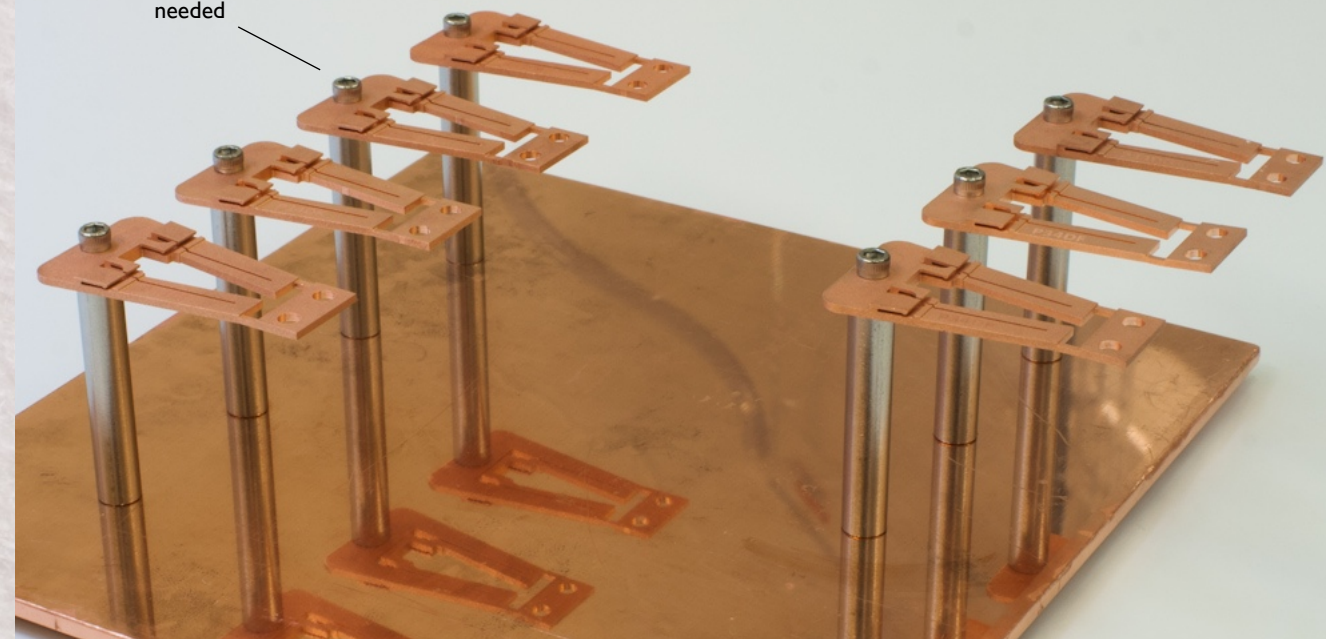
# Detector electronics

LMFE production



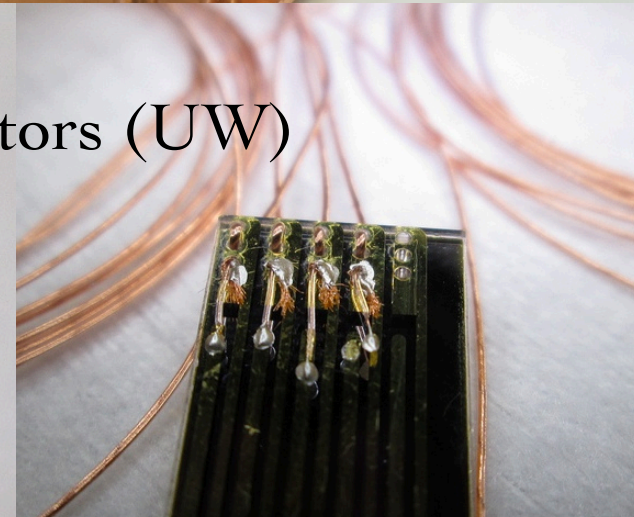
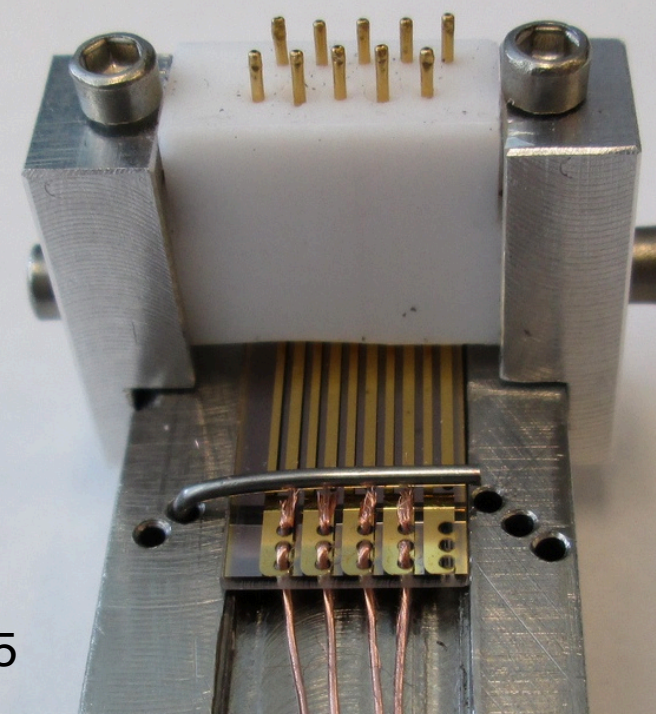
Negligible force  
applied so only  
single screws  
needed

Spring clips

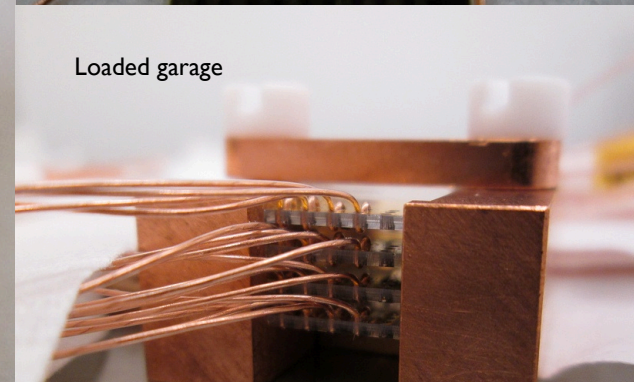


Epoxying jig

Signal connectors (UW)



Loaded garage



images from J. Loach

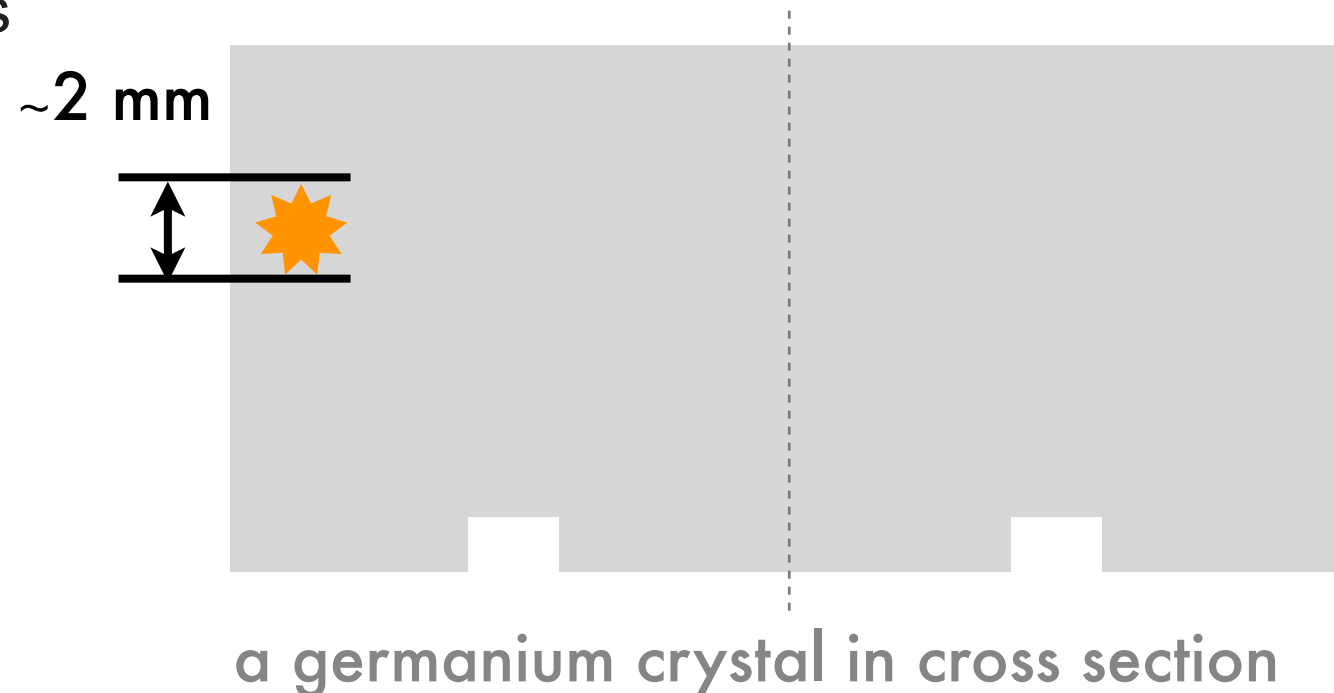
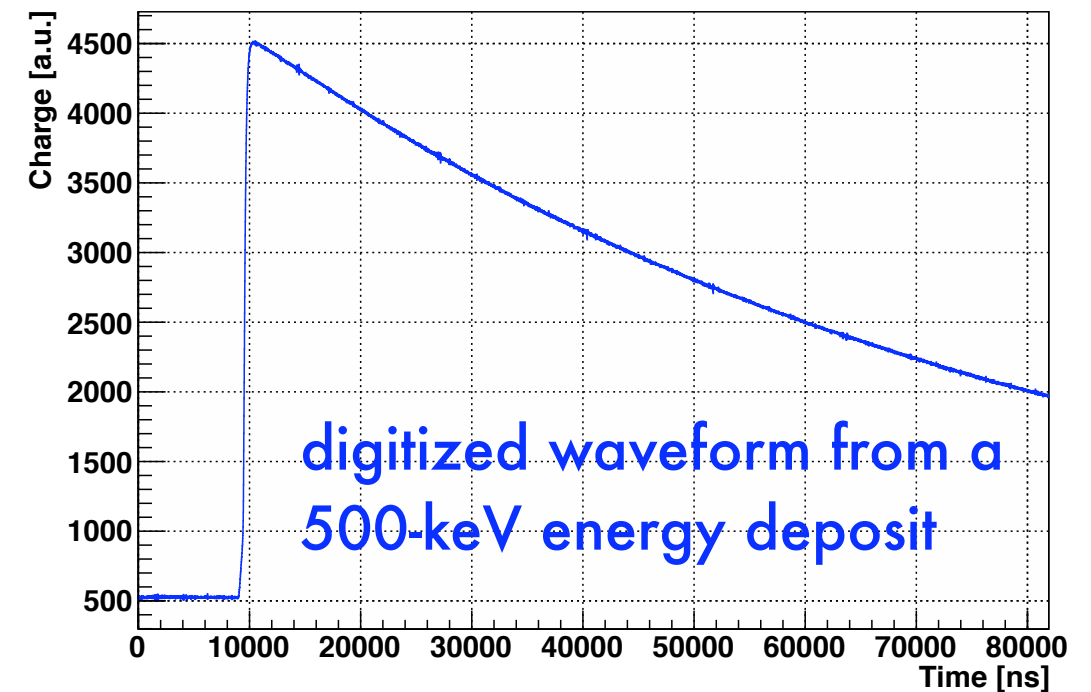


# detection of $0\nu\beta\beta$



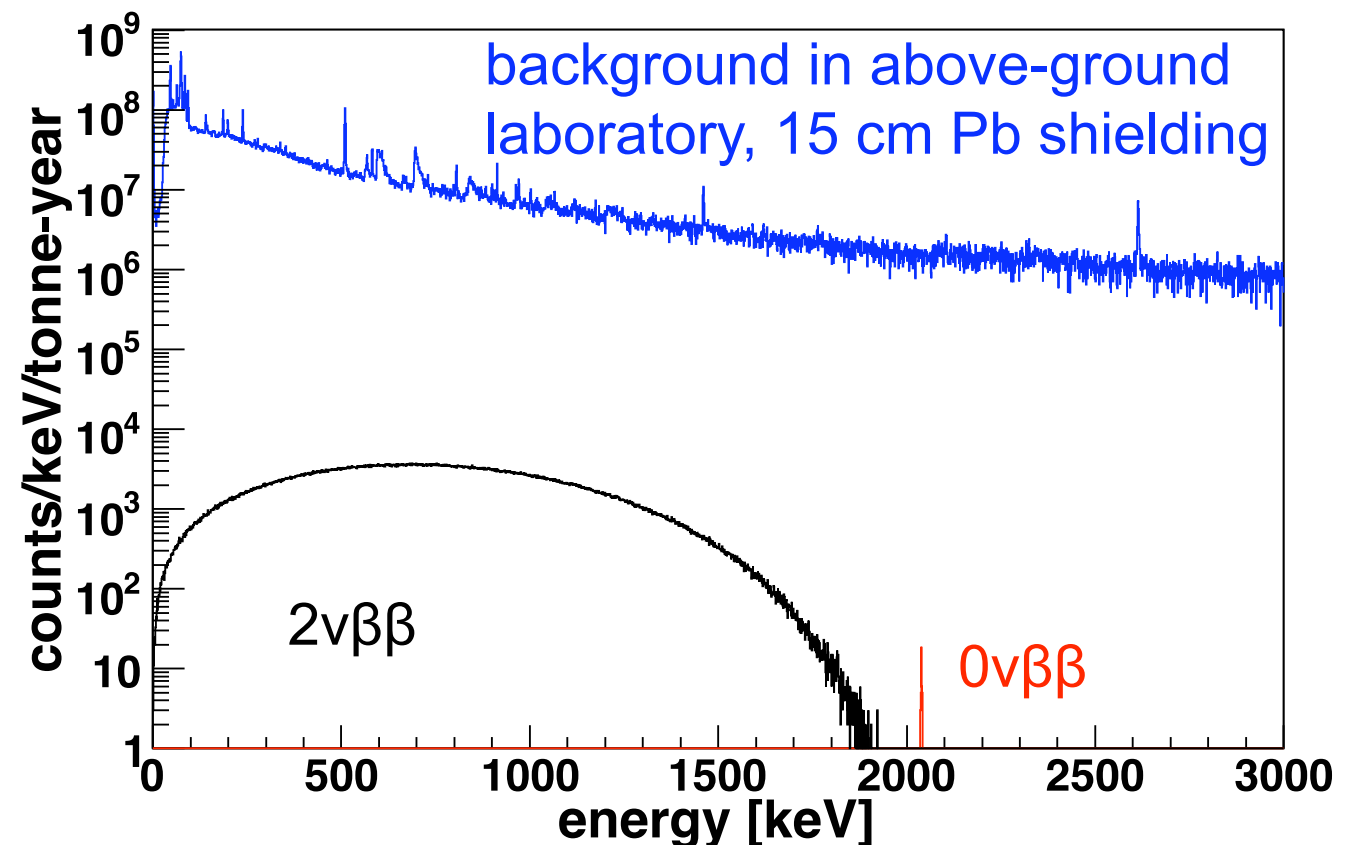
# The $0\nu\beta\beta$ signal in germanium

- Ionizing energy deposits in germanium produce a charge signal
- The  $0\nu\beta\beta$  signal:
  - Single site
  - Single crystal
  - Uncorrelated in time with other events
  - Near 2039-keV  $^{76}\text{Ge}$  endpoint



# Sources of background

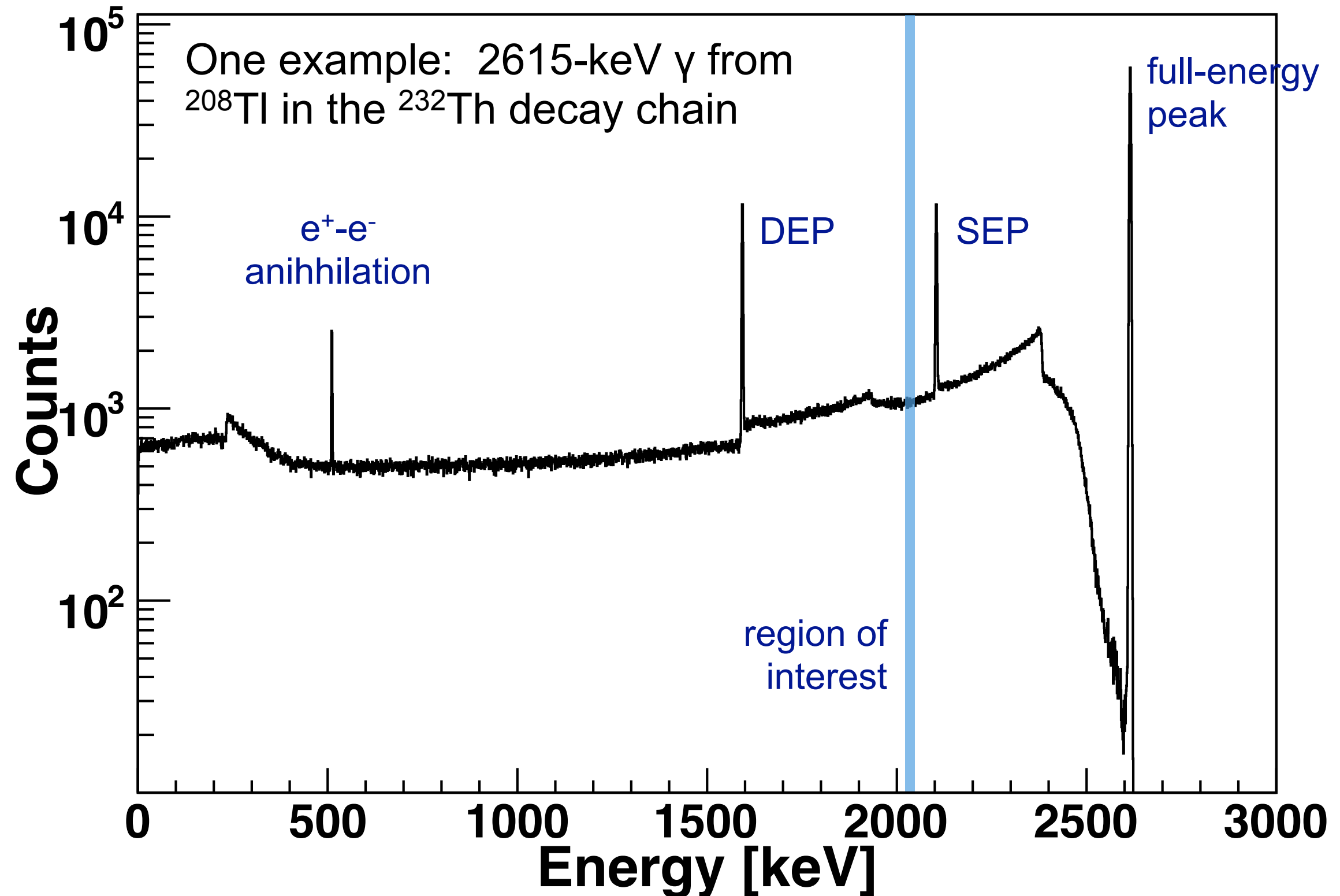
- Primordial contamination:  
 $^{40}\text{K}$ ,  $^{238}\text{U}$ ,  $^{232}\text{Th}$
- Long-lived cosmogenics:  
 $^{68}\text{Ge}$ ,  $^{60}\text{Co}$ ,  $^{65}\text{Zn}$
- Prompt cosmogenics:  
 $\mu$ ,  $\mu$ -induced neutrons
- other: anthropogenic contaminants, radon, solar neutrinos



## Background reduction techniques

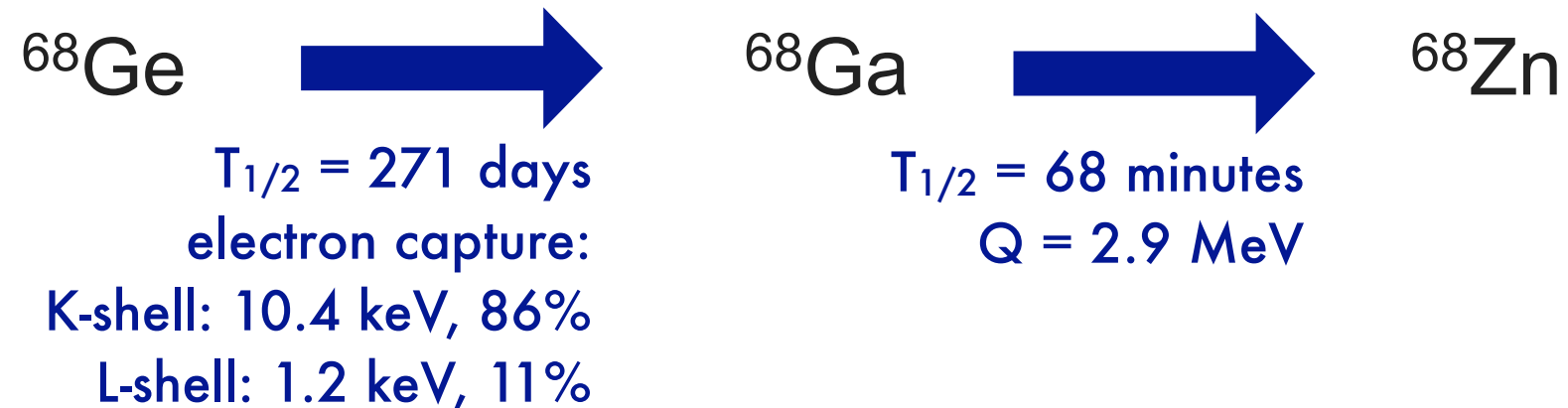
- Minimize mass of non-germanium components
- Use passive and active shielding
- Fabricate parts in shielding from clean copper and plastic
  - Electroform and machine copper underground at Sanford and PNNL

# Backgrounds in germanium detectors

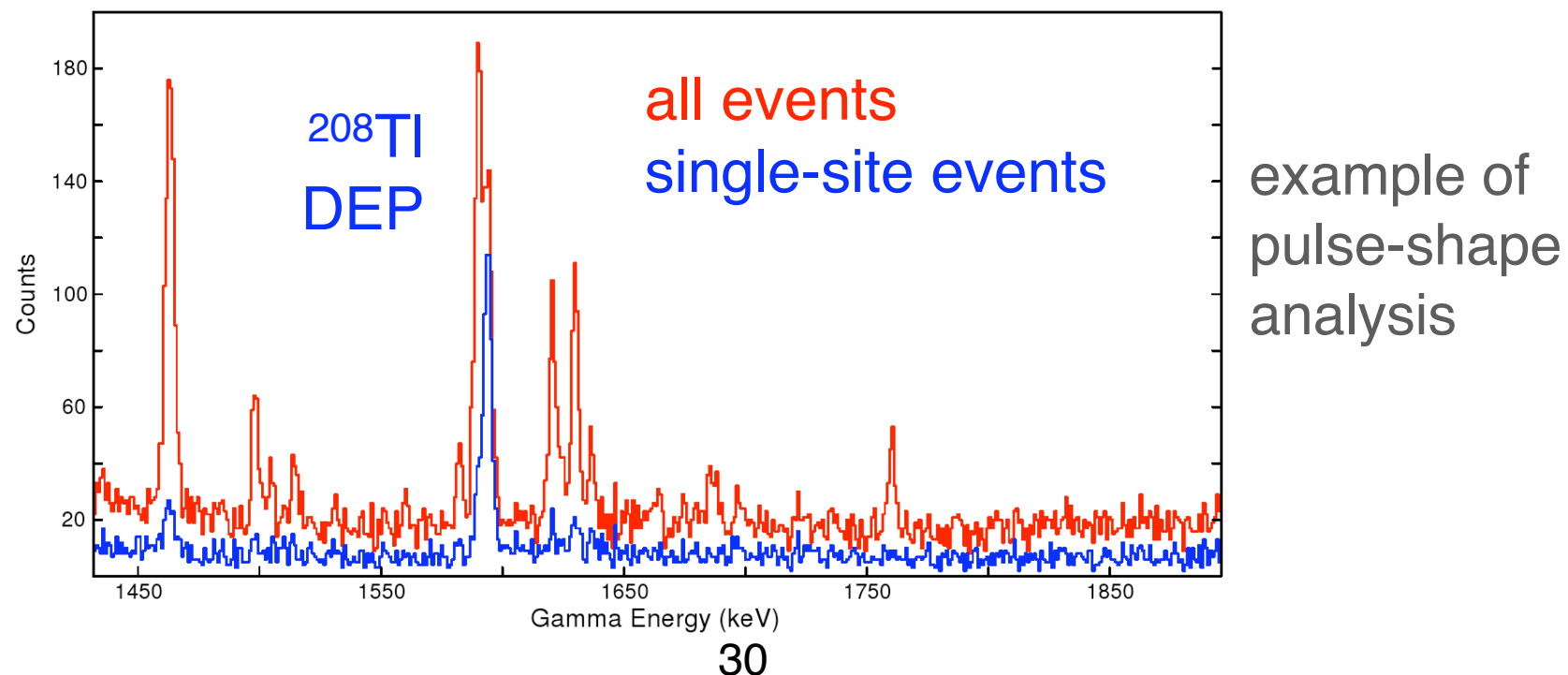


# Background mitigation techniques

- Time correlation: identify  $^{68}\text{Ge}$  to  $^{68}\text{Ga}$  decays



- Granularity: tag events that deposit energy in multiple detectors
- Pulse-shape discrimination: discriminate multi-site backgrounds

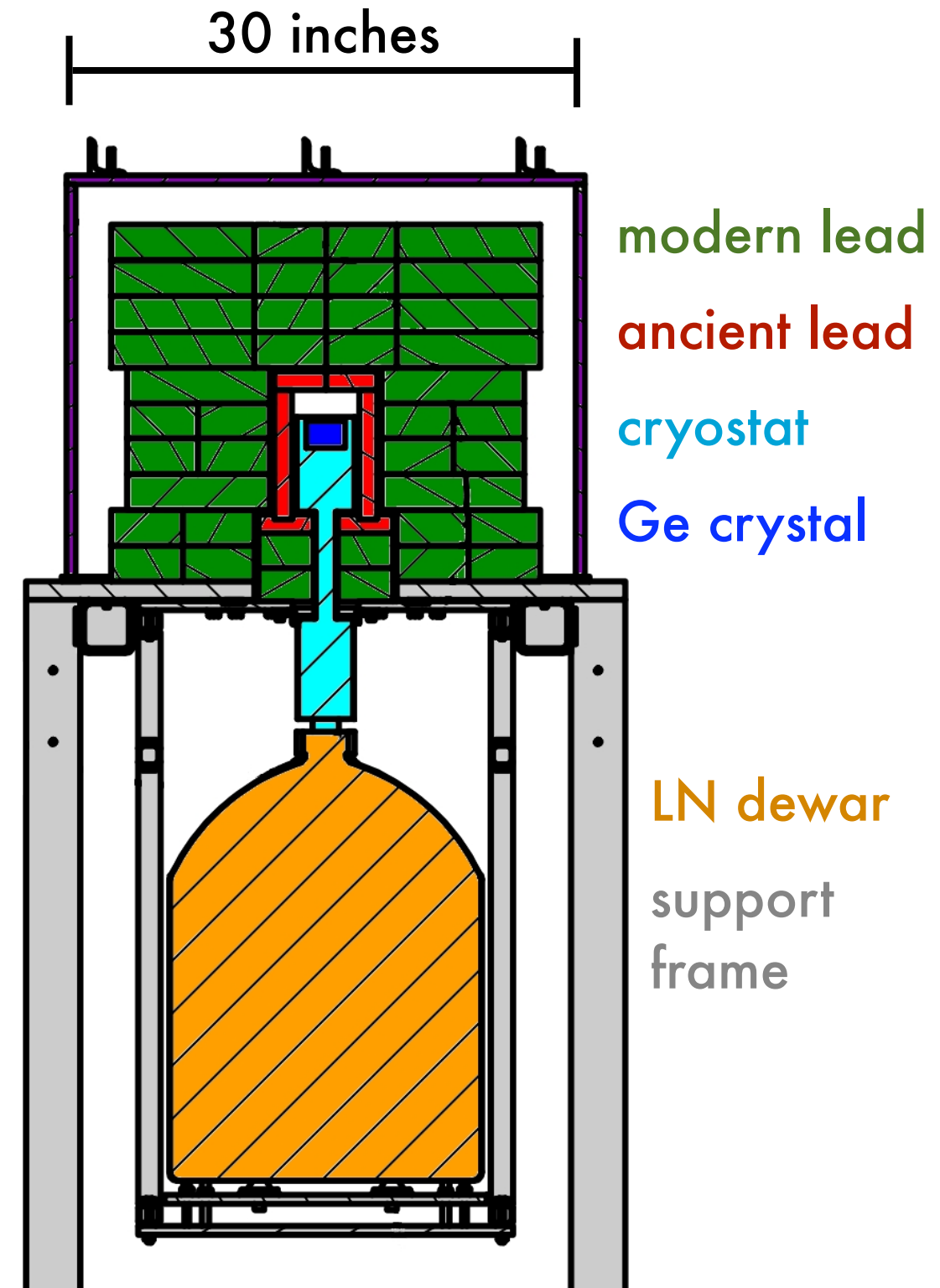




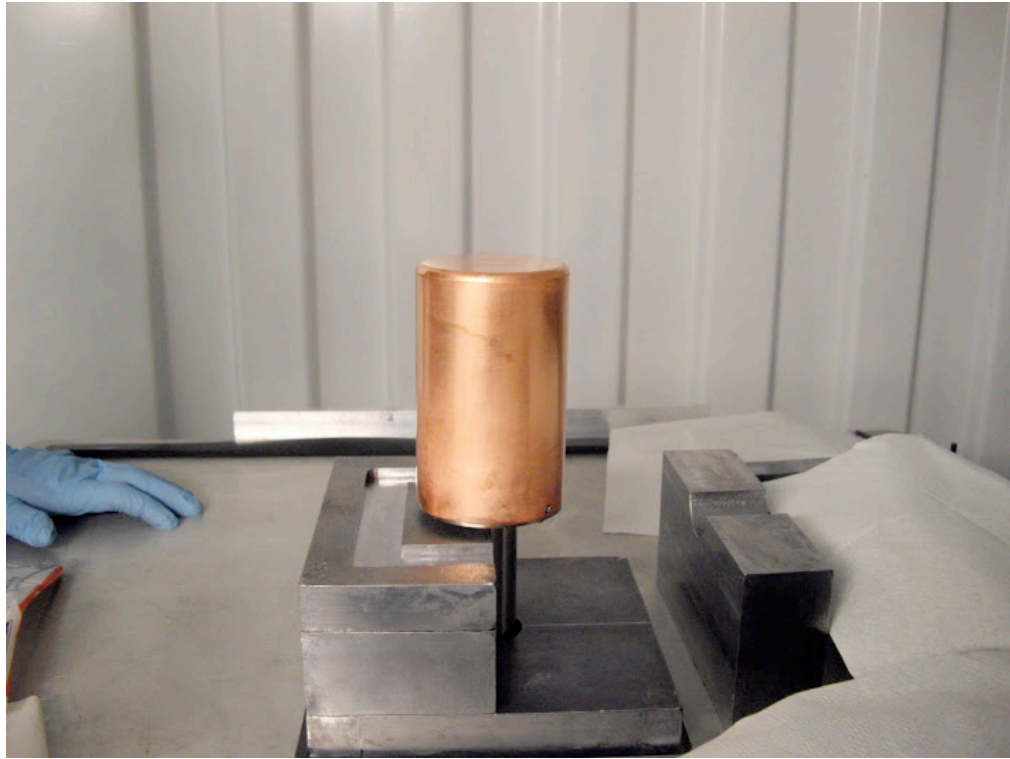
# R&D detector: MALBEK

## MAJORANA Low-background BEGe detector at Kimballton

- 0.4 kg natural germanium
- Customized CANBERRA Broad-Energy Germanium (BEGe) detector
  - Modified geometry of crystal ditch, surrounding components to minimize capacitance
  - Low-background copper cryostat
- Kimballton Underground Research Facility in Ripplemeade, VA (1700', 1400 m.w.e.)
- R&D:
  - Dark matter search
  - MAJORANA-like data-acquisition system
  - Measurement & model of background energy spectrum



# R&D detector: MALBEK





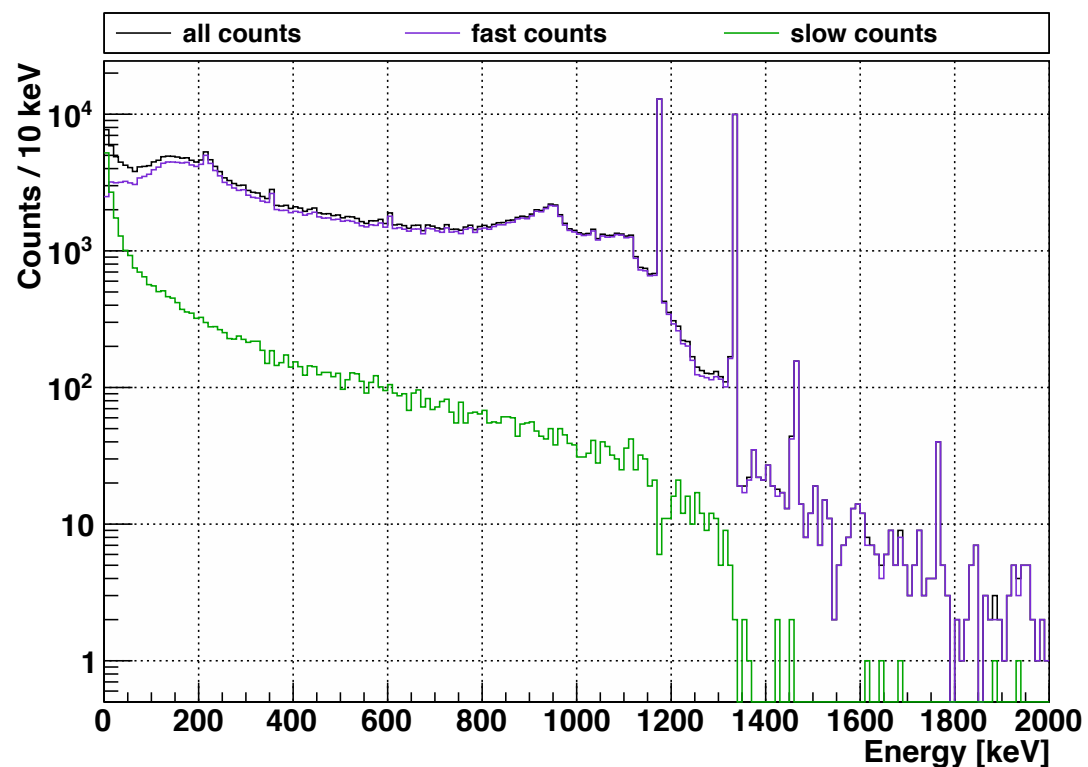
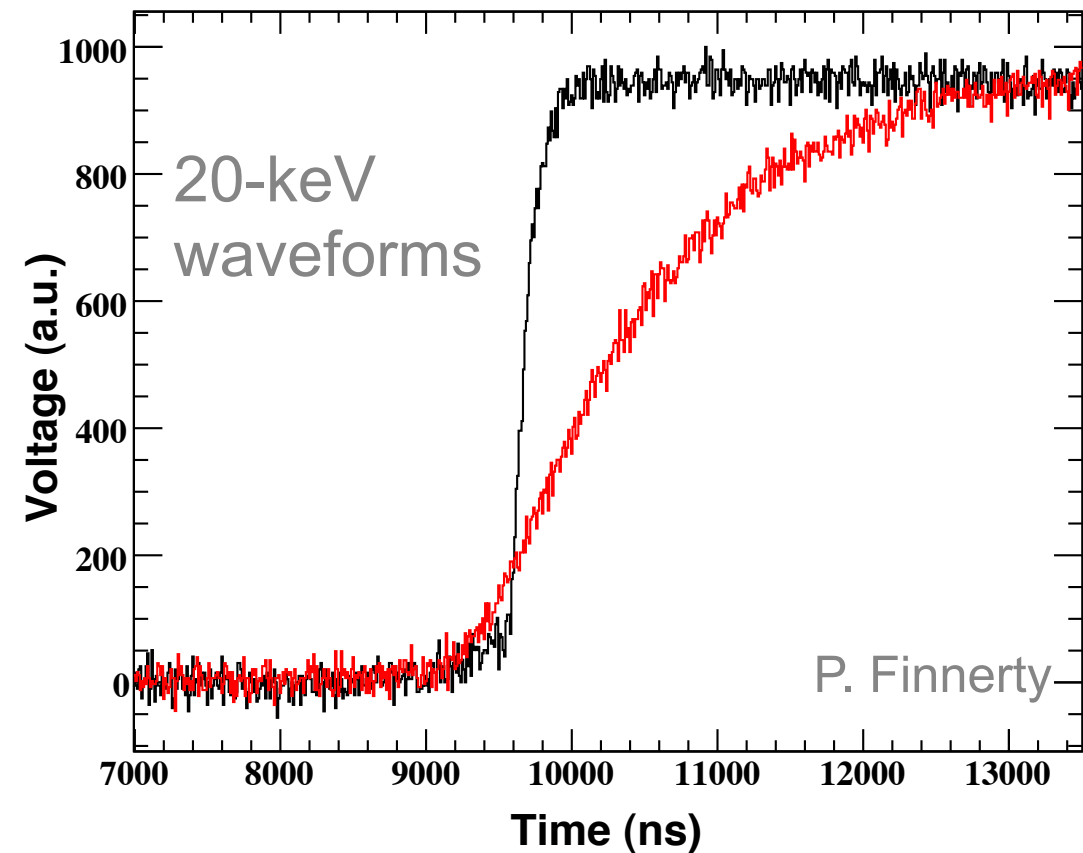
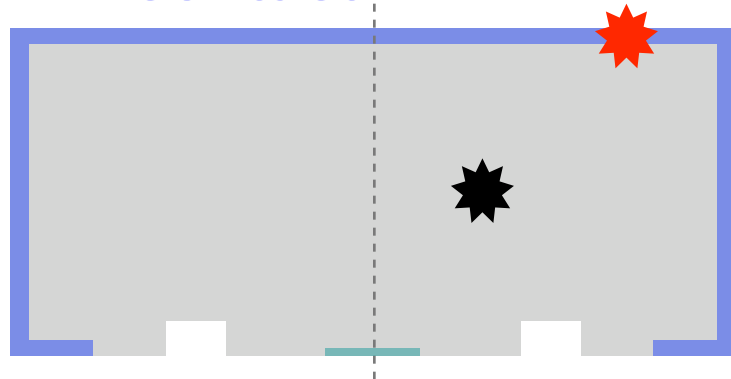
# Kimballton Underground Research Facility





# slow-rising, energy-degraded pulses

outer n+ contact



measured  $^{60}\text{Co}$  spectrum

slow pulses contribute throughout the energy spectrum

# MALBEK background model

## Material purity information

**cosmogenic contaminants:** production rates from literature

**primordial contaminants:** material assay data from literature and direct measurements of bulk and surface contamination

**cosmic muon flux:** from literature

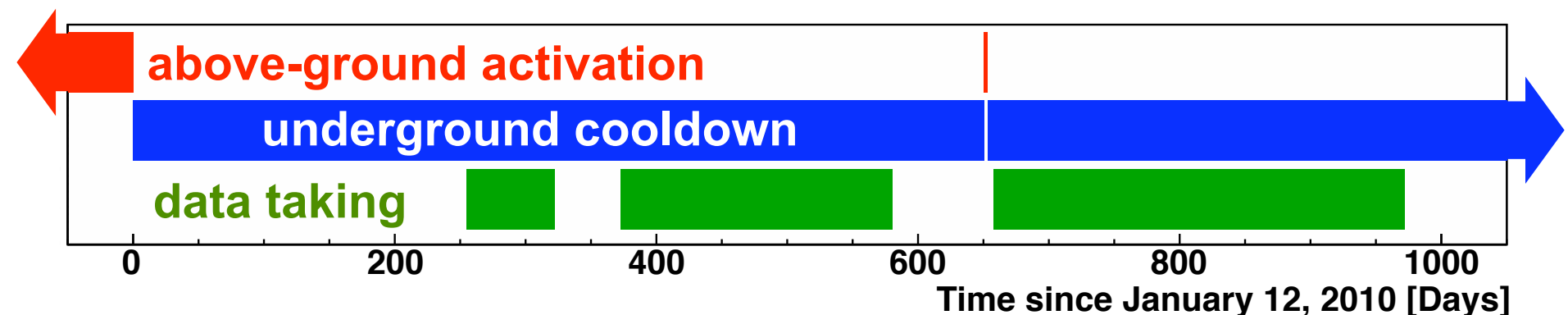
## Monte Carlo simulation results

**Geant4 simulations** to determine efficiencies for contamination to deposit energy in our detectors

50k CPU hours

8k+ runs, 40+ contaminants, 56 components, 21 materials

## Exposure history



## Detector characteristics

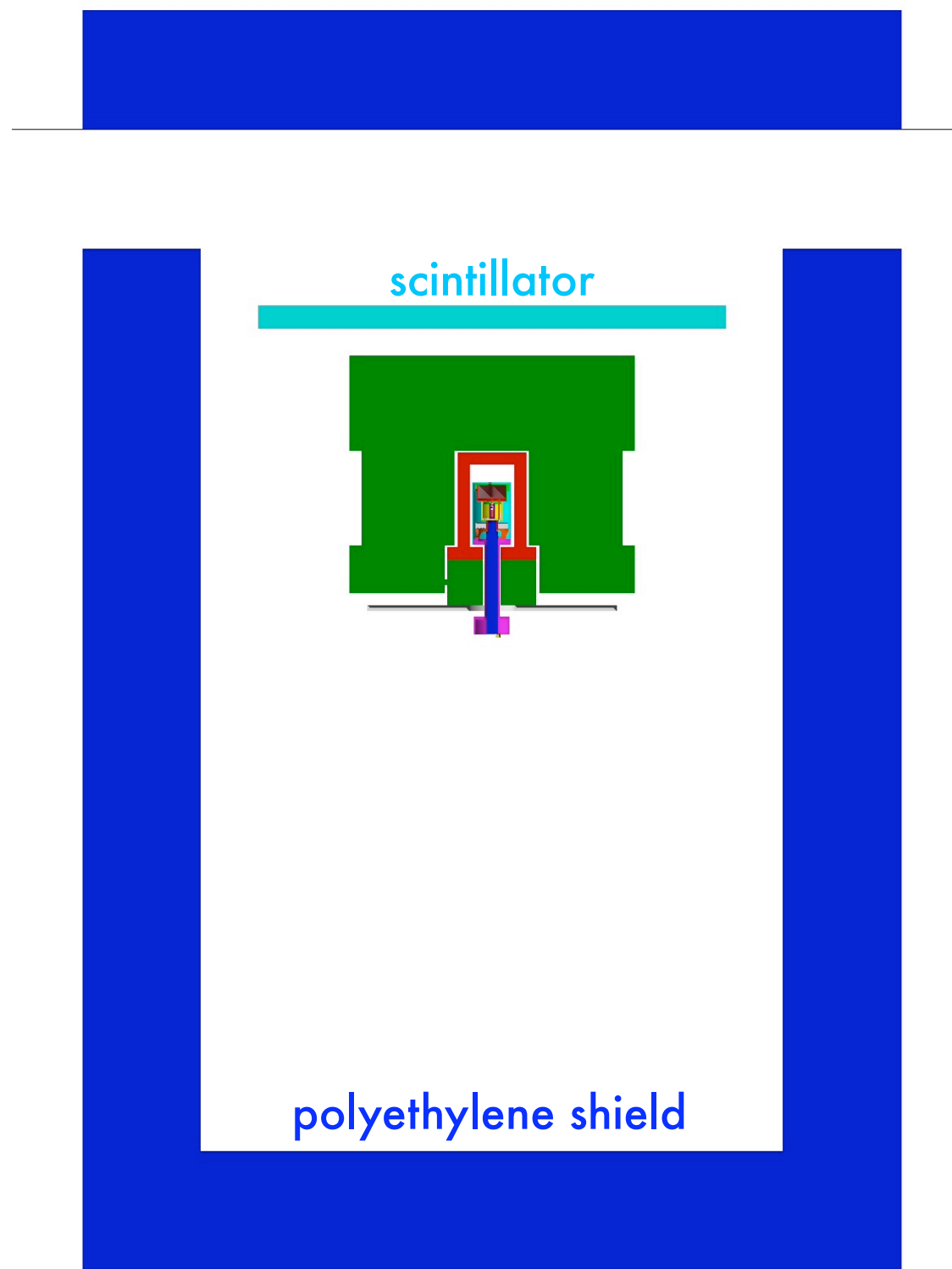
**energy resolution:**  $\sigma(E) = (0.12^2 + 0.09\epsilon E + E^2)^{1/2}$  keV

**dead layer properties:**  $0.93 \pm 0.09$  mm outer  $n^+$  contact [arXiv:1207.6716]

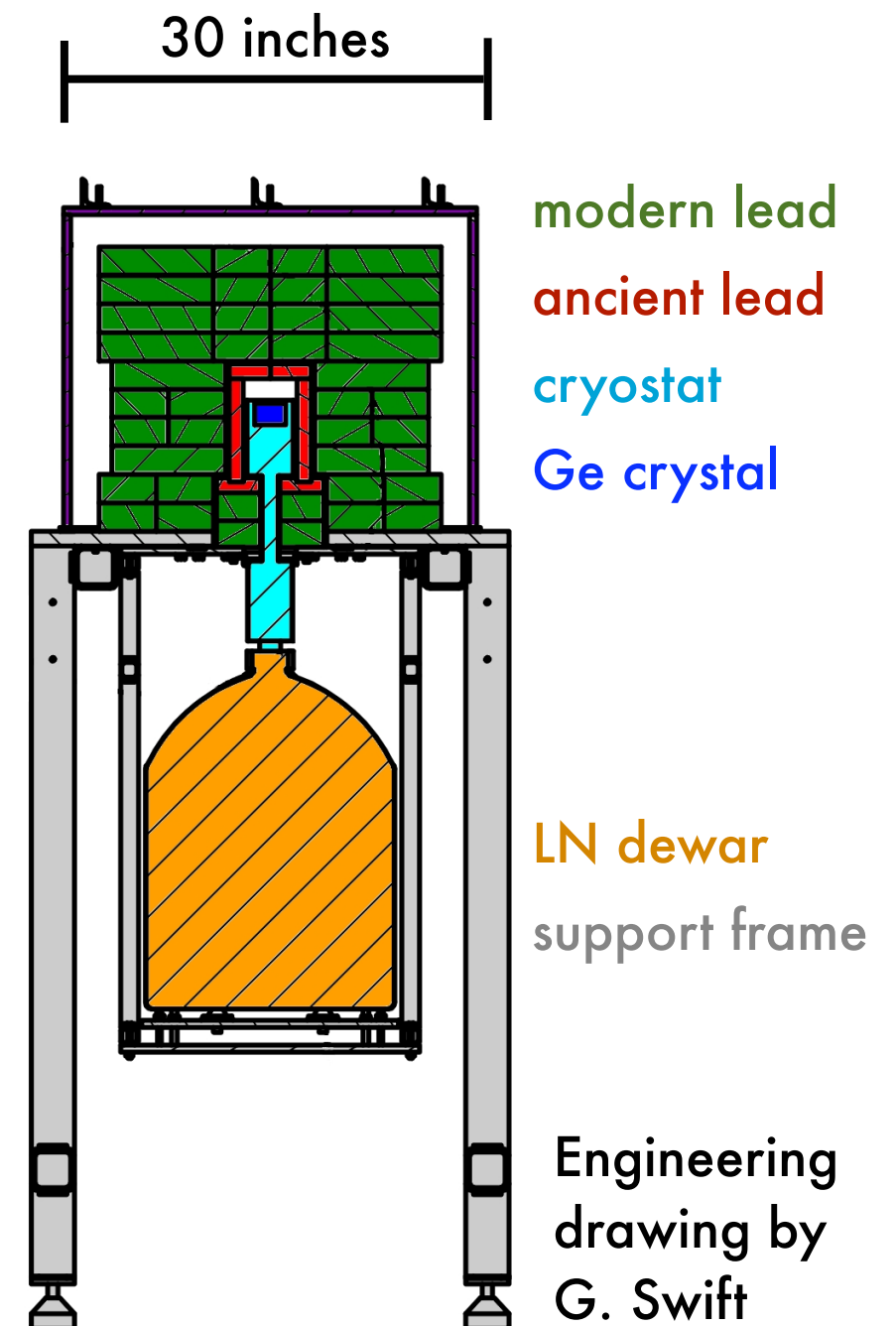
**preamplifier effects:** efficiency as a function of energy



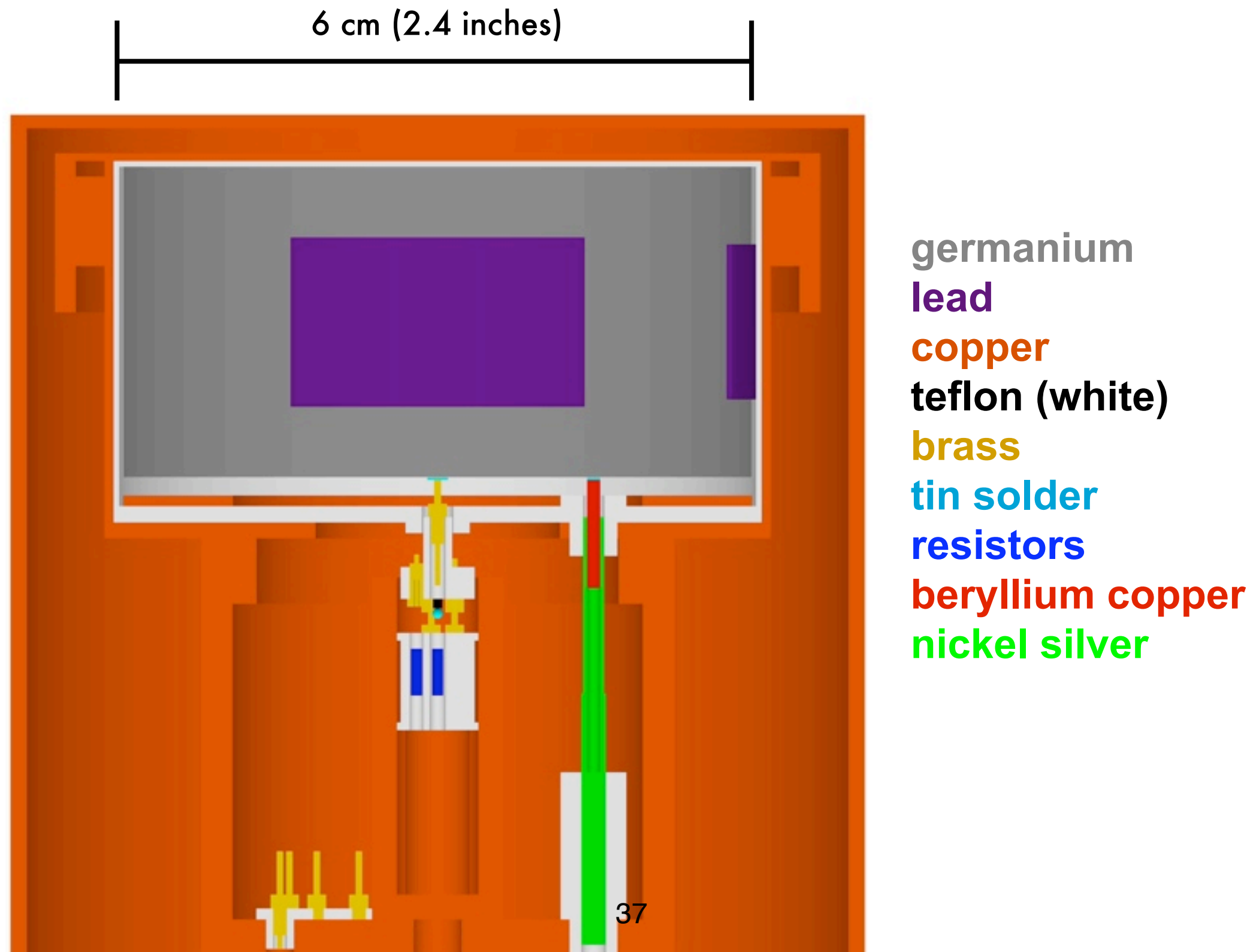
# Geant4 geometry model



Geant4 geometry

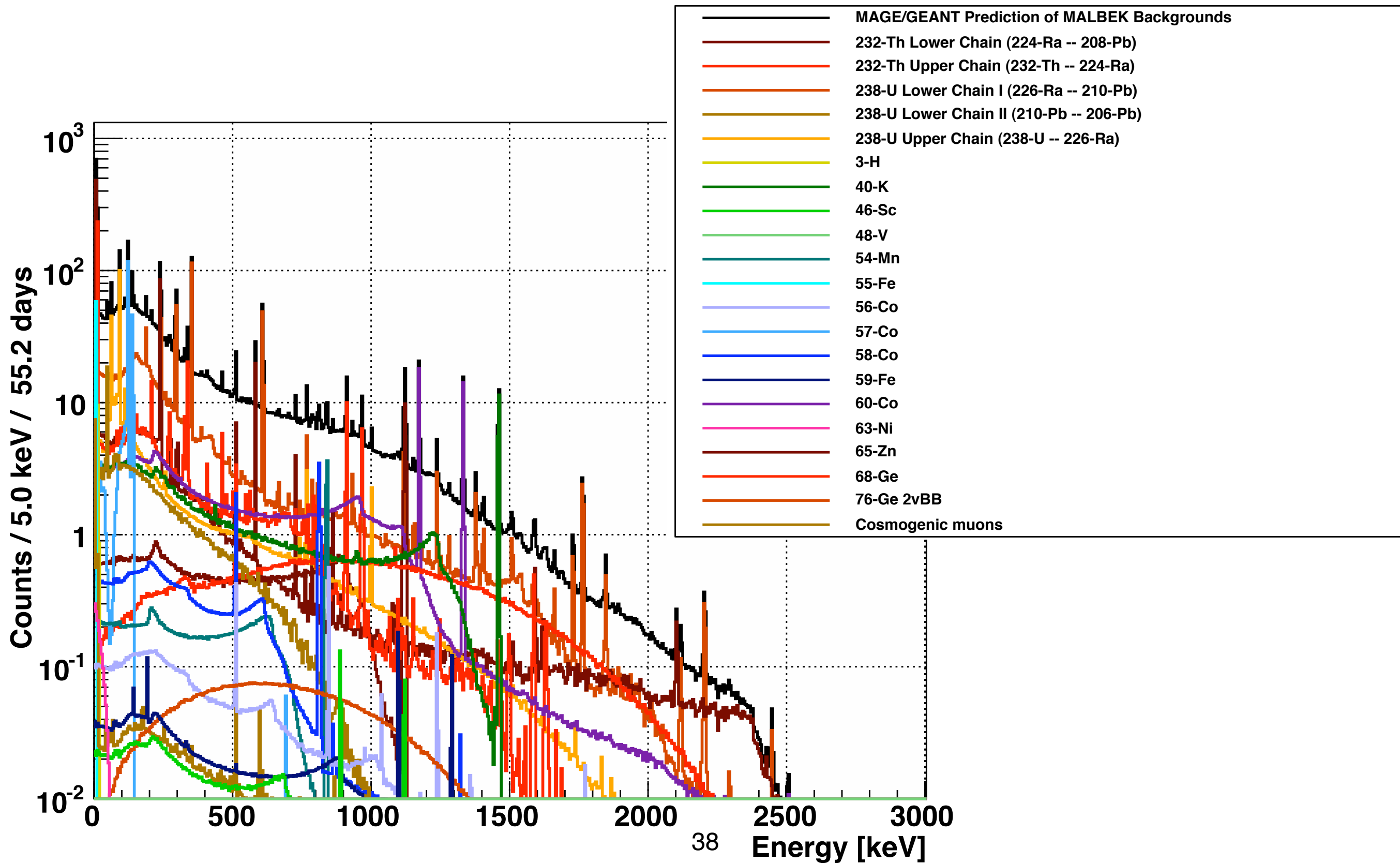


# Geant4 geometry model



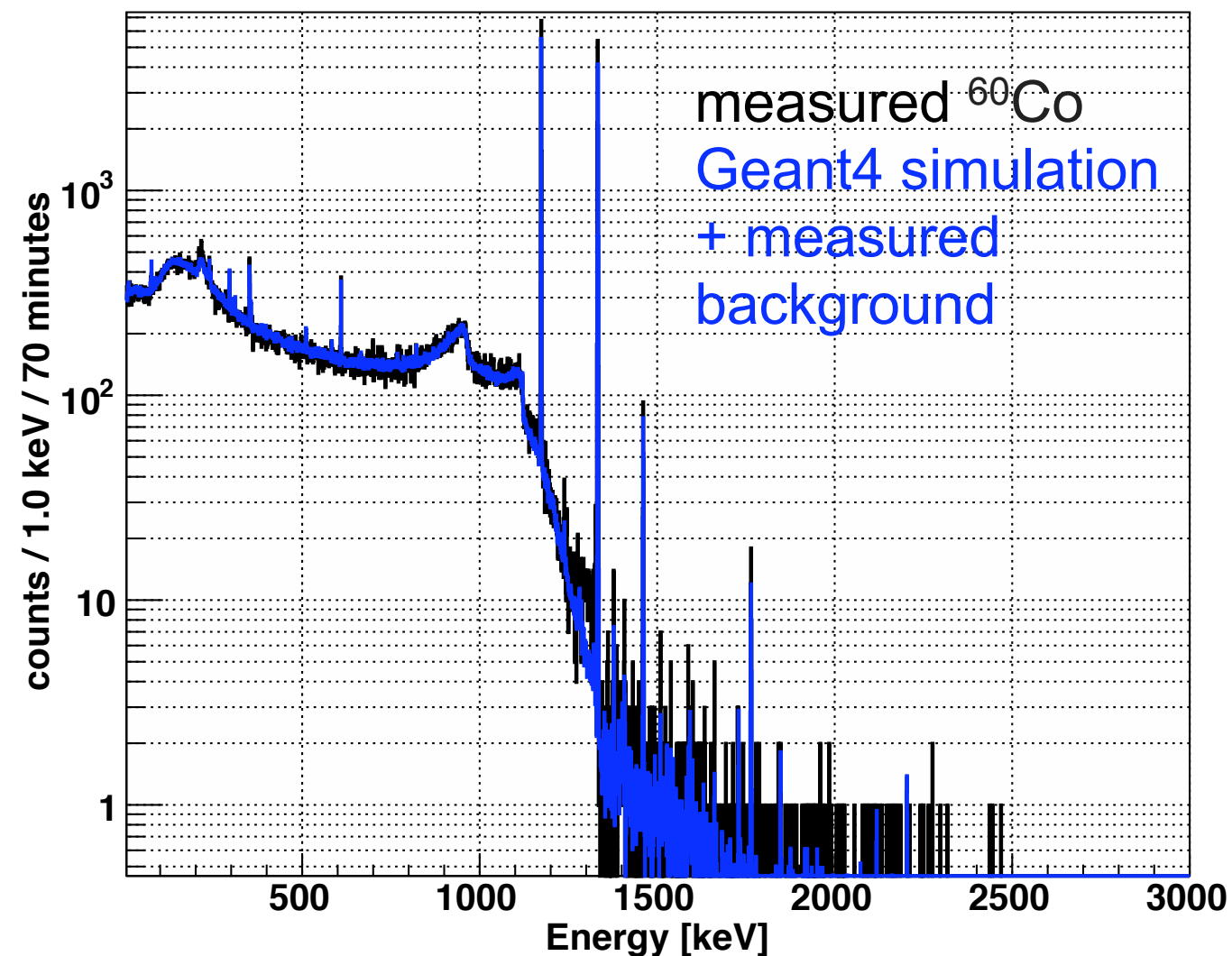


# MALBEK background model

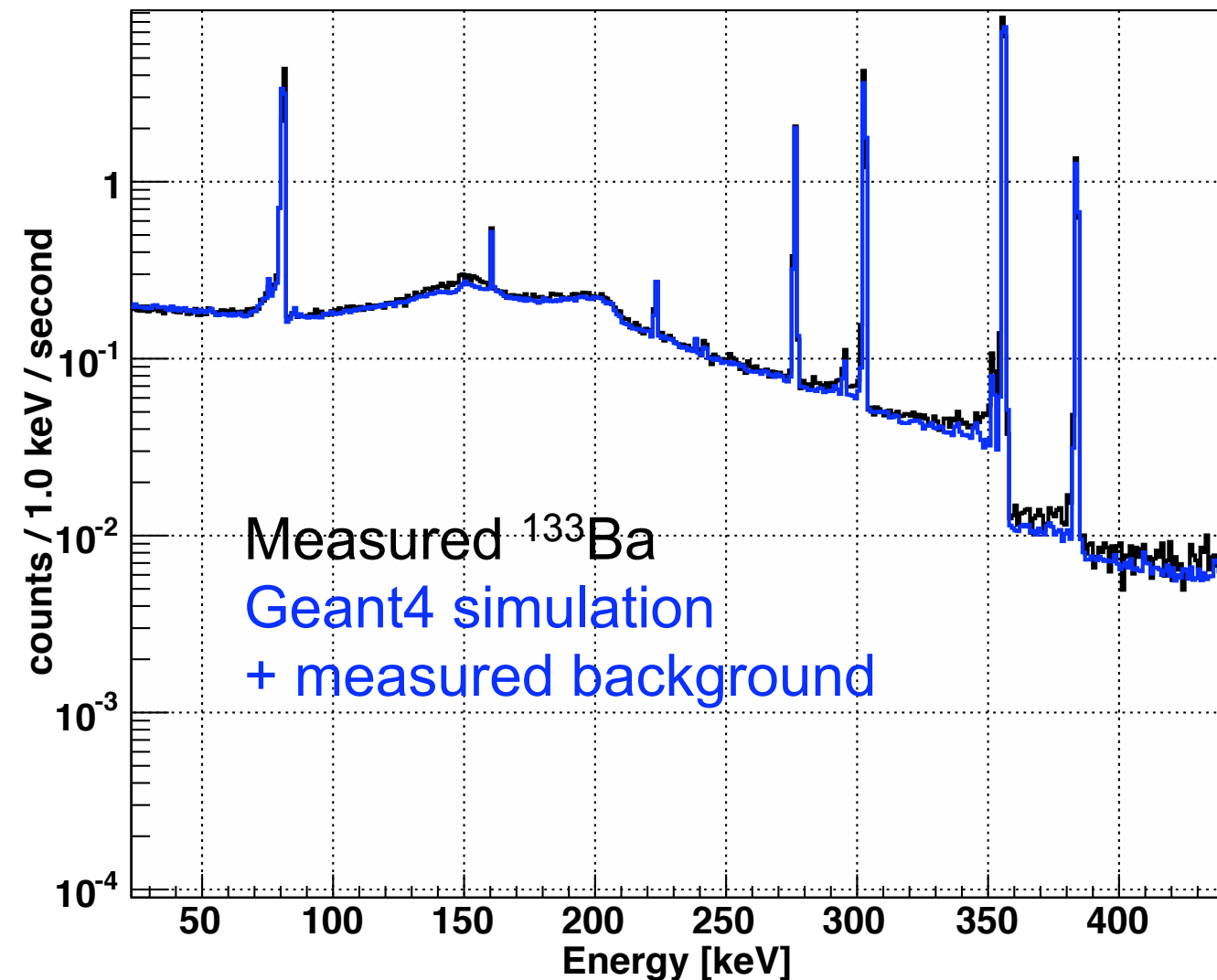


# validation tests

$^{60}\text{Co}$ : integral count rate agrees within 2% between 5 and 3000 keV

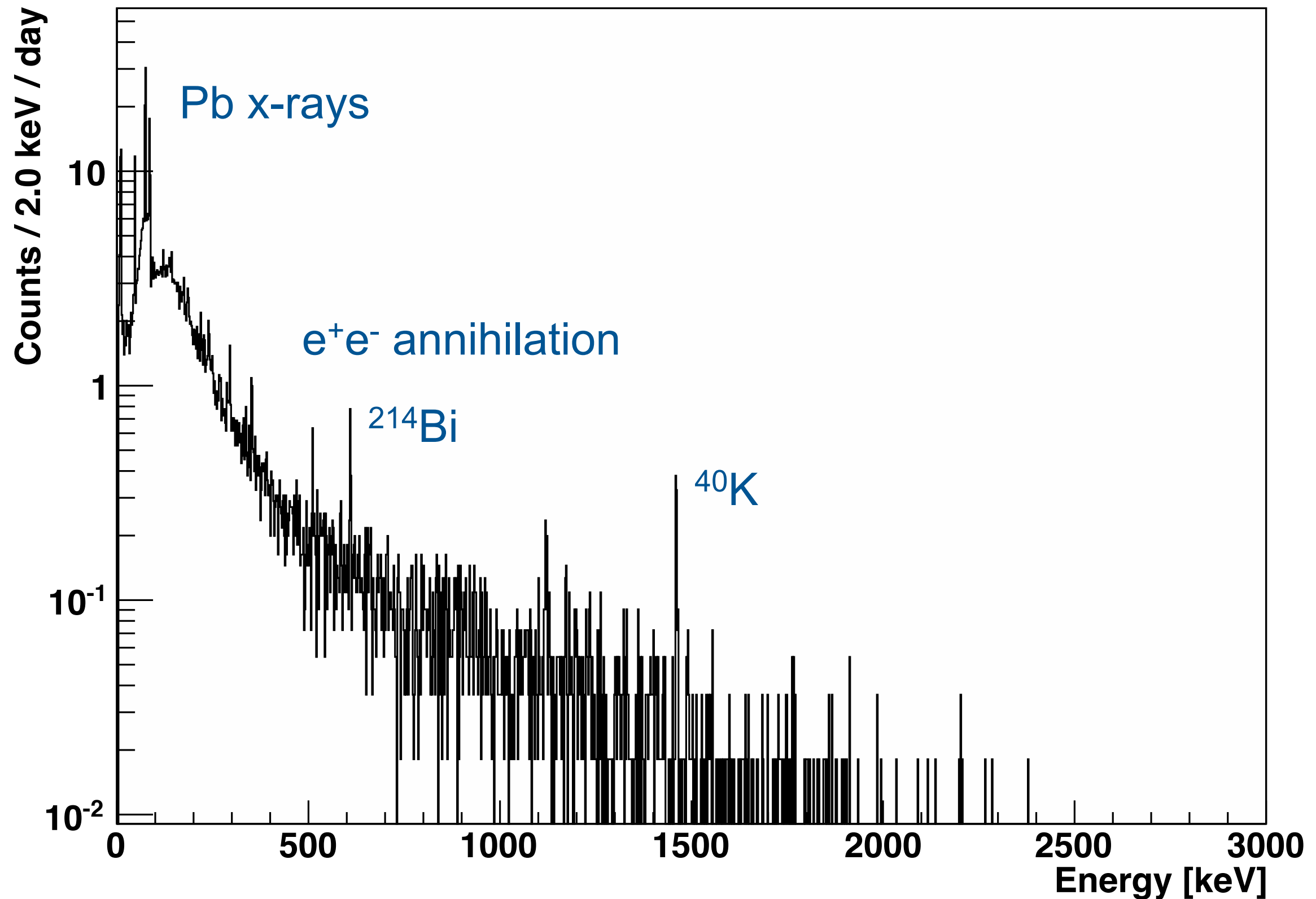


$^{133}\text{Ba}$ : integral count rate agrees within 3% between 5 and 400 keV

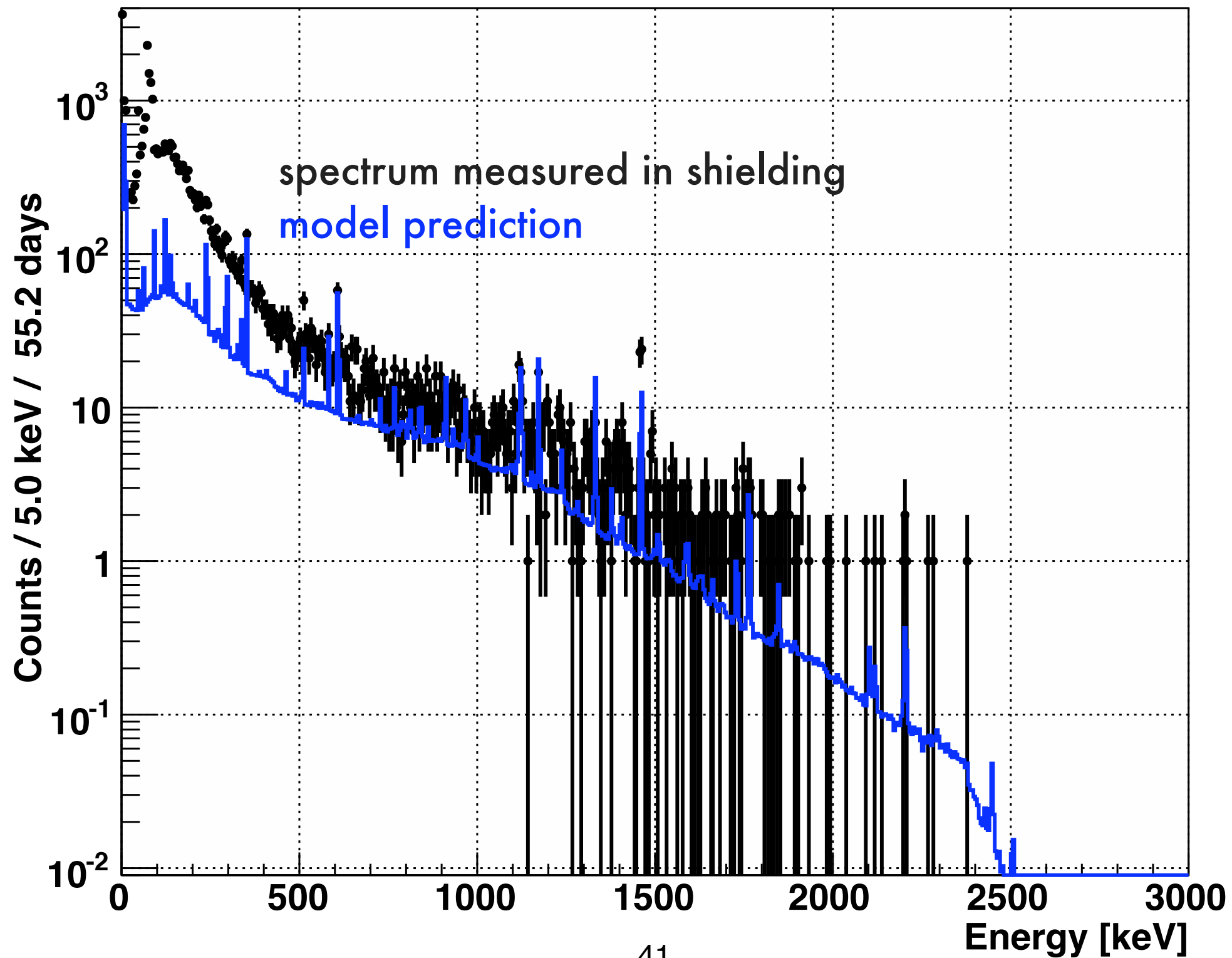




# Shielded background energy spectrum measured at Kimballton



# Measurement exceeded our expectations



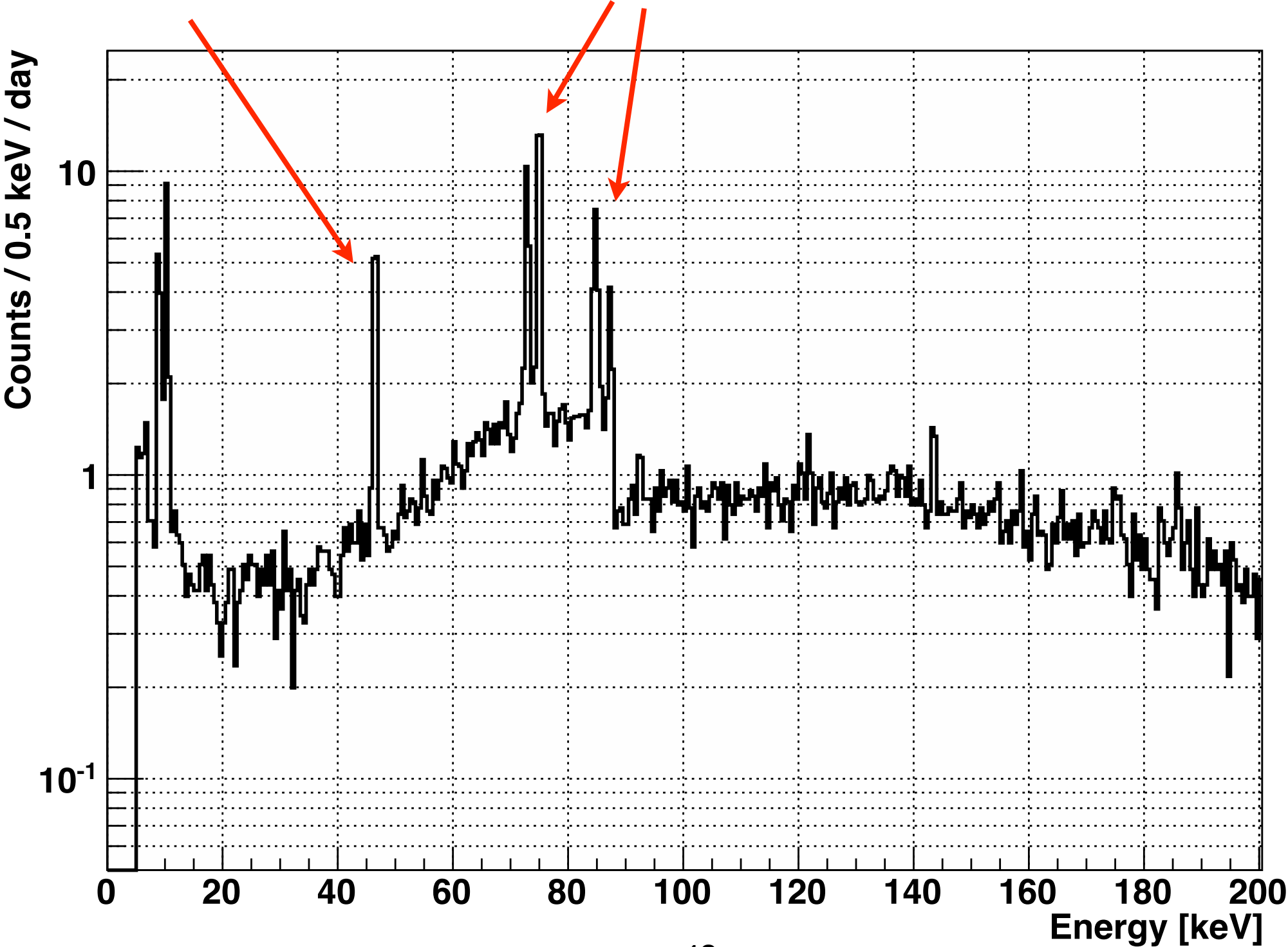




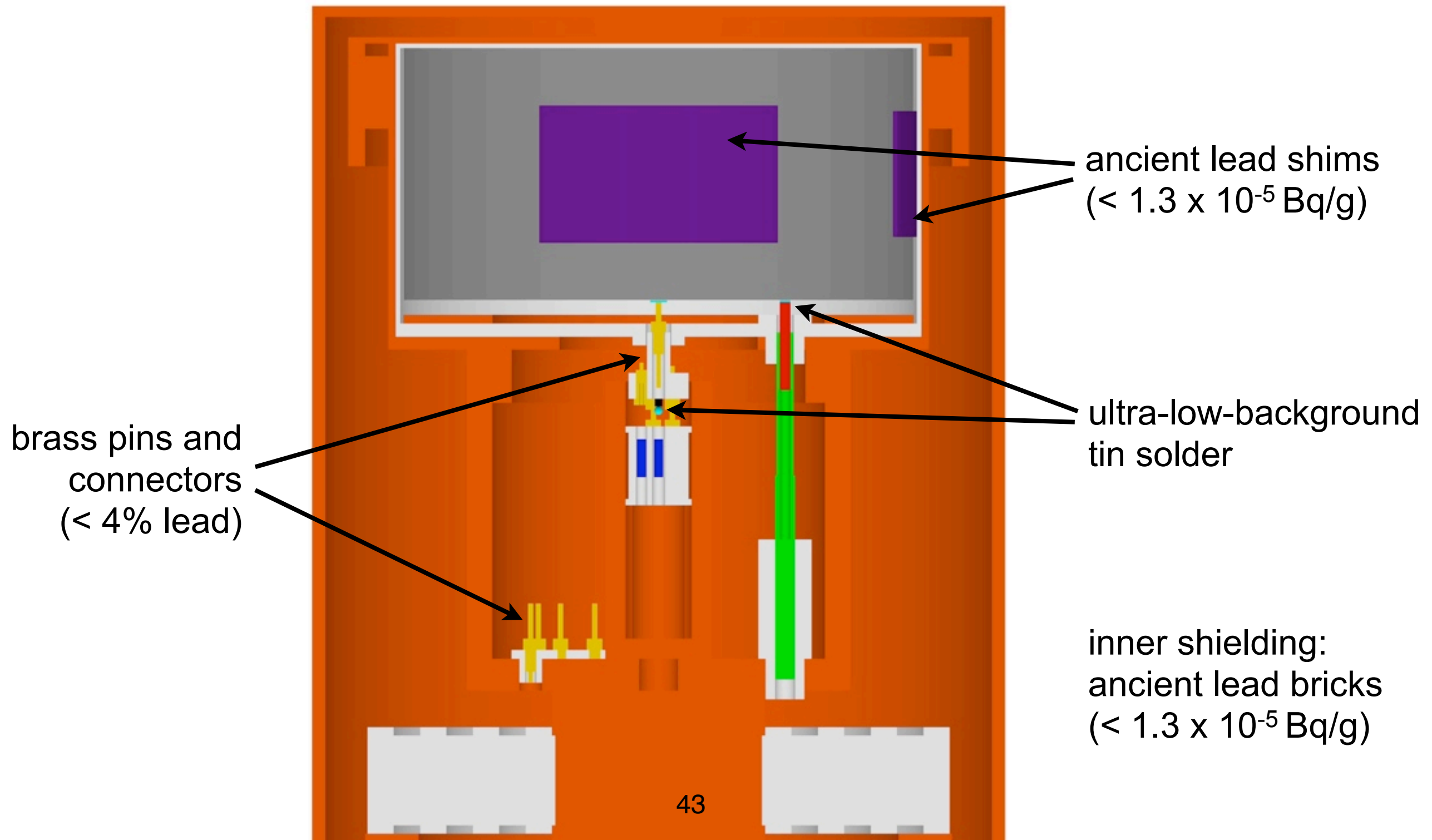
$T_{1/2} = 22 \text{ years}$   
 $\gamma \text{ } 46.5 \text{ keV}$

$T_{1/2} = 5 \text{ days}$   
 $\beta \text{ } 1162 \text{ keV end point}$

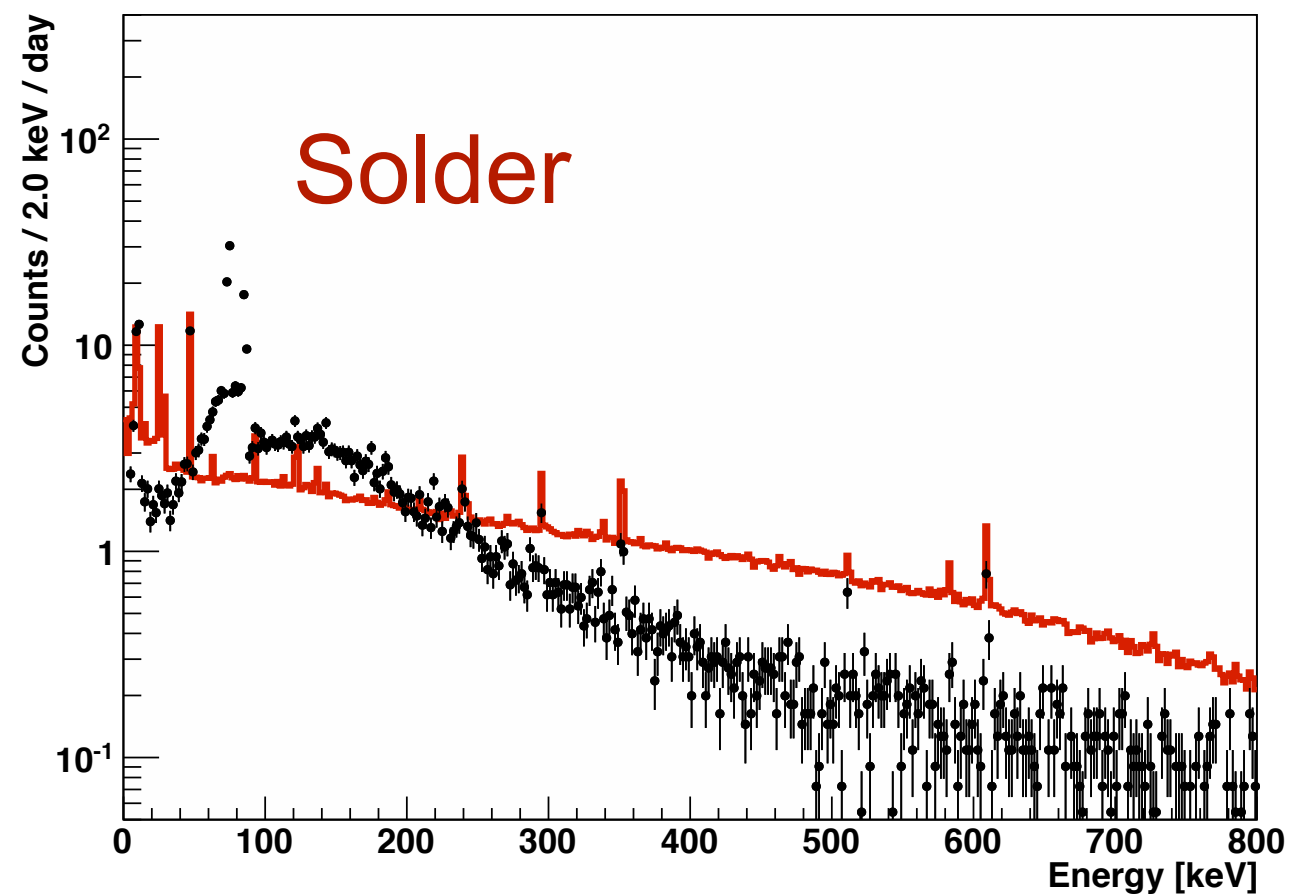
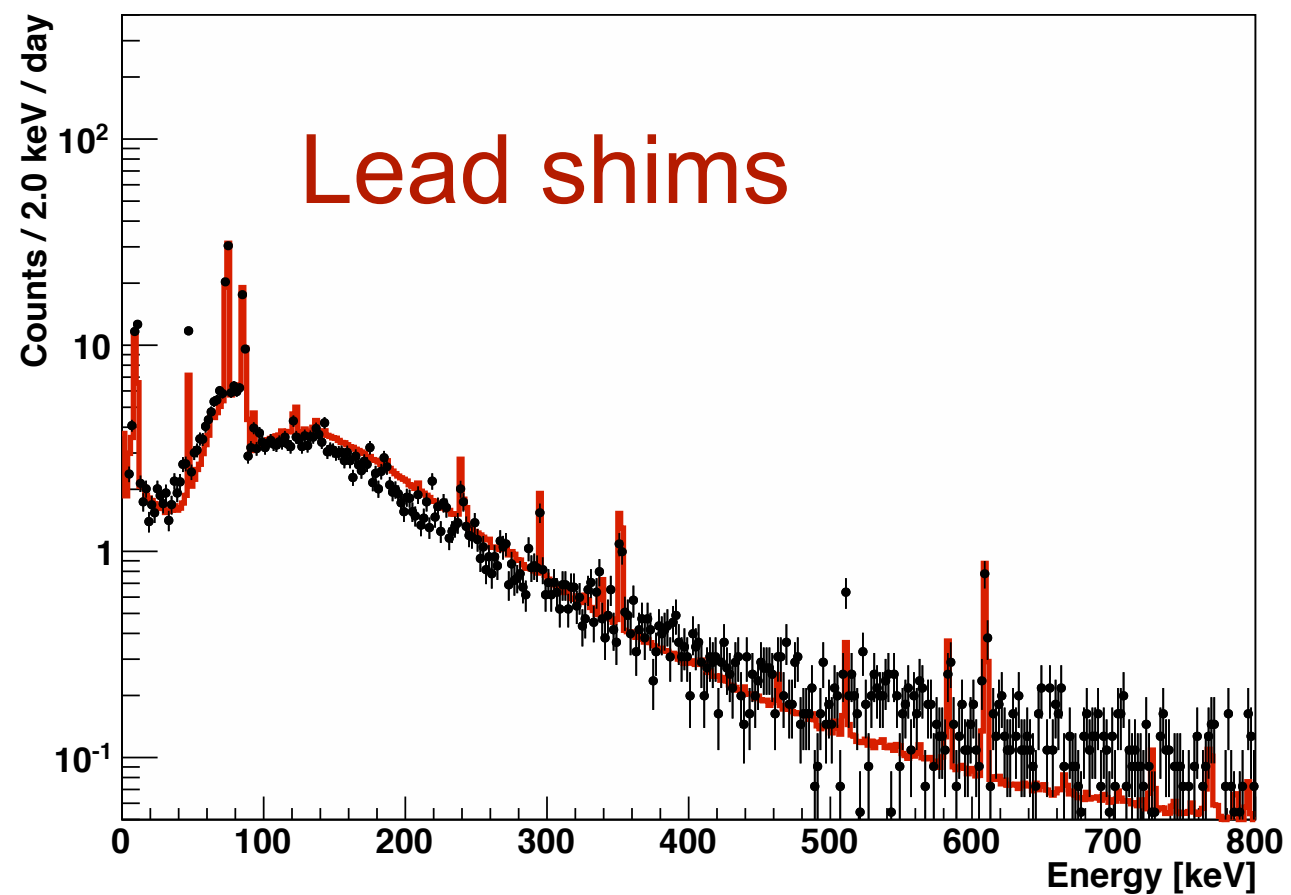
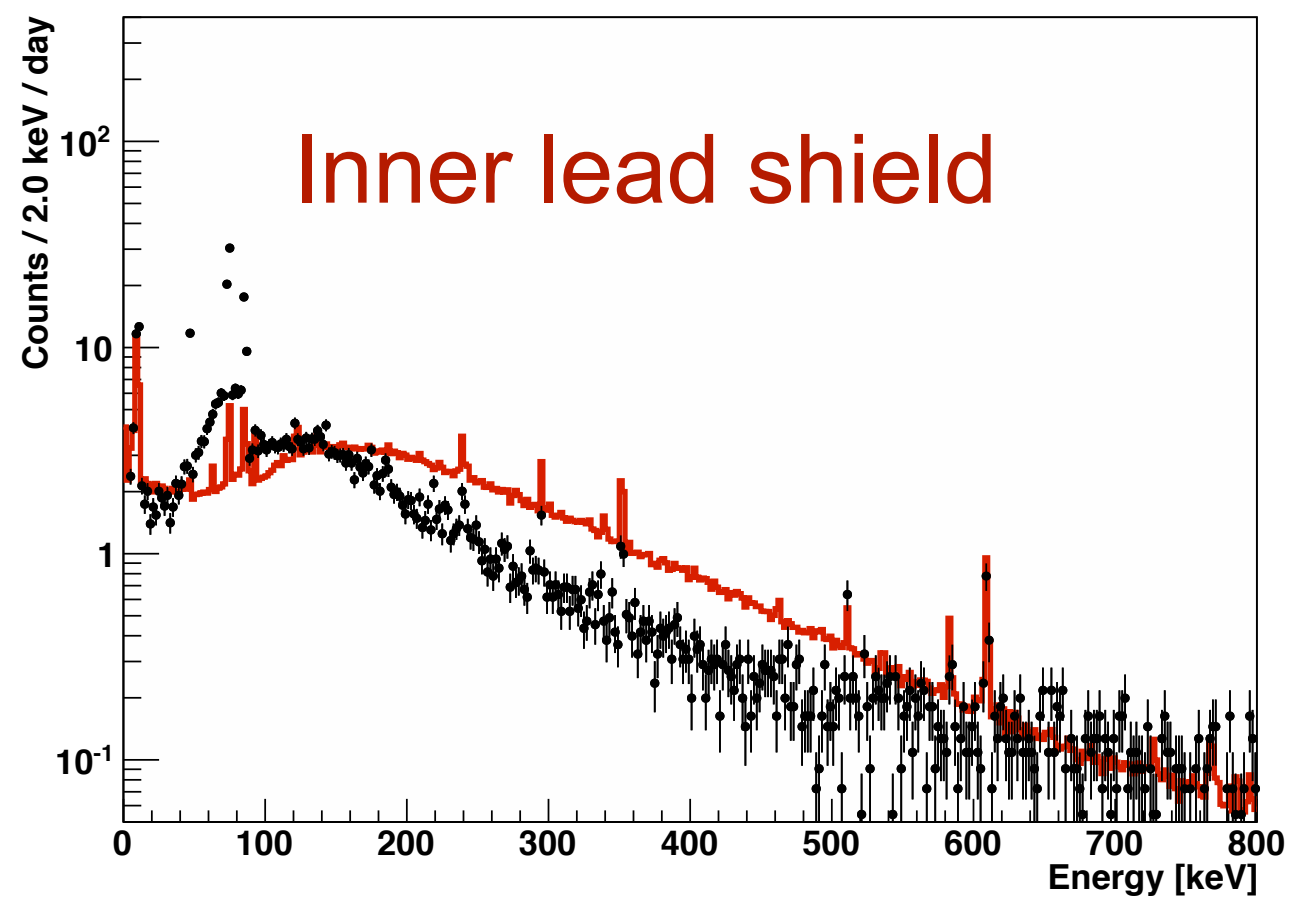
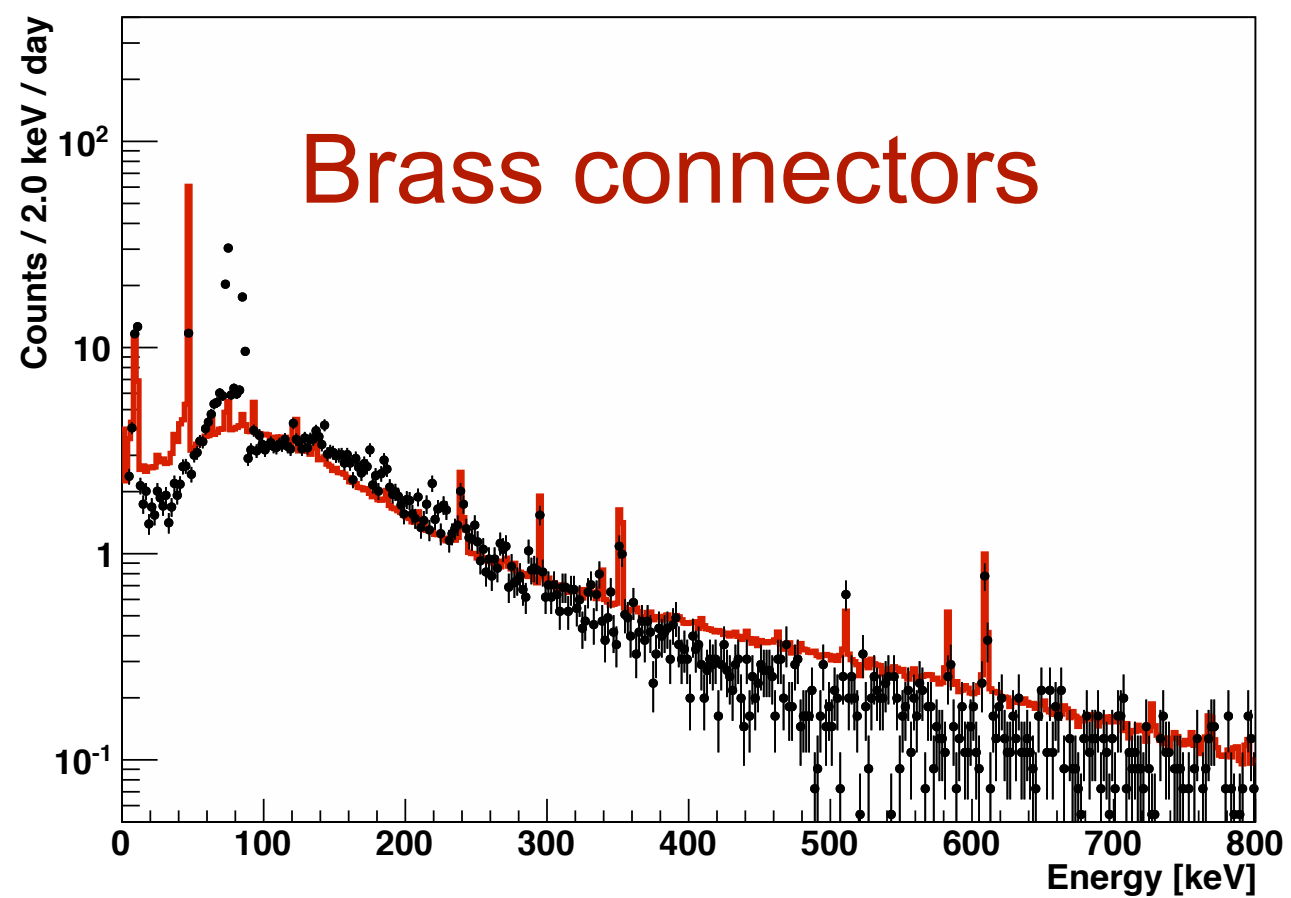
$T_{1/2} = 138 \text{ days}$   
 $\alpha \text{ } 5.3 \text{ MeV}$



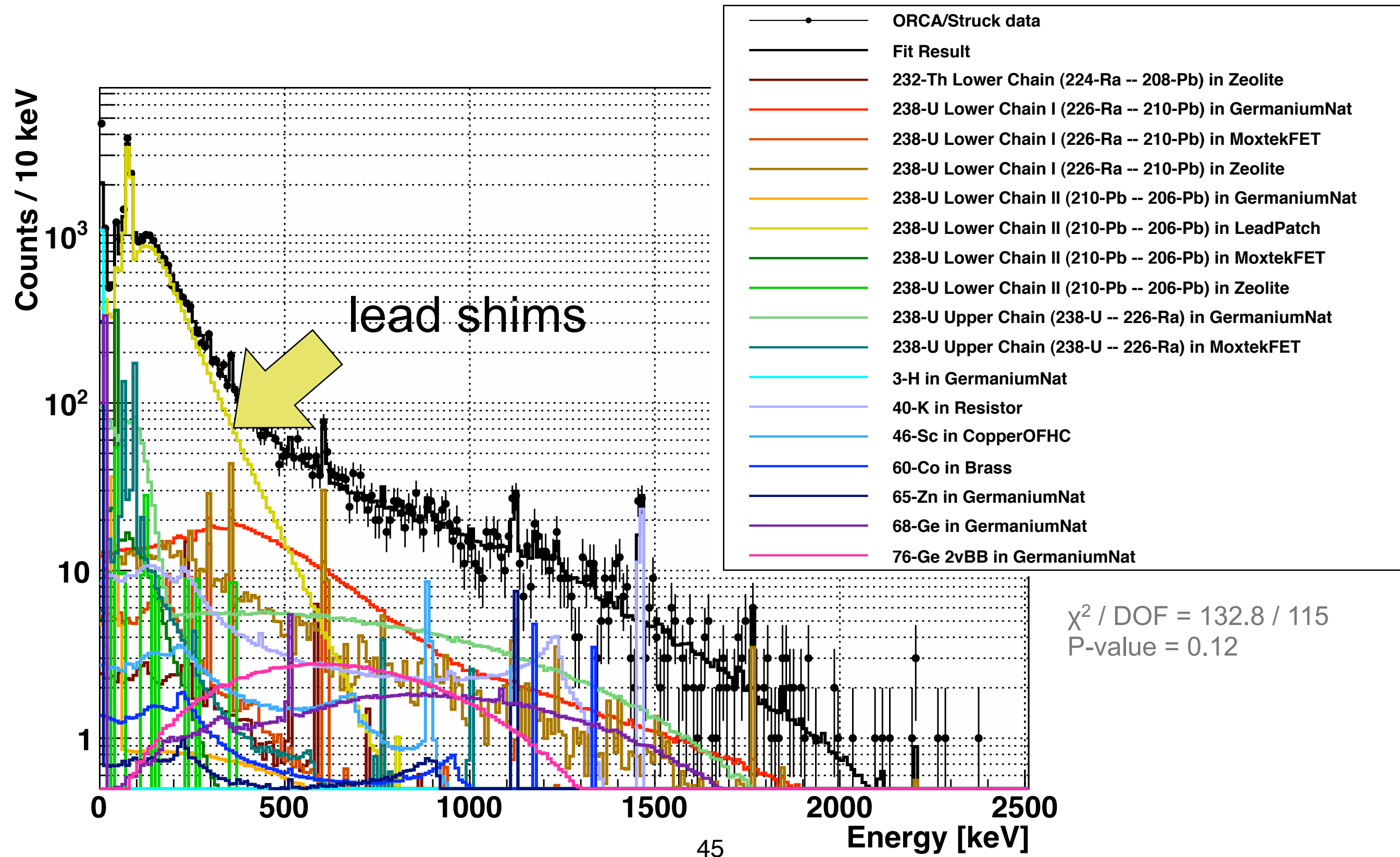
# Possible sources of $^{210}\text{Pb}$ contamination







# Background model fit of energy spectrum

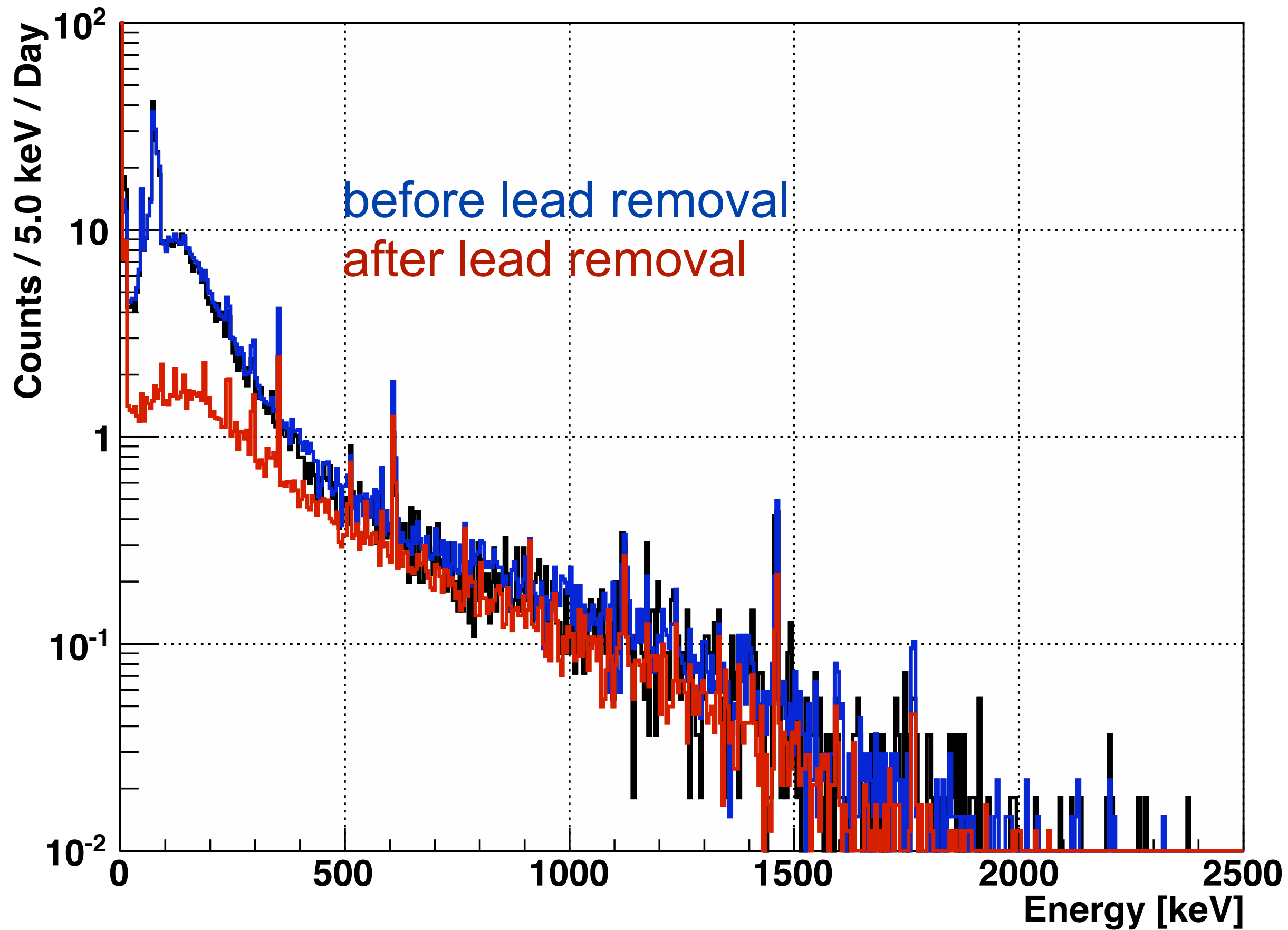




# Lead shim removal

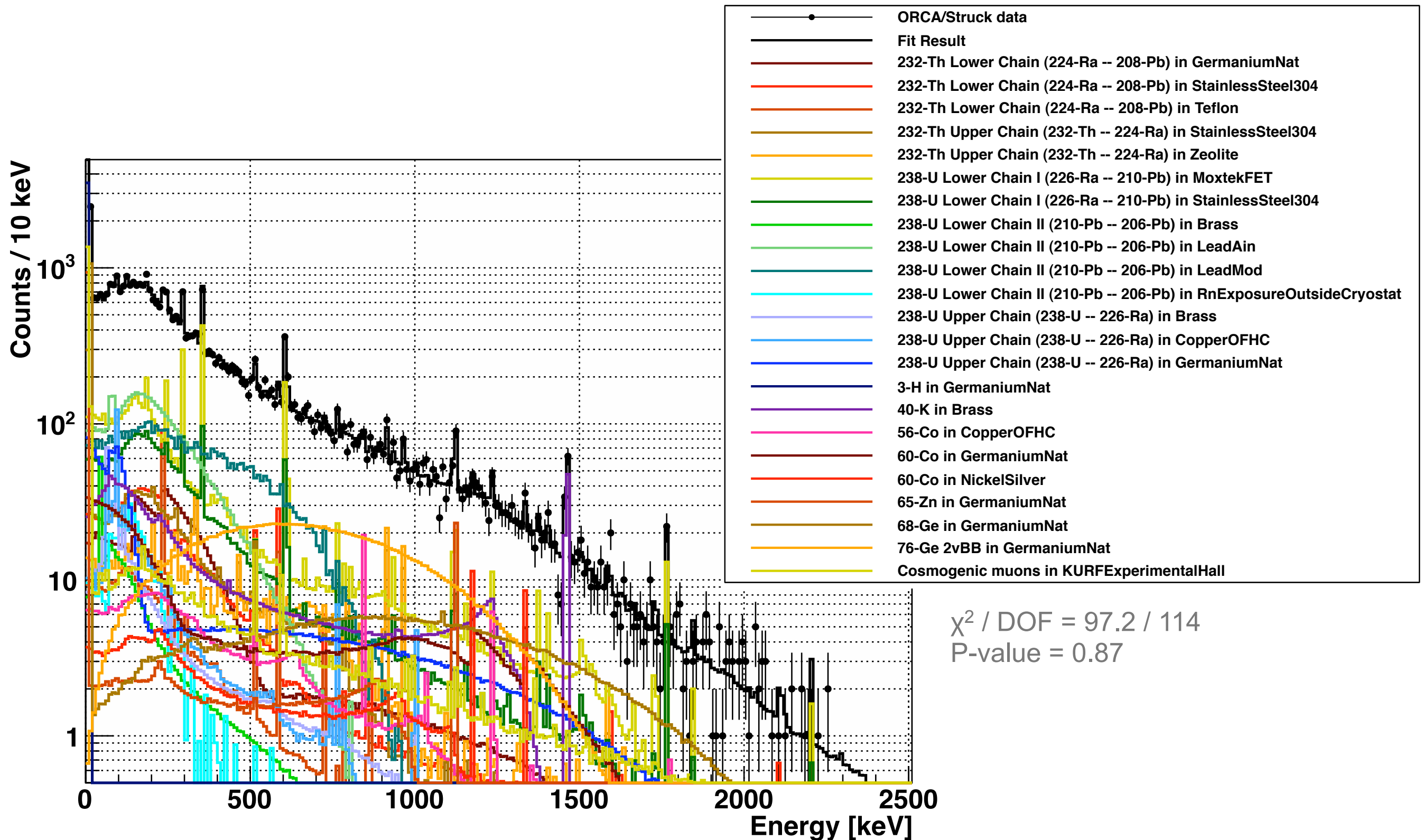


Removed 3g of lead from cryostat





# Background model fit of spectrum after shim removal

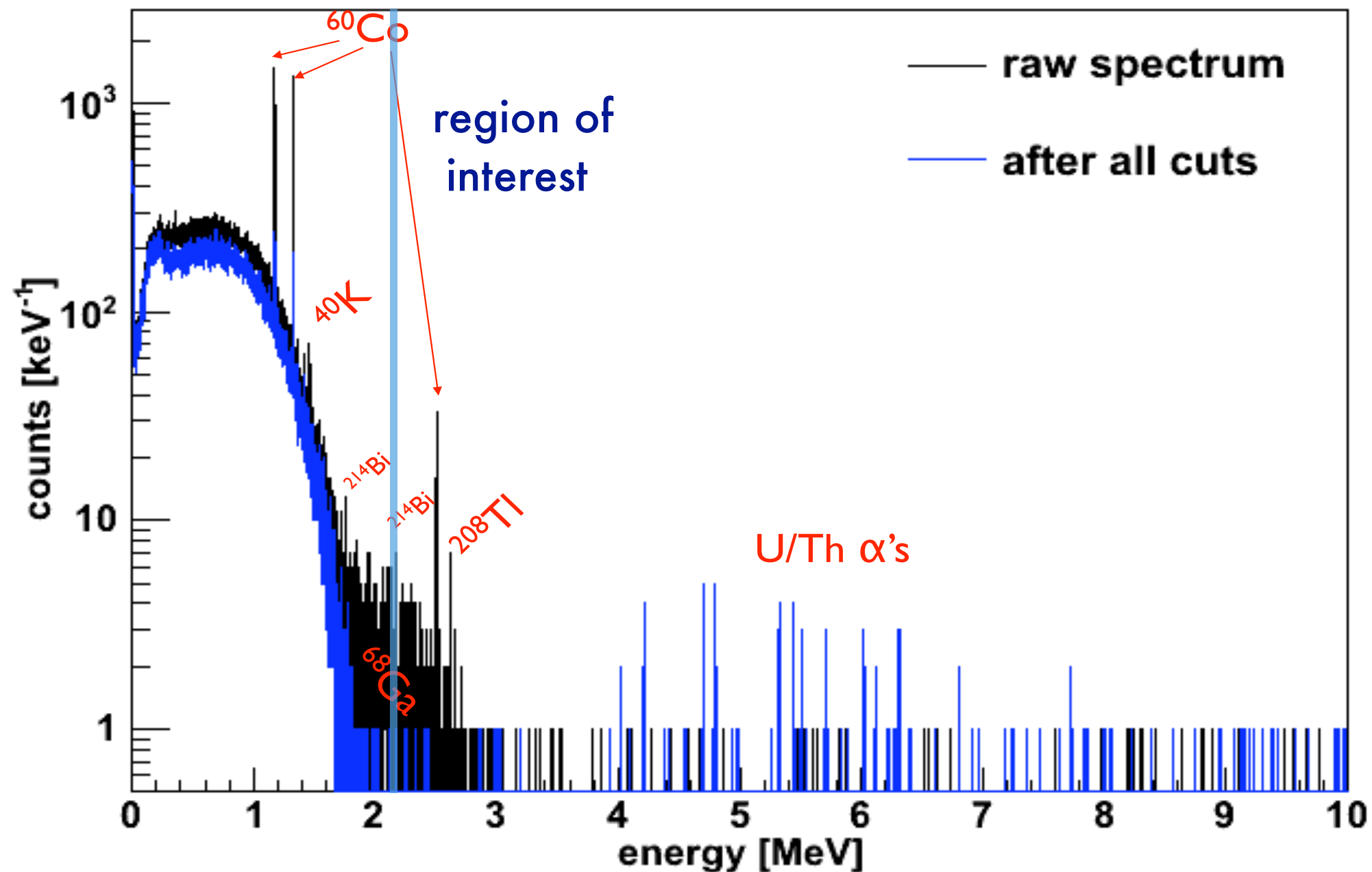


# Results from R&D detector

- Validated background model of the energy spectrum
- Identified and removed contaminated component from cryostat
- Studied slow pulses and developed a cut to remove them throughout the energy spectrum
- Tested MAJORANA data-acquisition system
- Developed and tested software for simulations and analysis of data

# DEMONSTRATOR background model

Simulated spectra, 40 kg yrs, detector resolution applied





# Conclusions

- Observation of neutrinoless double beta decay would determine Majorana nature of the neutrino, demonstrate lepton number violation, and provide information about neutrino mass
- MAJORANA DEMONSTRATOR is a 40-kg detector array searching for  $0\nu\beta\beta$  of  $^{76}\text{Ge}$ 
  - Under construction at Sanford Underground Laboratory
  - On track to begin taking data in September 2013
  - Tests of data taking, data analysis, and background modeling with an R&D detector have been successful



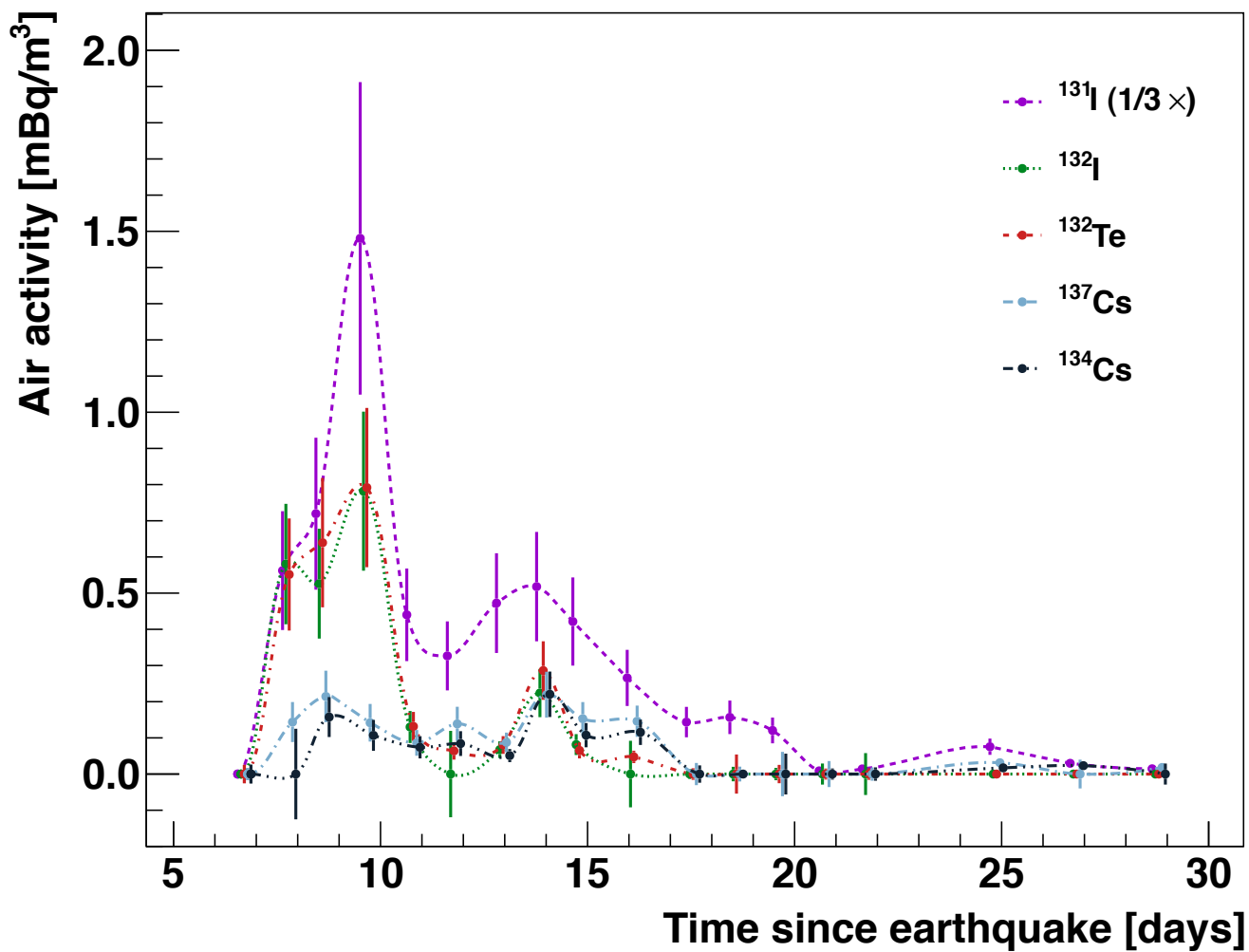
# Thank you!



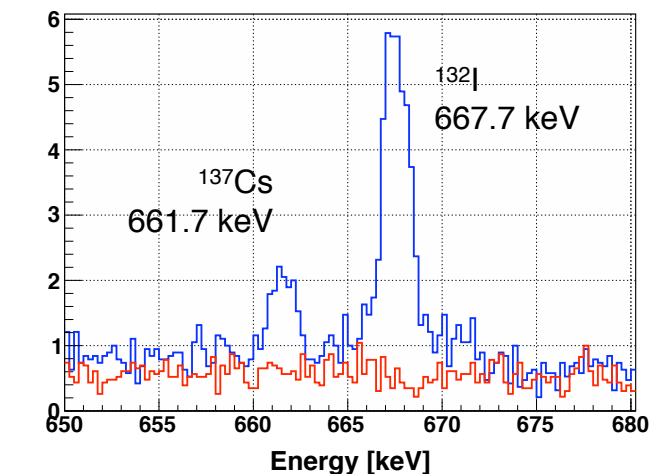
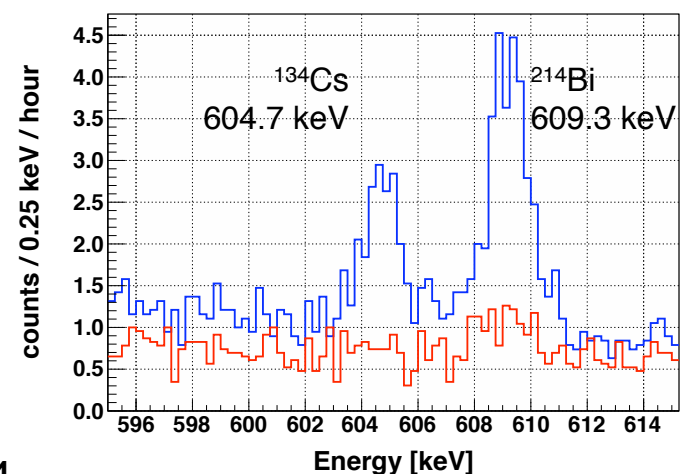
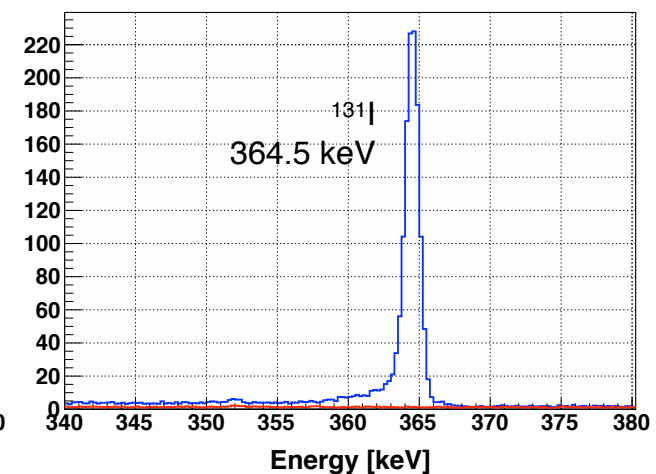
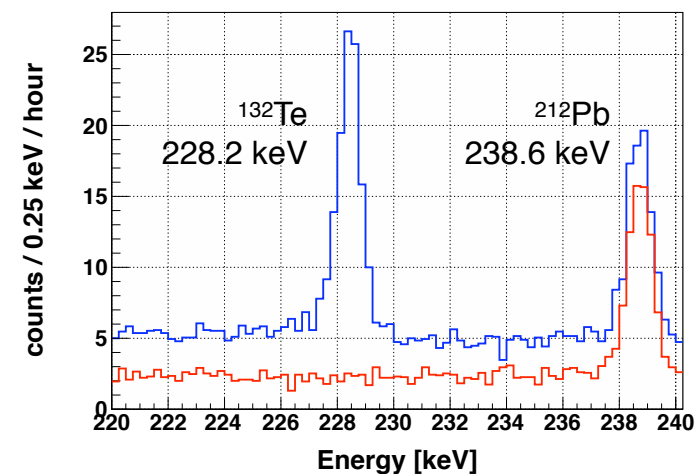
supplemental slides



# Airborne radioactivity in Seattle after the 2011 Fukushima earthquake



J. Diaz Leon et al., Journal of Environmental Radioactivity 102 (2011) 1032-1038  
arXiv:1103.4853



# energy calibration

51

Energy Linearity

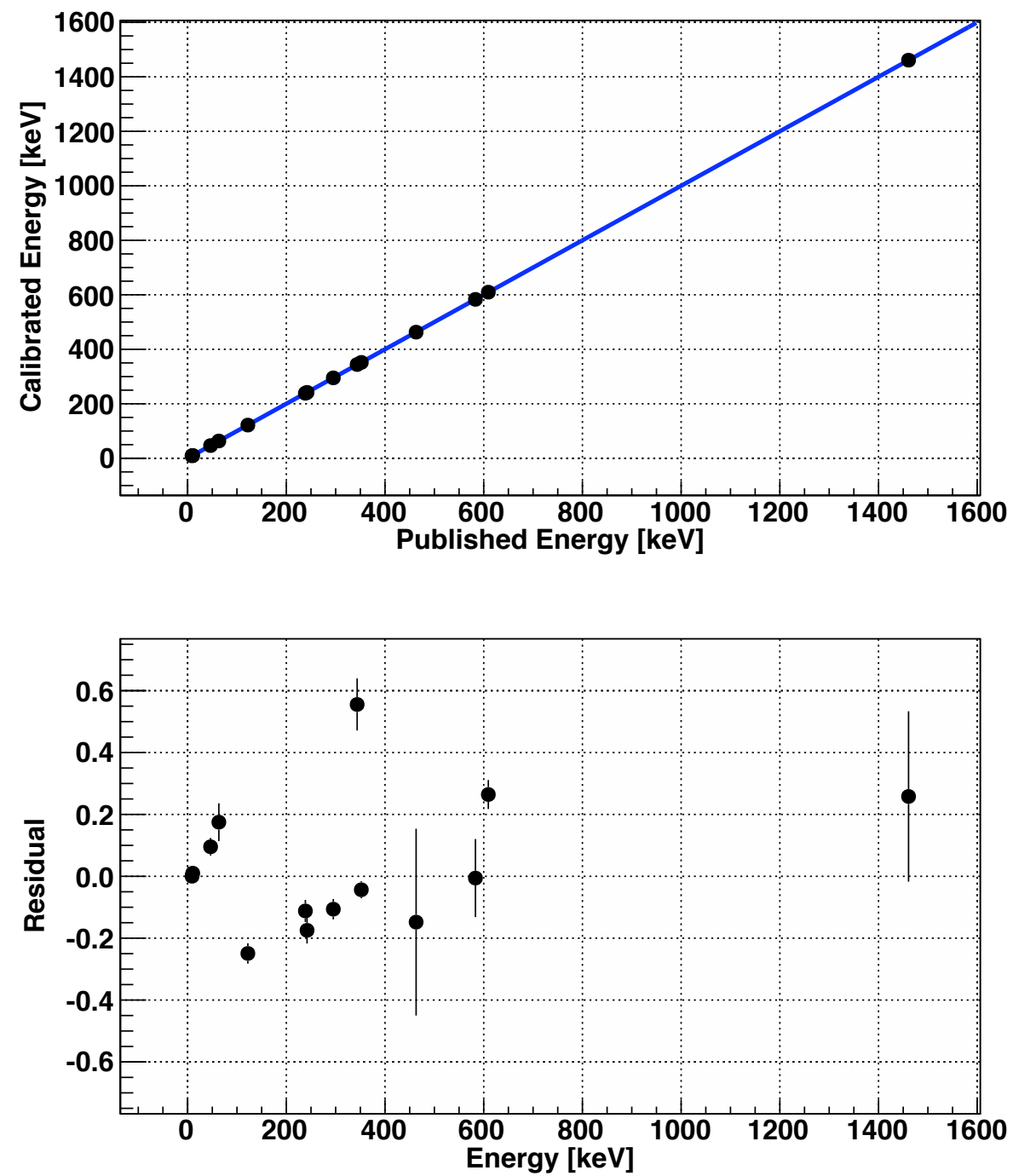


FIG. 47: Energy linearity.

$\chi^2/\text{DOF}$ : 194.21 / 14 ( 13.87), p-val: 8.33E-34

# identified peaks

Peak	Energy [keV]	Centroid [keV]	Sigma [eV]	Count Rate [μHz]	$\chi^2$ / DOF	P-value
$^{65}\text{Zn}^1$	8.98	$8.98 \pm 0.01$	$143.4 \pm 7.6$	$23.7 \pm 1.3$	39.3 / 53 ( 0.74 )	0.920
$^{68}\text{Ga}$	9.66	$9.66 \pm 0.03$	$141.7 \pm 33.3$	$3.8 \pm 0.8$		
$^{68}\text{Ge}^1$	10.37	$10.38 \pm 0.00$	$131.8 \pm 2.9$	$74.7 \pm 2.0$		
$^{210}\text{Pb}^1$	46.54	$46.63 \pm 0.03$	$162.1 \pm 25.0$	$4.1 \pm 0.6$	21.4 / 36 ( 0.59 )	0.975
$^{234}\text{U}, ^{72}\text{Ge}(n, \gamma)$	53.20, 53.53	$53.91 \pm 0.03$	$71.8 \pm 19.8$	$1.0 \pm 0.3$	19.5 / 36 ( 0.54 )	0.989
$^{234}\text{Th}^1$	63.29	$63.46 \pm 0.06$	$224.5 \pm 44.3$	$2.2 \pm 0.0$	19.3 / 36 ( 0.54 )	0.989
Bi $K_{\alpha 2}$	74.81	$75.06 \pm 0.06$	$223.4 \pm 54.0$	$2.3 \pm 0.6$	34.3 / 59 ( 0.58 )	0.996
Bi $K_{\alpha 1}$	77.11	$77.23 \pm 0.06$	$221.7 \pm 50.5$	$2.7 \pm 0.6$		
$^{234}\text{Th}$	92.38, 92.80	$92.76 \pm 0.05$	$440.9 \pm 41.4$	$9.0 \pm 0.0$	28.3 / 36 ( 0.79 )	0.816
$^{57}\text{Co}^1$	122.06	$121.81 \pm 0.03$	$280.4 \pm 27.8$	$8.3 \pm 0.0$	43.1 / 36 ( 1.20 )	0.193
$^{57}\text{Co}$ + atomic	143.58	$143.52 \pm 0.06$	$305.8 \pm 65.7$	$4.0 \pm 0.8$	26.1 / 36 ( 0.73 )	0.887
$\gamma^1$	-	$185.60 \pm 0.04$	$334.7 \pm 30.6$	$8.9 \pm 0.0$	19.0 / 36 ( 0.53 )	0.991
$^{228}\text{Ac}$	209.25	$209.56 \pm 0.16$	$226.5 \pm 97.8$	$0.7 \pm 0.0$	20.5 / 36 ( 0.57 )	0.982
$^{212}\text{Pb}^1$	238.63	$238.52 \pm 0.03$	$304.2 \pm 33.2$	$8.0 \pm 0.8$	33.9 / 54 ( 0.63 )	0.985
$^{214}\text{Pb}^1$	242.00	$241.82 \pm 0.04$	$338.1 \pm 41.1$	$7.3 \pm 0.8$		
$^{214}\text{Pb}^1$	295.22	$295.12 \pm 0.03$	$408.5 \pm 29.7$	$13.8 \pm 1.0$	14.0 / 36 ( 0.39 )	1.000
$^{228}\text{Ac}$	338.32	$338.13 \pm 0.10$	$265.5 \pm 88.3$	$1.3 \pm 0.4$	28.1 / 36 ( 0.78 )	0.822
$^{206}\text{Pb}(n, n'\gamma)^1$	343.51	$344.07 \pm 0.08$	$387.7 \pm 72.4$	$2.9 \pm 0.6$	26.3 / 36 ( 0.73 )	0.881
$^{214}\text{Pb}^1$	351.93	$351.89 \pm 0.03$	$442.9 \pm 21.5$	$20.8 \pm 0.0$	25.4 / 36 ( 0.71 )	0.905
$^{228}\text{Ac}^1$	463.00	$462.86 \pm 0.30$	$1022.9 \pm 217.2$	$2.1 \pm 0.0$	29.5 / 36 ( 0.82 )	0.768
$^{208}\text{Tl}$ + annih.	510.77, 511.00	$510.88 \pm 0.20$	$1217.0 \pm 190.1$	$5.2 \pm 0.0$	23.5 / 36 ( 0.65 )	0.946
$^{208}\text{Tl}^1$	583.19	$583.18 \pm 0.13$	$620.8 \pm 100.0$	$2.7 \pm 0.5$	17.2 / 36 ( 0.48 )	0.997
$^{214}\text{Bi}^1$	609.32	$609.58 \pm 0.05$	$703.7 \pm 40.2$	$15.6 \pm 1.0$	21.2 / 36 ( 0.59 )	0.976
$^{214}\text{Bi}$	768.36	$768.36 \pm 0.24$	$1106.2 \pm 216.0$	$2.6 \pm 0.5$	15.1 / 36 ( 0.42 )	0.999
$^{206}\text{Pb}(n, n'\gamma)$	803.10	$803.61 \pm 0.35$	$1042.7 \pm 280.8$	$1.3 \pm 0.4$	14.4 / 36 ( 0.40 )	0.999
$^{228}\text{Ac}$	911.20	$911.80 \pm 0.22$	$1025.3 \pm 183.3$	$2.3 \pm 0.5$	13.9 / 36 ( 0.39 )	1.000
?	-	$1086.22 \pm 0.24$	$608.4 \pm 161.2$	$0.7 \pm 0.0$	13.4 / 36 ( 0.37 )	1.000
$^{214}\text{Bi}$	1120.29	$1122.00 \pm 0.52$	$2129.8 \pm 477.1$	$2.5 \pm 0.6$	23.6 / 36 ( 0.65 )	0.945
$^{40}\text{K}^1$	1460.82	$1461.08 \pm 0.28$	$1887.3 \pm 244.6$	$3.3 \pm 0.5$	12.7 / 36 ( 0.35 )	1.000
$^{214}\text{Bi}$	1764.49	$1765.05 \pm 0.41$	$1042.3 \pm 354.4$	$0.8 \pm 0.2$	11.1 / 36 ( 0.31 )	1.000
$^{214}\text{Bi}$	2204.06	$2257.12 \pm 0.00$	$14567.5 \pm 0.0$	$1.3 \pm 0.0$	2.4 / 36 ( 0.07 )	1.000



# sigma vs. energy

55

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Sigma vs. Energy

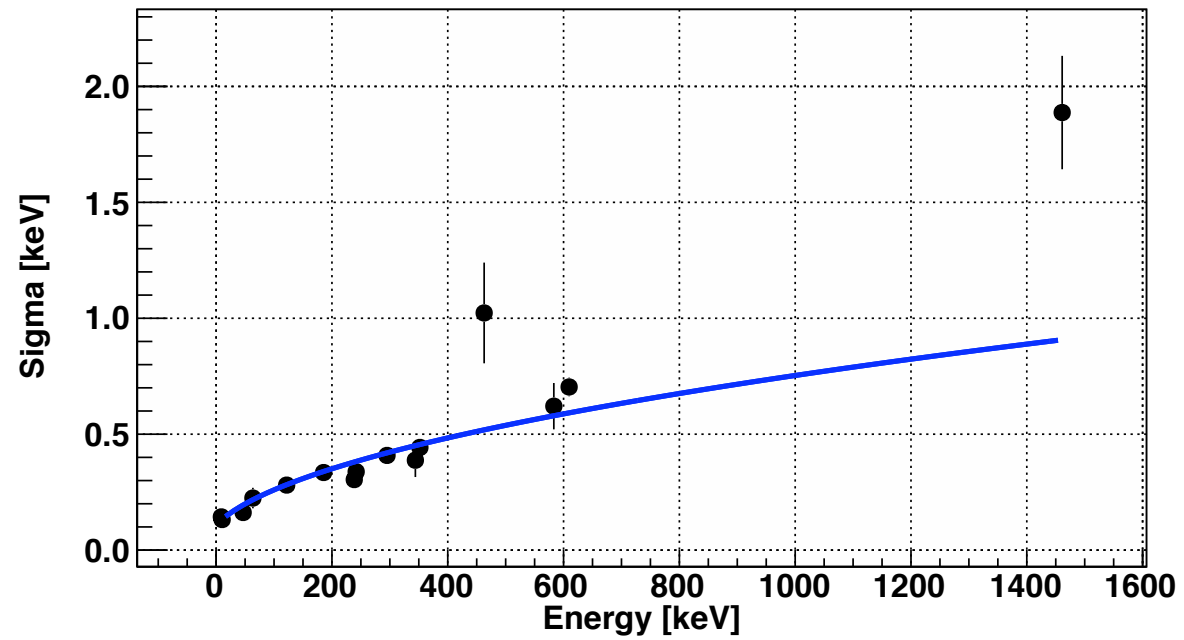


FIG. 51:

Sigma vs. Energy (with linear term)

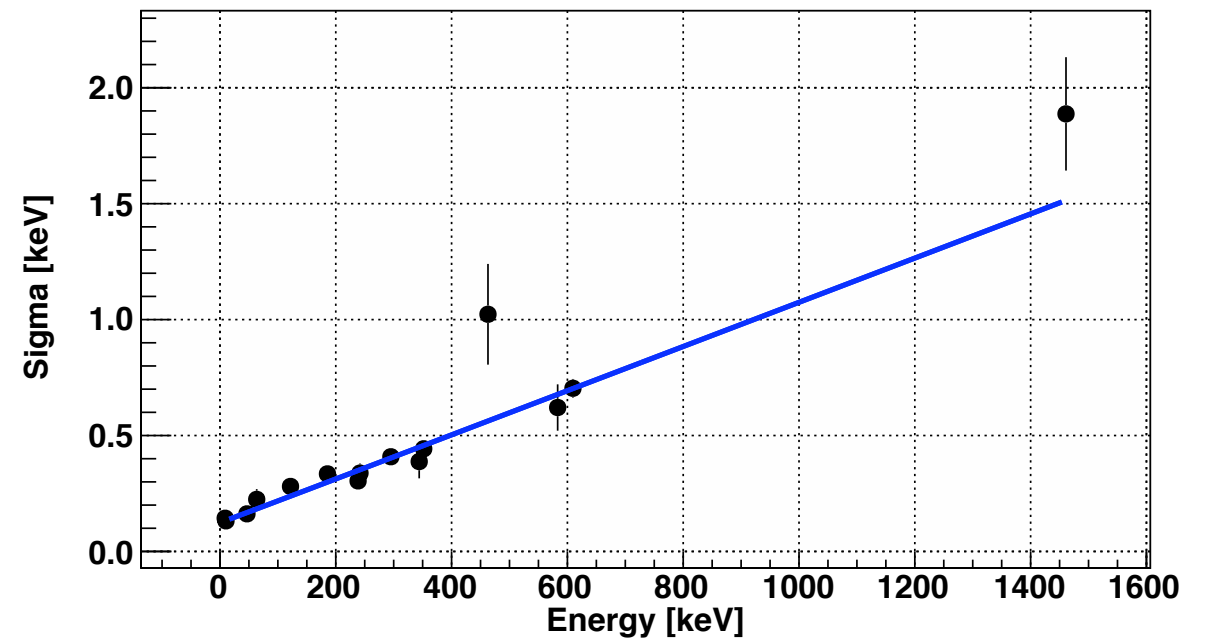
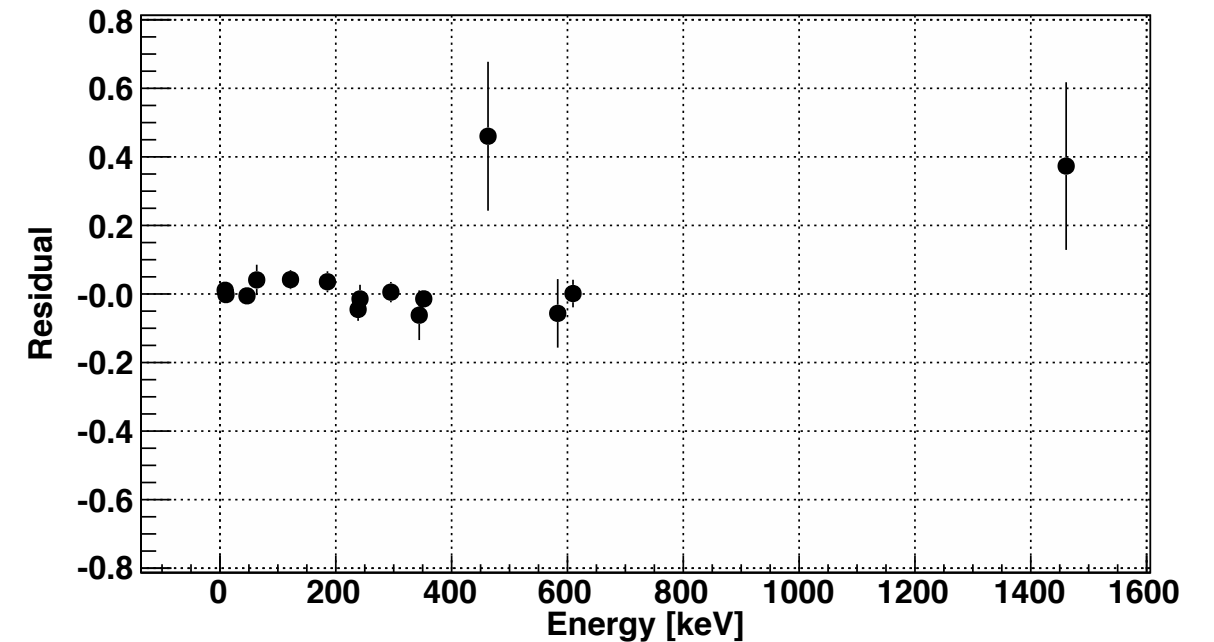
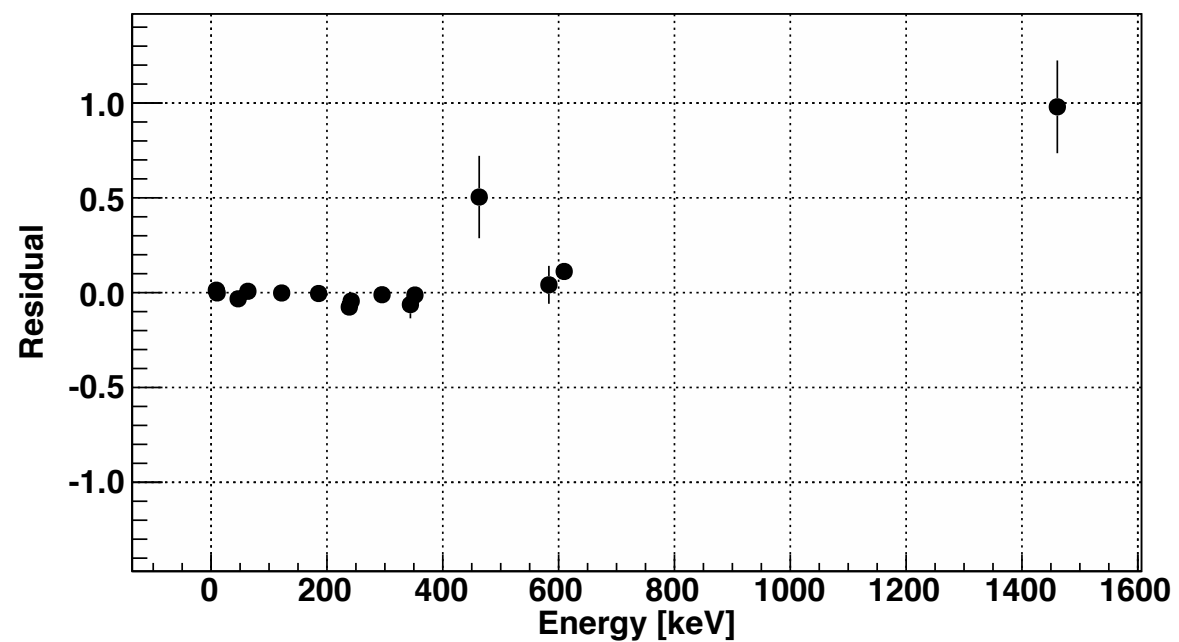


FIG. 52:



# radiopurity info

Table J.4: Radiopurity information for Copper-OFHC. Table generated by *MJBMDbInfo.-ComponentsStore*.

Contaminant	Activity/ Production Rate	Reference
$^{232}\text{Th}$ to $^{228}\text{Ra}$ ( $^{232}\text{Th}$ step 1)	9.00E-01 $\mu\text{Bq/kg}$	DEMONSTRATOR Table [78]
$^{228}\text{Ra}$ to $^{228}\text{Th}$ ( $^{232}\text{Th}$ step 2)	9.00E-01 $\mu\text{Bq/kg}$	DEMONSTRATOR Table [78]
$^{228}\text{Th}$ to $^{224}\text{Ra}$ ( $^{232}\text{Th}$ step 3)	9.00E-01 $\mu\text{Bq/kg}$	DEMONSTRATOR Table [78]
$^{224}\text{Ra}$ to $^{208}\text{Pb}$ ( $^{232}\text{Th}$ step 4)	9.00E-01 $\mu\text{Bq/kg}$	DEMONSTRATOR Table [78]
$^{238}\text{U}$ to $^{234}\text{Th}$ ( $^{238}\text{U}$ step 1)	3.00E+00 $\mu\text{Bq/kg}$	DEMONSTRATOR Table [78]
$^{234}\text{Th}$ to $^{234}\text{U}$ ( $^{238}\text{U}$ step 2)	3.00E+00 $\mu\text{Bq/kg}$	DEMONSTRATOR Table [78]
$^{234}\text{U}$ to $^{230}\text{Th}$ ( $^{238}\text{U}$ step 3)	3.00E+00 $\mu\text{Bq/kg}$	DEMONSTRATOR Table [78]
$^{230}\text{Th}$ to $^{226}\text{Ra}$ ( $^{238}\text{U}$ step 4)	3.00E+00 $\mu\text{Bq/kg}$	DEMONSTRATOR Table [78]
$^{226}\text{Ra}$ to $^{222}\text{Rn}$ ( $^{238}\text{U}$ step 5)	3.00E+00 $\mu\text{Bq/kg}$	DEMONSTRATOR Table [78]
$^{222}\text{Rn}$ to $^{210}\text{Tl}$ or $^{210}\text{Pb}$ ( $^{238}\text{U}$ step 6)	3.00E+00 $\mu\text{Bq/kg}$	DEMONSTRATOR Table [78]
$^{210}\text{Tl}$ to $^{210}\text{Pb}$ ( $^{238}\text{U}$ step 6a)	6.30E-04 $\mu\text{Bq/kg}$	DEMONSTRATOR Table [78]
$^{210}\text{Pb}$ to $^{210}\text{Bi}$ or $^{206}\text{Pb}$ ( $^{238}\text{U}$ step 7)	3.00E+00 $\mu\text{Bq/kg}$	DEMONSTRATOR Table [78]
$^{210}\text{Bi}$ to $^{210}\text{Po}$ or $^{206}\text{Pb}$ ( $^{238}\text{U}$ step 8)	3.00E+00 $\mu\text{Bq/kg}$	DEMONSTRATOR Table [78]
$^{210}\text{Po}$ to $^{206}\text{Pb}$ ( $^{238}\text{U}$ step 9)	3.00E+00 $\mu\text{Bq/kg}$	DEMONSTRATOR Table [78]
$^{40}\text{K}$	1.24E+01 $\mu\text{Bq/kg}$	EXO [75]
$^{46}\text{Sc}$	4.58E+00 atoms/kg/day	Heusser et al. [79]
$^{48}\text{V}$	9.50E+00 atoms/kg/day	Heusser et al. [79]
$^{56}\text{Co}$	1.99E+01 atoms/kg/day	Heusser et al. [79]
$^{57}\text{Co}$	1.56E+02 atoms/kg/day	Heusser et al. [79]
$^{58}\text{Co}$	1.43E+02 atoms/kg/day	Heusser et al. [79]
$^{59}\text{Fe}$	3.93E+01 atoms/kg/day	Heusser et al. [79]
$^{60}\text{Co}$	2.00E+02 atoms/kg/day	DEMONSTRATOR Table [78]

Table J.5: Radiopurity information for Germanium-Nat. Table generated by *MJBMDbInfo.-ComponentsStore*.

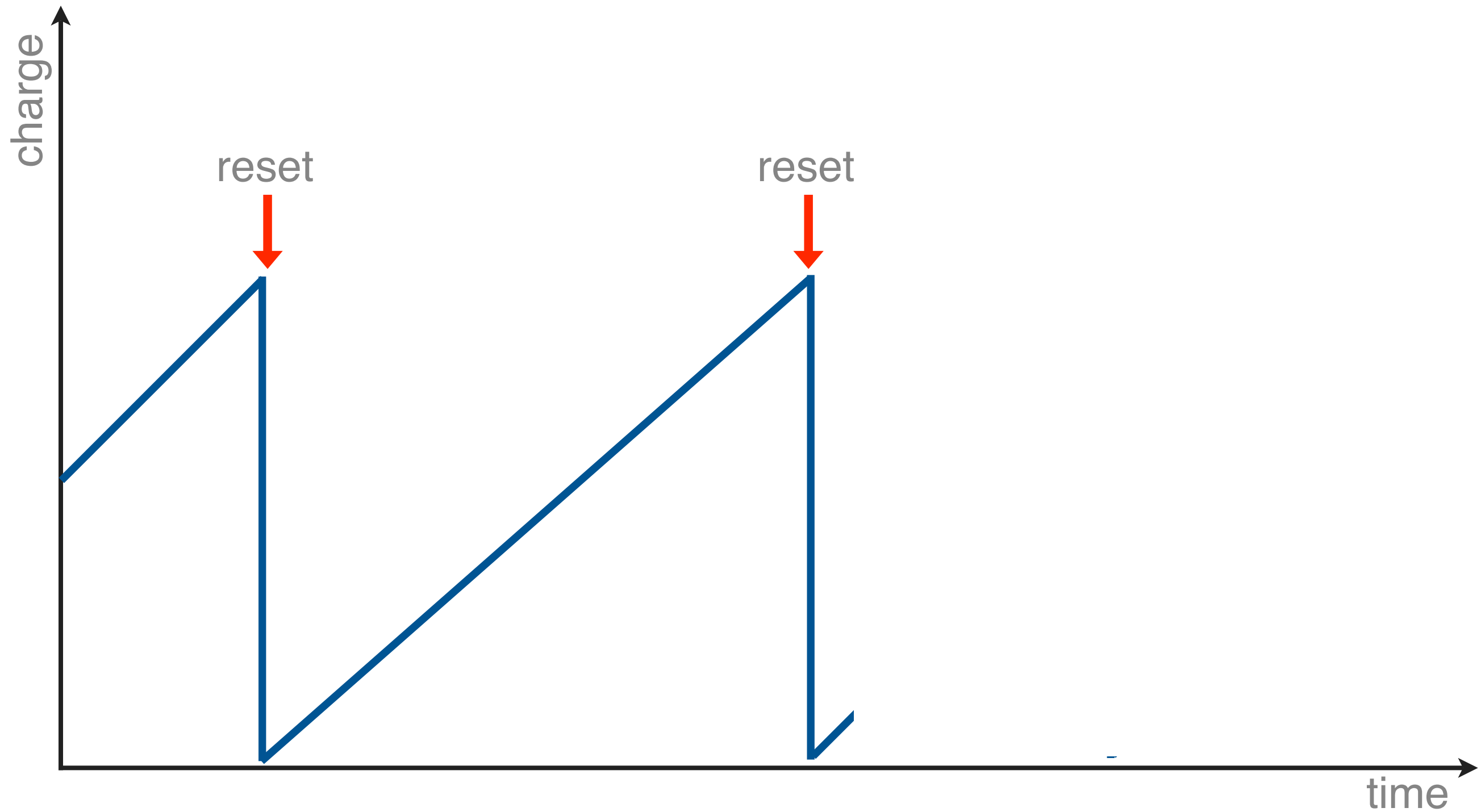
Contaminant	Activity/ Production Rate	Reference
$^{232}\text{Th}$ to $^{228}\text{Ra}$ ( $^{232}\text{Th}$ step 1)	1.42E-02 $\mu\text{Bq/kg}$	DEMONSTRATOR Table [78]
$^{228}\text{Ra}$ to $^{228}\text{Th}$ ( $^{232}\text{Th}$ step 2)	1.42E-02 $\mu\text{Bq/kg}$	DEMONSTRATOR Table [78]
$^{228}\text{Th}$ to $^{224}\text{Ra}$ ( $^{232}\text{Th}$ step 3)	1.42E-02 $\mu\text{Bq/kg}$	DEMONSTRATOR Table [78]
$^{224}\text{Ra}$ to $^{208}\text{Pb}$ ( $^{232}\text{Th}$ step 4)	1.42E-02 $\mu\text{Bq/kg}$	DEMONSTRATOR Table [78]
$^{238}\text{U}$ to $^{234}\text{Th}$ ( $^{238}\text{U}$ step 1)	1.38E-02 $\mu\text{Bq/kg}$	DEMONSTRATOR Table [78]
$^{234}\text{Th}$ to $^{234}\text{U}$ ( $^{238}\text{U}$ step 2)	1.38E-02 $\mu\text{Bq/kg}$	DEMONSTRATOR Table [78]
$^{234}\text{U}$ to $^{230}\text{Th}$ ( $^{238}\text{U}$ step 3)	1.38E-02 $\mu\text{Bq/kg}$	DEMONSTRATOR Table [78]
$^{230}\text{Th}$ to $^{226}\text{Ra}$ ( $^{238}\text{U}$ step 4)	1.38E-02 $\mu\text{Bq/kg}$	DEMONSTRATOR Table [78]
$^{226}\text{Ra}$ to $^{222}\text{Rn}$ ( $^{238}\text{U}$ step 5)	1.38E-02 $\mu\text{Bq/kg}$	DEMONSTRATOR Table [78]
$^{222}\text{Rn}$ to $^{210}\text{Tl}$ or $^{210}\text{Pb}$ ( $^{238}\text{U}$ step 6)	1.38E-02 $\mu\text{Bq/kg}$	DEMONSTRATOR Table [78]
$^{210}\text{Tl}$ to $^{210}\text{Pb}$ ( $^{238}\text{U}$ step 6a)	2.89E-06 $\mu\text{Bq/kg}$	DEMONSTRATOR Table [78]
$^{210}\text{Pb}$ to $^{210}\text{Bi}$ or $^{206}\text{Pb}$ ( $^{238}\text{U}$ step 7)	1.38E-02 $\mu\text{Bq/kg}$	DEMONSTRATOR Table [78]
$^{210}\text{Bi}$ to $^{210}\text{Po}$ or $^{206}\text{Pb}$ ( $^{238}\text{U}$ step 8)	1.38E-02 $\mu\text{Bq/kg}$	DEMONSTRATOR Table [78]
$^{210}\text{Po}$ to $^{206}\text{Pb}$ ( $^{238}\text{U}$ step 9)	1.38E-02 $\mu\text{Bq/kg}$	DEMONSTRATOR Table [78]
$^3\text{H}$	2.77E+01 atoms/kg/day	D.-M. Mei [80]
$^{54}\text{Mn}$	9.10E+00 atoms/kg/day	Avg. from Table I [81]
$^{55}\text{Fe}$	8.40E+00 atoms/kg/day	MAJORANA BSD – GENIUS [61]
$^{57}\text{Co}$	6.84E+00 atoms/kg/day	Avg. from Table I [81]
$^{58}\text{Co}$	1.61E+01 atoms/kg/day	MAJORANA BSD – GENIUS [61]
$^{60}\text{Co}$	5.00E+00 atoms/kg/day	DEMONSTRATOR Table [78]
$^{63}\text{Ni}$	4.60E+00 atoms/kg/day	MAJORANA BSD – GENIUS [61]
$^{65}\text{Zn}$	7.90E+01 atoms/kg/day	MAJORANA BSD – GENIUS [61]
$^{68}\text{Ge}$	3.00E+01 atoms/kg/day	DEMONSTRATOR Table [78]
$^{76}\text{Ge } 2\nu\beta\beta$	9.03E+00 $\mu\text{Bq/kg}$	A.S. Barabash [1]

# Pulsed-reset preamplifier

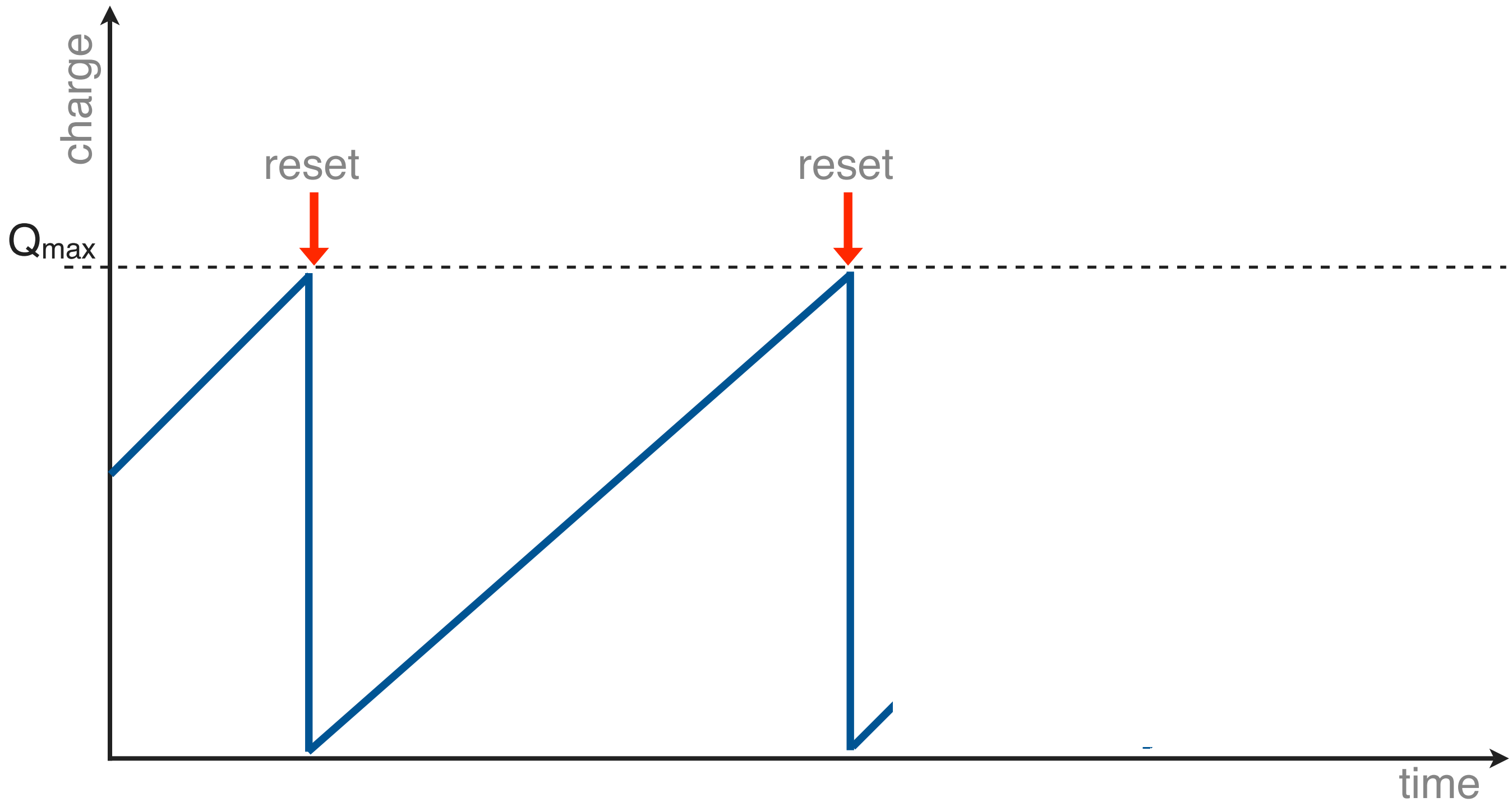




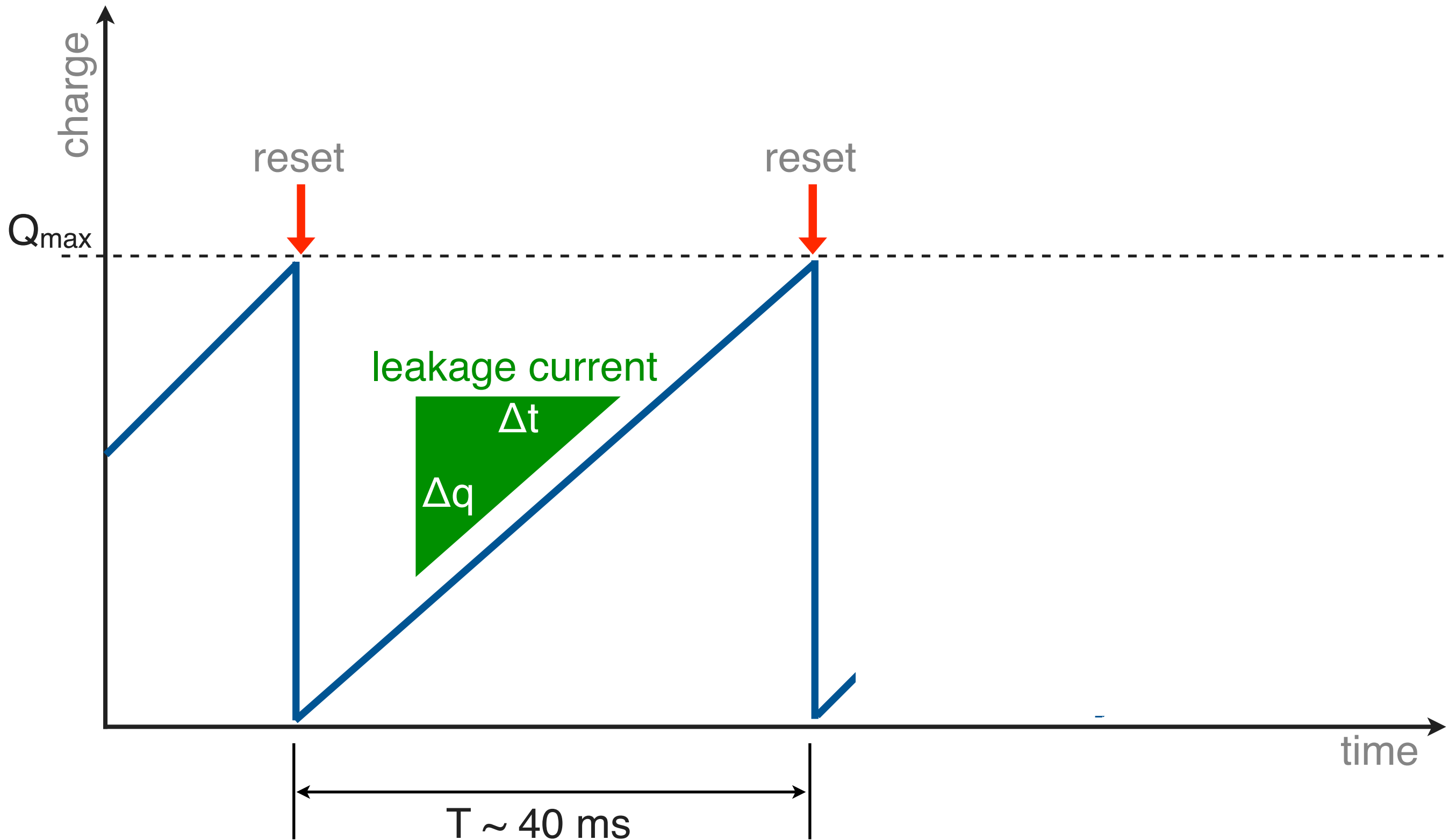
# Pulsed-reset preamplifier



# Pulsed-reset preamplifier

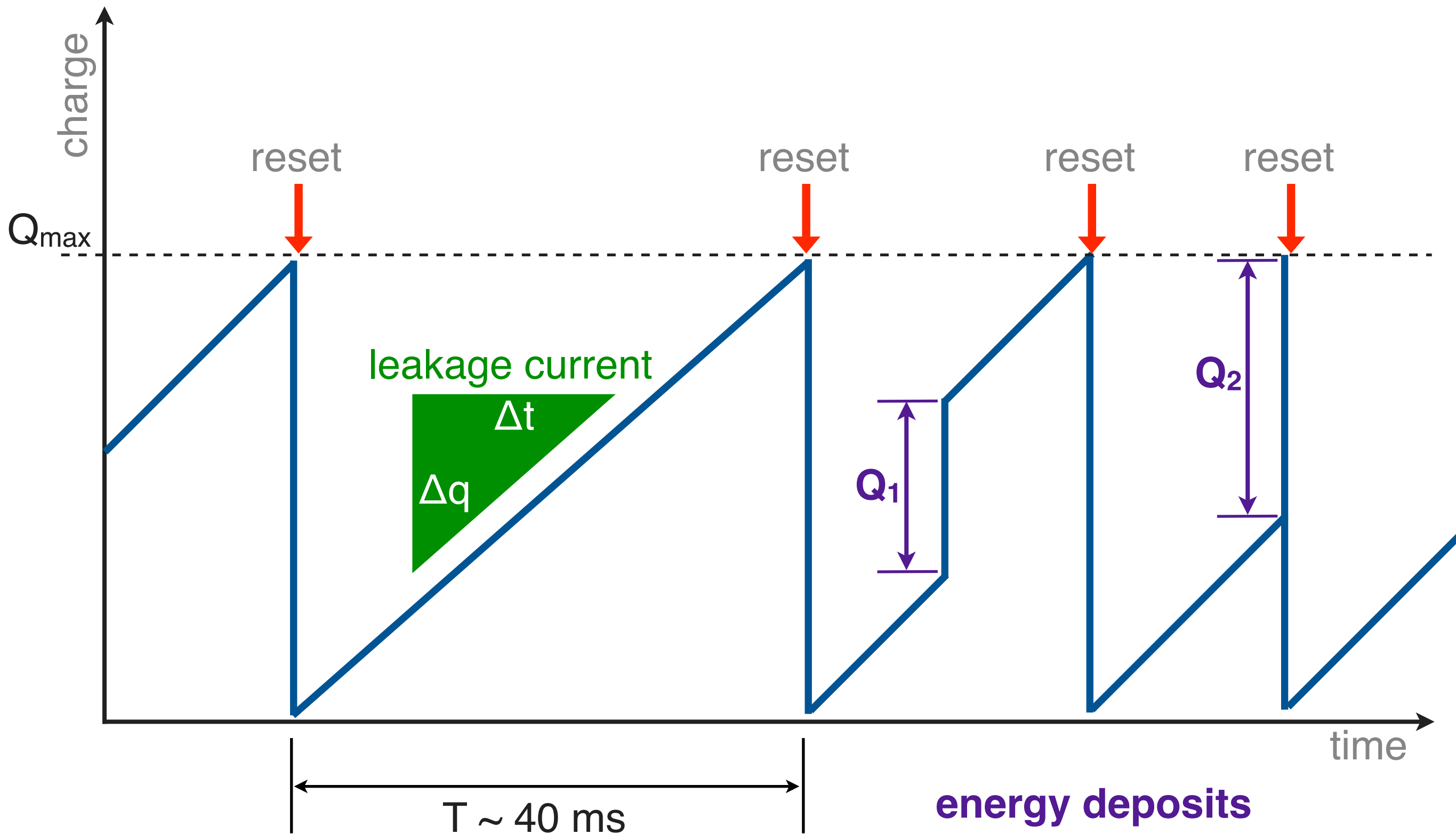


# Pulsed-reset preamplifier

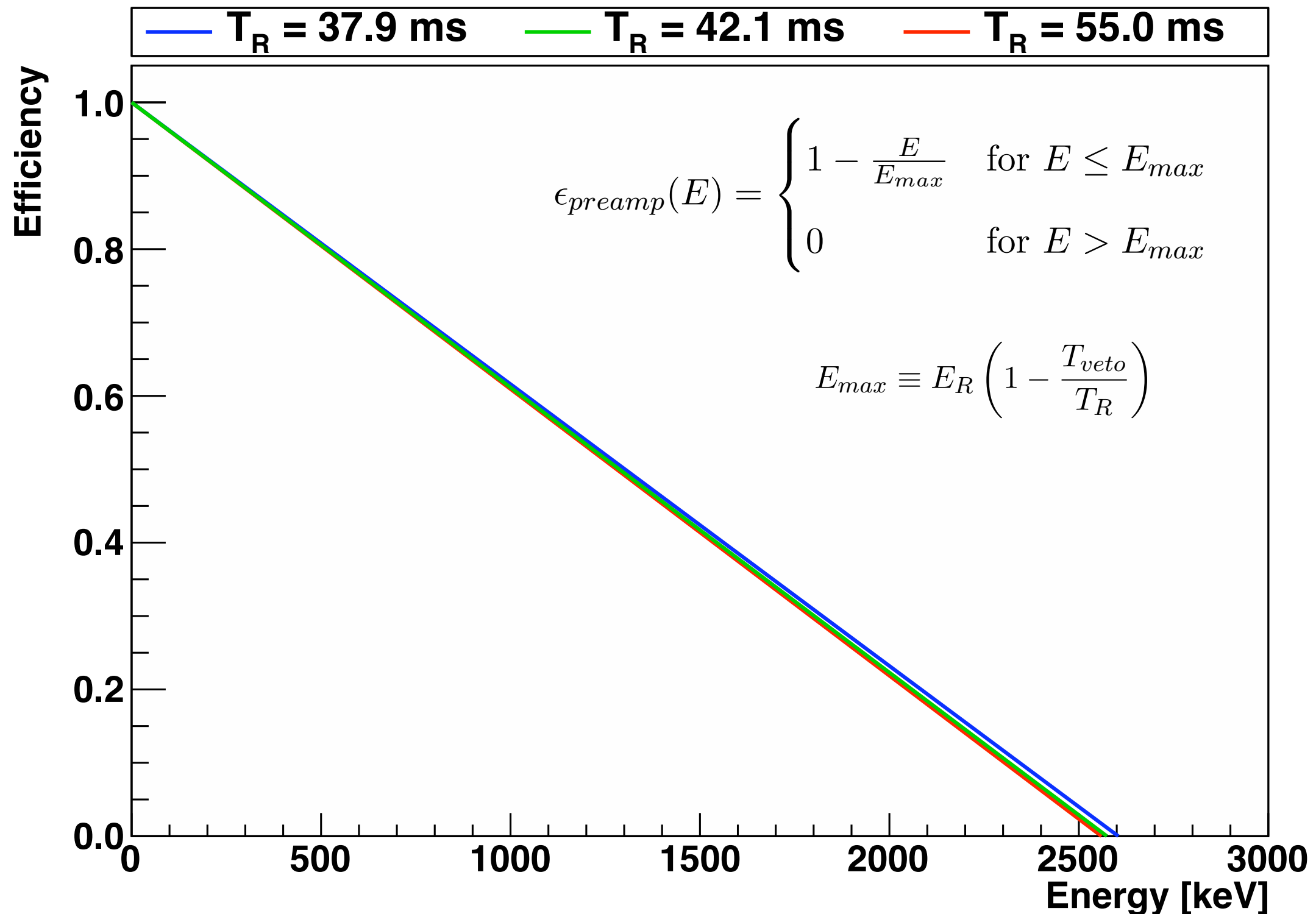




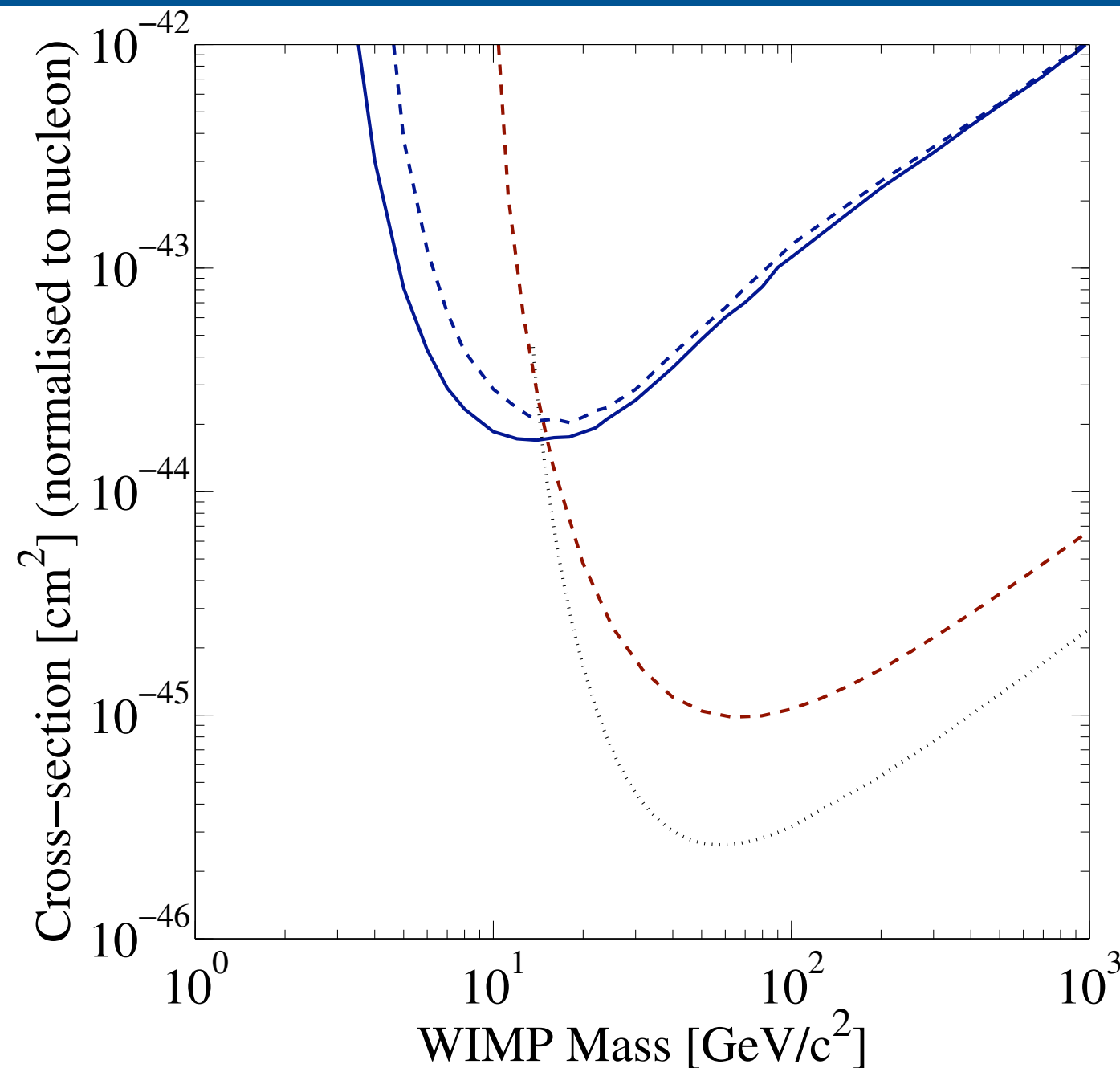
# Pulsed-reset preamplifier



# Energy-dependent efficiency



# DEMONSTRATOR WIMP sensitivity

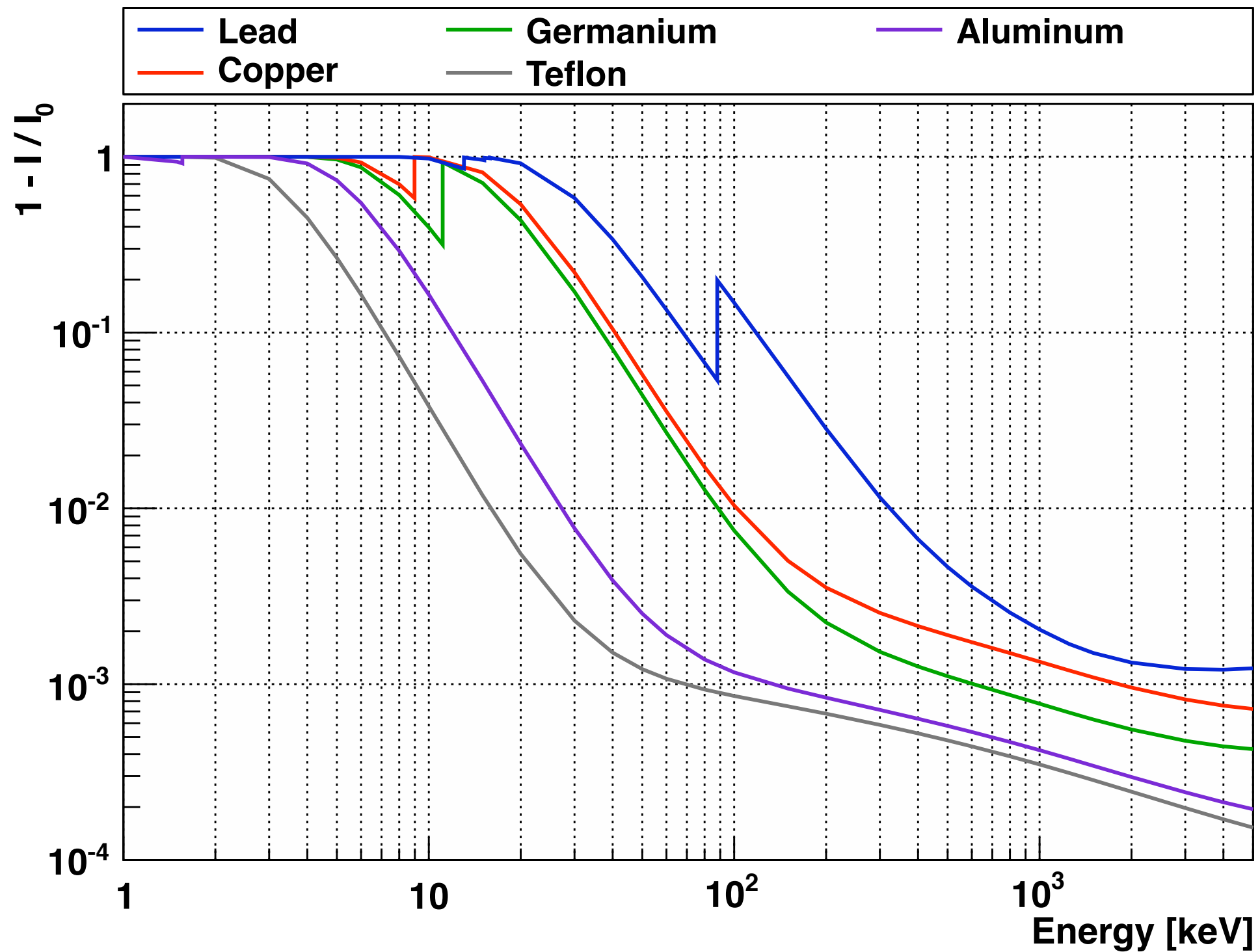


Mike Marino  
PhD. Thesis, 2010

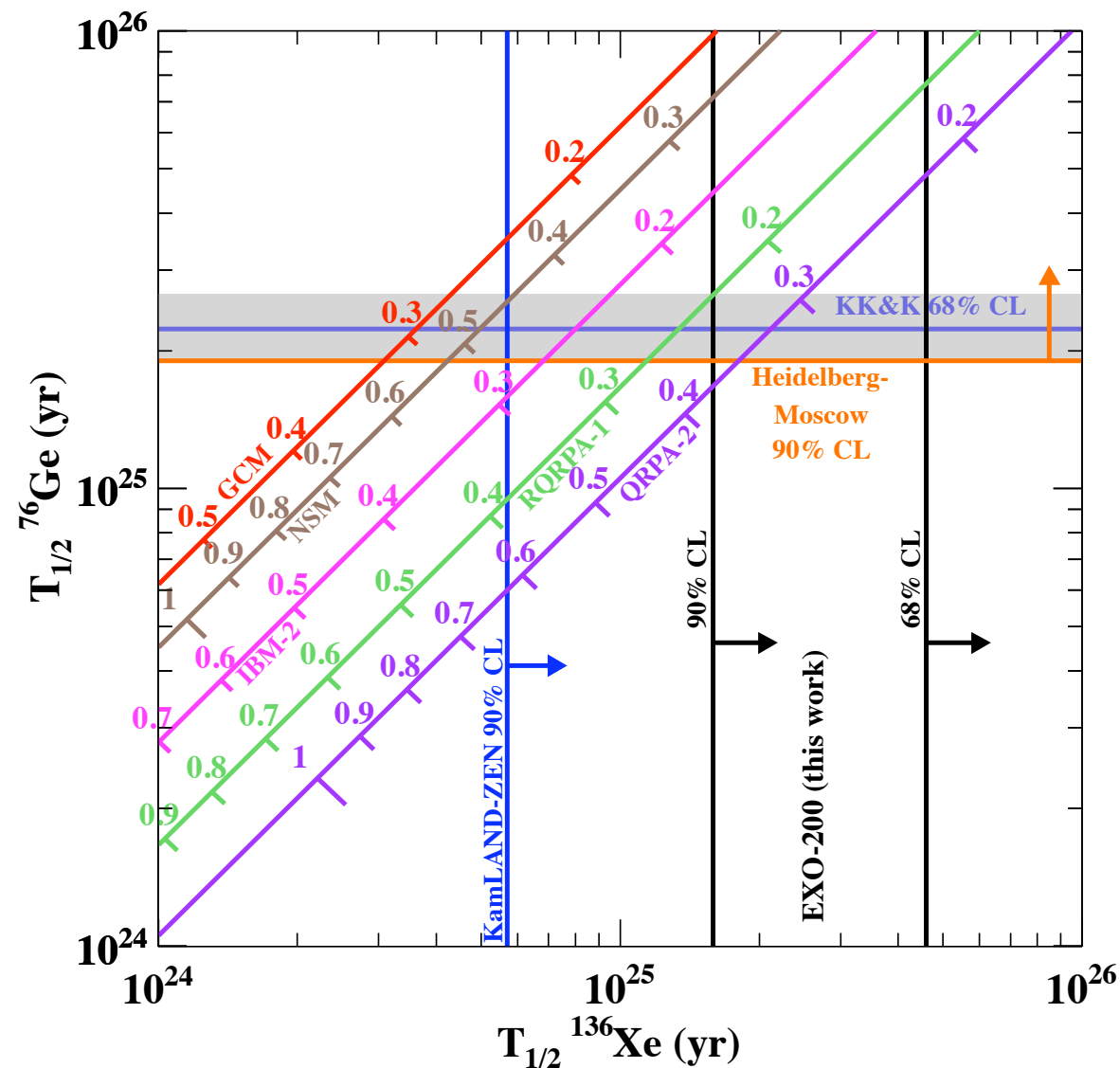
Figure 5.7: MAJORANA DEMONSTRATOR sensitivity to a WIMP signal (blue solid, 0.3 keV threshold; blue dashed, 0.5 keV threshold), comparing to SuperCDMS Phase A [96] (red dashed) and LUX 300 [97] (black dotted). Plot generated with DMTools [86], lines are 90% CLexclusions.



# Photon attenuation



# EXO $0\nu\beta\beta$ limit

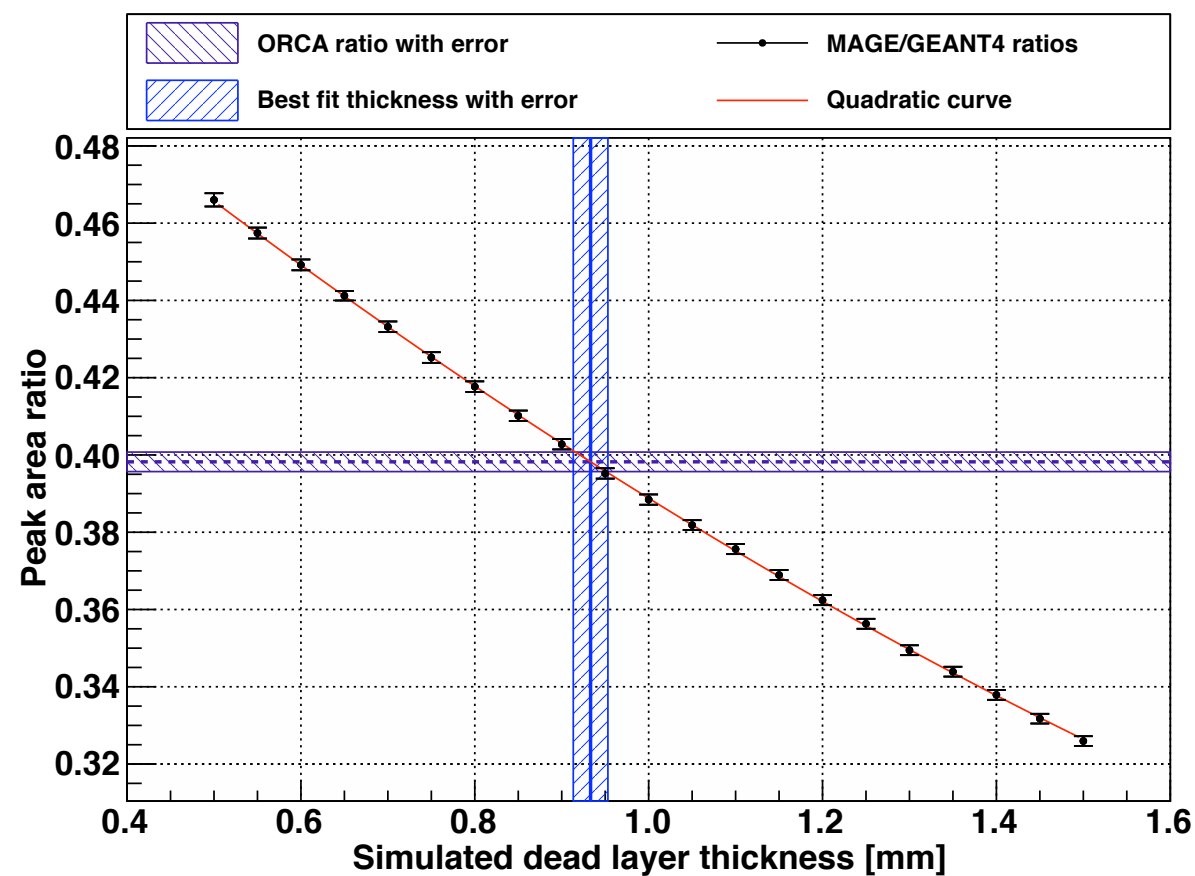


EXO-200: Phys. Rev. Lett. 109 (2012) 032505

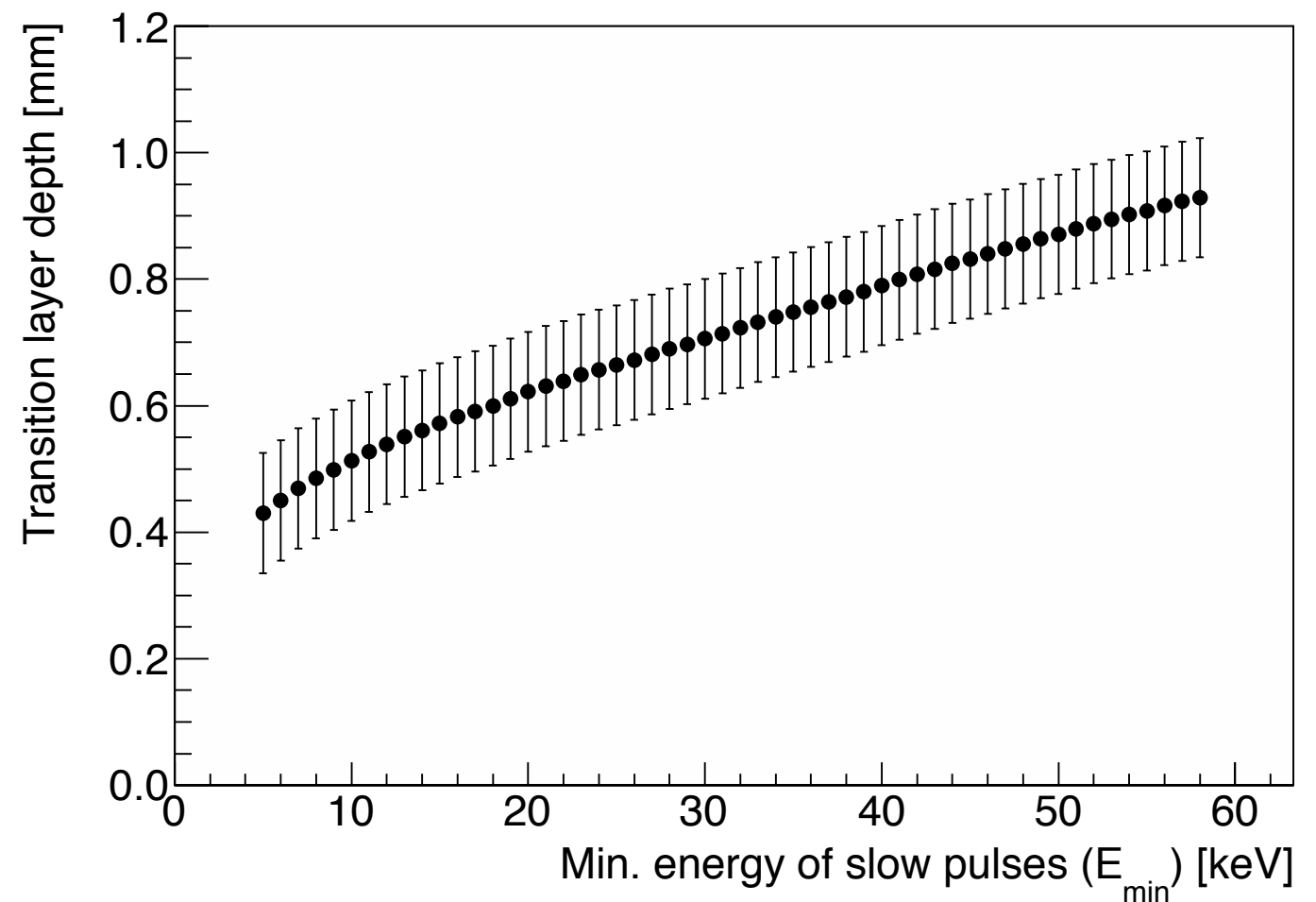
IG. 6: Relation between the  $T_{1/2}^{0\nu\beta\beta}$  in  $^{76}\text{Ge}$  and  $^{136}\text{Xe}$  for different matrix element calculations (GCM [20], NSM [21], 3M-2 [22], RQRPA-1 [23] and QRPA-2 [5]). For each matrix element  $\langle m \rangle_{\beta\beta}$  is also shown (eV). The claim [4] is represented by the grey band, along with the best limit for  $^{76}\text{Ge}$  [19]. The result reported here is shown along with that from [7].

# Slow energy-degraded pulses from the MALBEK transition dead layer

## Total dead layer measurement

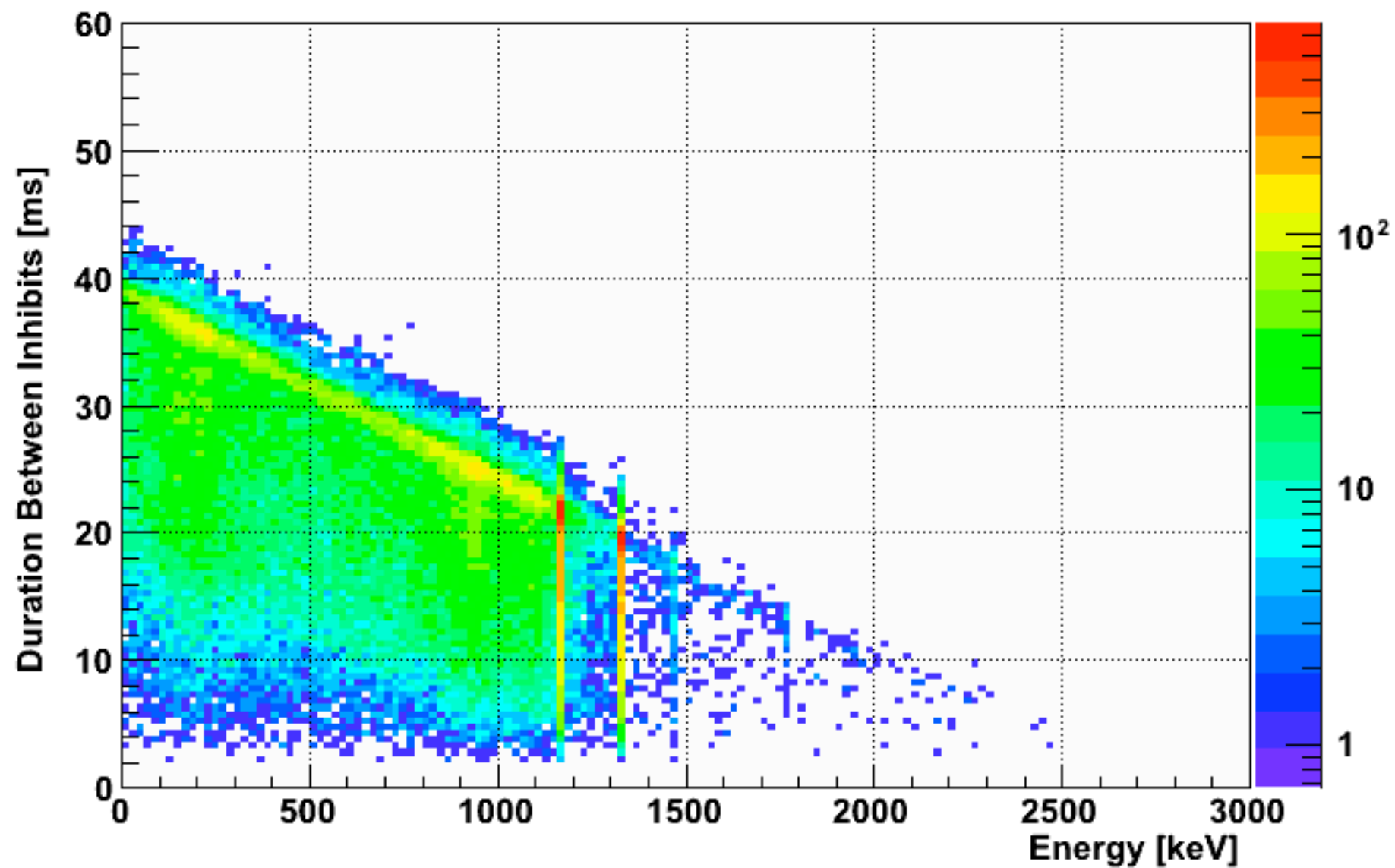


## Fractional charge collection vs. depth

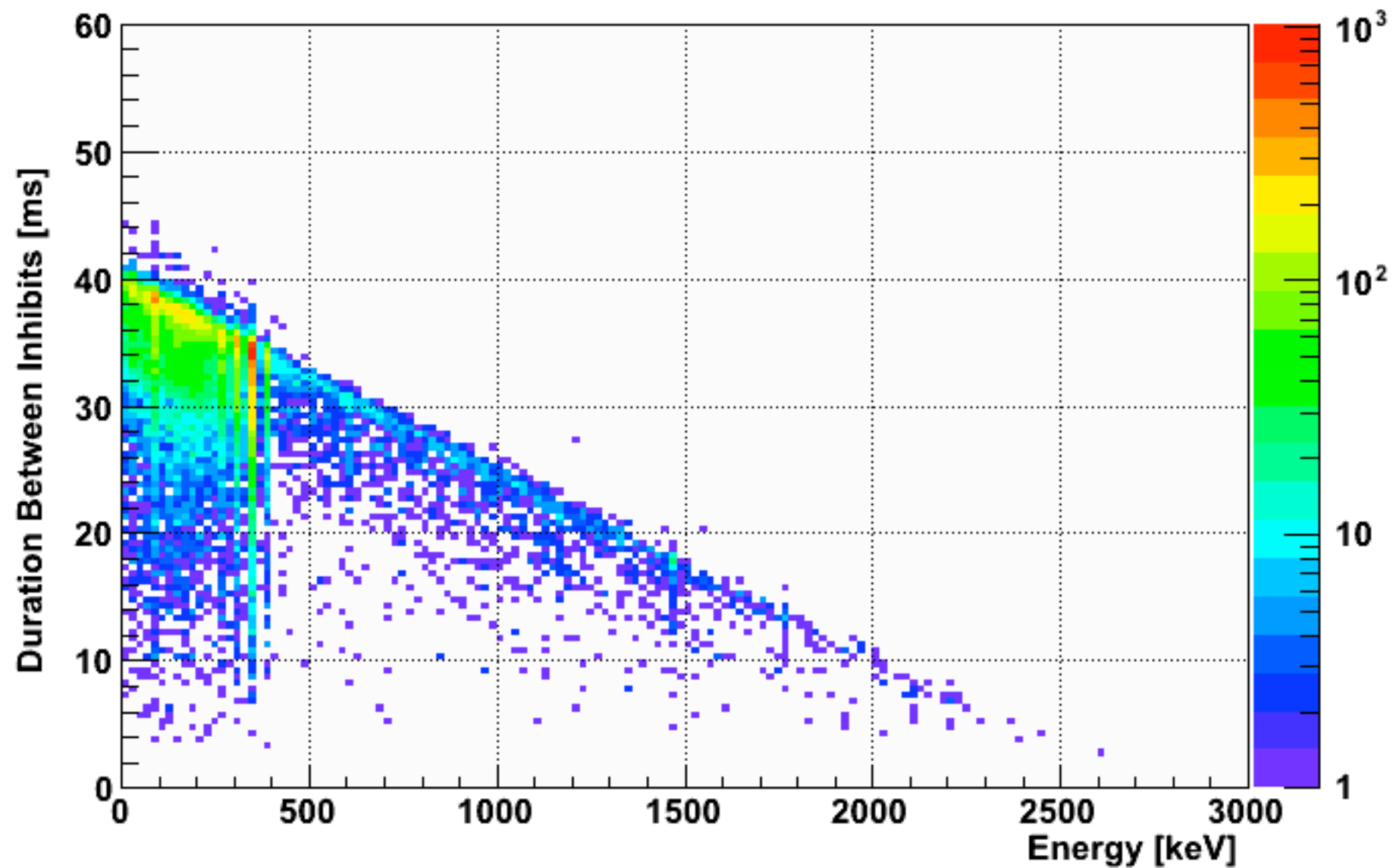




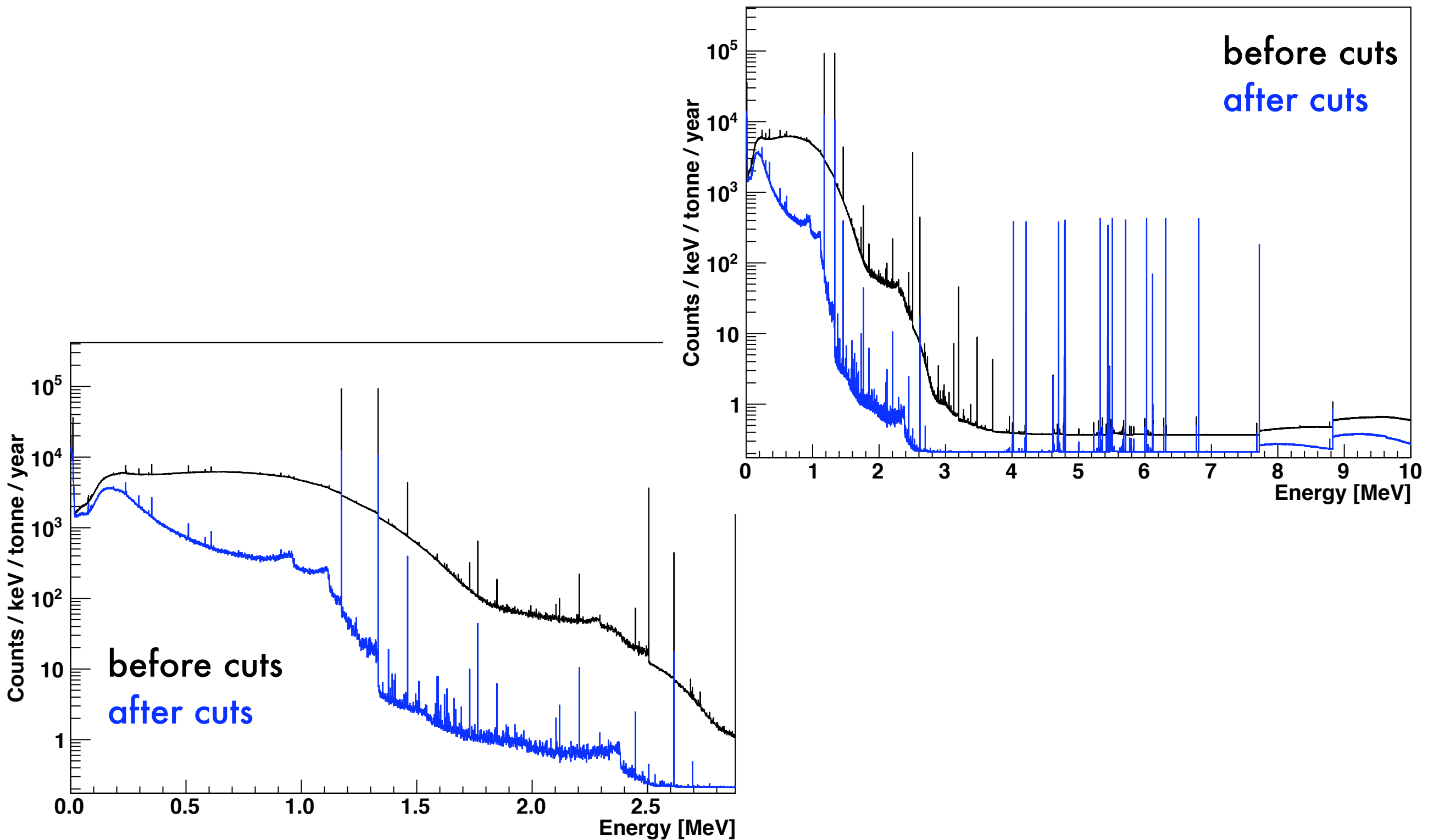
# Pulsed-reset preamp



# Pulsed-reset preamp



# DEMONSTRATOR background model





# questions

- GERDA: P1=18kg enr-Ge, P2 = 20kg enr-Ge?
- Cu purity: limits or measurements?
- PPC technology
- enr. Ge
- add DM/PPC slide
- add assay achievements/
- put status bullets on pics

Make sure slide # is visible on every slide!